

EFFECTS OF ASTEROIDS ON THE ORBITAL MOTIONS OF TERRESTRIAL PLANETS

S. ALJBAAE, J. SOUCHAY

Observatoire de Paris - SYRTE

e-mail: safwan.aljbaae@obspm.fr

ABSTRACT. The perturbations from the largest ~ 300 asteroids which were taken into account in the ephemerides DE403 (Standish et al. 1995), EPM98 (Pitjeva 2001), and INPOP08 (Fienga et al. 2009) are a major problem in the construction of theses ephemeris. Therefore, it looks important to evaluate of the individual effects of largest asteroids of the solar system on the orbits of the terrestrial planets (Mercury, Venus, Earth and Mars) because these effects could reach a few kilometers in several decades in the case of Mars. This is the purpose of this work. For that our methodology consists in several stages:

- ★ A numerical integration of the orbits of the planets at short and long time scales with and without the disturbing asteroid from which we want to know the effects.
- ★ A determination of the signal representing the effects, by simple subtraction.
- ★ The analysis of the signal by the method of FFT (Fast Fourier Transform);
- ★ The adjustment of the signal by a set of sinusoids determined in the previous step.

We analyze in detail the influences of 43 among the largest asteroids on the six orbital elements a , e , i , Ω , $\tilde{\omega}$ and λ of Mercury, Venus, the EMB (Earth-Moon barycenter) and Mars. In addition we study their influence on two fundamental parameters: the distance and the orientation vector from the EMB to each of the terrestrial planets. This type of study is interesting in many fields, such as planetary ephemerides, as well as space navigation, to understand better the effects of each asteroid taken individually on the terrestrial planets. Note that this type of study is a continuation of previous ones (Williams, 1984; Mouret et al. 2009).

1. INTRODUCTION

The motion of a given planet around the Sun can be considered at first approximation as a Keplerian motion perturbed by the other planets and the small bodies of the solar system. Each of these perturbations must be treated either analytically or numerically, and can be measured as a change of the planet's osculating orbital elements ($a, e, i, \Omega, \varpi = \Omega + w$ and $L = \varpi + M$) determined from the perturbing function \mathfrak{R} , according to the Lagrange formula. The corresponding analytical developments of the perturbing function \mathfrak{R} as a function of the orbital elements of the two bodies considered are particularly complex. On the contrary it is easy to use the numerical integration (Runge-Kutta of the 12th order), in computing the 9-body problem (without asteroids), then of the 10-body problem (with the given asteroid). Then we determine the effect of the asteroid considered on each orbital parameter of the planet studied by simple subtraction of the two signals. We focus our efforts in performing the frequency analysis of the data, using fast Fourier transform (FFT) to determine the leading frequencies. At last we carry out a nonlinear regression in which the differential data are modeled by the least-square method following an equation of the type: $F(t) = a_o + a_1 t + a_2 t^2 + \sum_{i=1}^N A_i \sin(f_i t) + B_i \cos(f_i t) + C_i t \sin(f_i t) + D_i t \cos(f_i t)$

We have also calculated the individual effects of the leading asteroids on the Earth-Moon distance, showing that they are quite larger than the level of precision of the LLR (Lunar Laser Ranging), and the individual influences of each asteroid on the distance from the EMB to each terrestrial planet and their orientation vector as seen from the EMB, which are very important parameters in space navigation and astrometry. Below, we present below an example of our results which consists in tables with the Fourier and Poisson components for the perturbations of the orbital elements of each terrestrial planet due to each asteroid, and the corresponding curves (the initial signal, the adjustment determined by our FFT analysis and the residuals). In each case (planet, asteroid, orbital element) we find that our fit is satisfactory, since the post-fit residuals are significantly lower than the original signal.

2. RESULTS

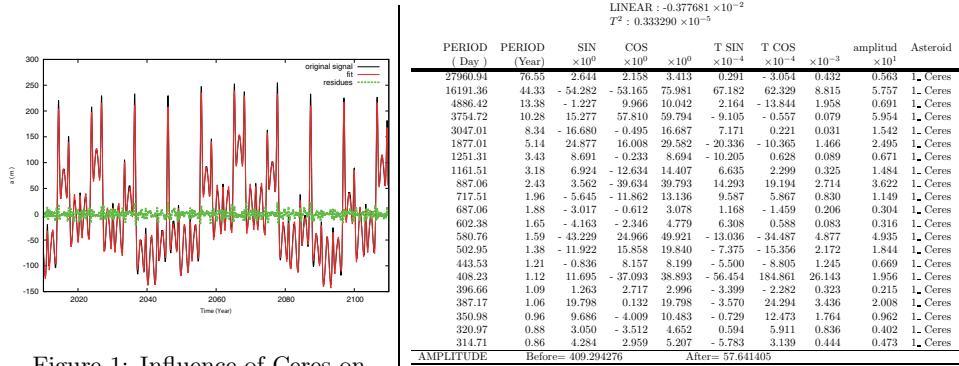


Figure 1: Influence of Ceres on Mars semi major axis (in red)

Table 1: Influence of Ceres on Mars' semi major axis. Decomposition in Fourier and Poisson series

asteroide	Mercury				Venus				Mars					
	EMB-Mercury $\times \frac{\Delta M}{M}$	asteroide	EMB-Venus $\times \frac{\Delta M}{M}$	asteroide	EMB-Mars $\times \frac{\Delta M}{M}$	asteroide	EMB-Mars $\times \frac{\Delta M}{M}$	asteroide	EMB-Mars $\times \frac{\Delta M}{M}$	asteroide	EMB-Mars $\times \frac{\Delta M}{M}$	asteroide		
1. Ceres	388.99745	1018.52641	857.47915	5710.02652	17631.53663	1941.76261	694.01010	46a. Hestia	1164.60660	111. Até	12356.22000			
2. Pallas	193.21460	413.14394	141.88830	12173.16362	370.64521	4182.51821	390.65110	111. Até	380.29540	46. Hestia	248.66100			
3. Juno	29.79749	47.96241	35.77520	869.53154	430.46224	318.86497	77.53379	3. Juno	141.31080	19. Fortuna	1485.34900			
4. Vesta	404.55975	1636.34551	586.32212	8301.55218	1455.61346	6457.48006	61.31650	165. Loreley	84.56231	24. Themis	1291.46300			
5. Hebe	61.28165	319.84333	74.70642	1275.36429	297.28411	485.497680	88. Thise	35.62771	18a. Melpomene	70.03936	3. Juno	1221.21800		
7. Iris	63.86719	130.36684	10.91732	97.25229	10.91732	10.91732	6. Hebe	32.65011	88. Thise	65.21992	6. Hebe	1109.21300		
8. Flora	66.22072	163.49586	11.77010	876.91818	955.95446	295.91801	18a. Melpomene	25.56861	28a. Bellona	49.66069	9. Metis	1058.72800		
9. Metis	66.22072	163.49586	11.77010	876.91818	955.95446	295.91801	20. Massalia	19.39845	28. Bellona	47.19845	29. Amphitrite	676.18000		
10. Hygiea	40.41443	244.13825	38.81212	1254.49565	2798.83823	2998.78852	20. Massalia	18.50000	28. Bellona	26.10475	1. Irene	529.31010		
11. Parthenope	7.79214	51.54678	7.89198	42.35986	505.30615	461.49872	45. Cybèle	12.18488	42. Egeria	26.70767	20. Massalia	349.37370		
13. Egeria	11.49283	23.51995	6.00169	1628.31059	361.99825	619.36387	45. Cybèle	11.15585	45. Eugenia	26.37660	8. Flora	342.20010		
14. Irene	12.92370	31.03000	18.34228	485.36482	271.56245	21.52470	2. Palas	10.79248	24. Amphitrite	26.10475	16. Loreley	311.62700		
15. Eunomia	22.84730	71.25431	64.06555	1455.53785	21.52470	2028.37351	20. Massalia	10.47764	20. Massalia	23.67039	14. Irene	263.50100		
16. Psyche	6.99493	26.21138	9.41352	229.74675	161.67922	645.65713	48. Doris	10.30252	7. Iris	23.64386	28. Bellona	261.16760		
17. Thise	2.17981	7.78699	2.21243	5.41765	68.21111	219.01810	47. Aglaia	9.39724	10. Hygiea	21.92538	7. Iris	236.89170		
18. Melpomene	6.50754	13.38326	14.99738	180.46621	105.09344	119.62900	10. Themis	8.70575	24. Themis	10.41845	2. Palas	226.52210		
19. Fortuna	25.64017	117.05157	14.50262	241.74259	1004.49812	6000.85764	16. Psyche	8.32899	65. Cybèle	10.41845	28. Bellona	189.73600		
20. Matilde	43.48745	30.19208	0.82468	10.74254	10.74254	10.74254	2.03200	2.03200	1.42205	1.42205	1.42205	138.93990		
21. Letitia	4.85095	36.77710	0.95140	13.73545	290.04054	488.48782	7. Iris	7.55930	42. Egeria	13.51992	13. Egeria	185.14450		
22. Kalliope	2.24835	8.23317	3.41134	335.01872	66.82759	193.04942	14. Irene	7.49398	19. Fortuna	12.13824	10. Hygiea	168.26760		
24. Themis	8.29007	35.55968	0.45115	40.77149	263.71601	2145.52121	10. Hygiea	7.42384	9. Metis	10.88295	15. Eunomia	165.01150		
28. Bellona	13.00662	27.20348	20.89924	241.39297	481.31076	666.97955	31. Euphrosyne	5.37681	12. Ceres	9.52207	16. Psyche	152.56800		
29. Amphitrite	14.22070	50.16000	21.52470	273.76838	582.30369	1189.20269	15. Eunomia	4.71414	16. Psyche	9.41188	65. Cybèle	144.84320		
31. Euphrosyne	1.12241	22.00084	5.45233	10.74254	10.74254	10.74254	4. Vesta	4.70596	49. Pales	9.38890	45. Eugenia	129.91330		
35. Eugenia	4.08374	12.87875	6.41856	289.28267	119.82741	613.18380	13. Egeria	3.79445	48. Doris	6.92858	31. Euphrosyne	86.37314		
46. Hestia	17.96786	7.74860	23.94010	139.99005	233.53623	299.17144	40.04949	18. Elektra	3.03469	18. Elektra	4.60745	10. Metis	66.72900	
47. Aglaia	0.56693	1.52783	0.98627	13.06890	14.46586	3.34038	2.55691	2.55691	4.91414	21. Leucothea	0.21603	21. Leucothea		
48. Doris	2.88513	13.51154	4.41259	137.73545	293.43396	293.43396	49. Pales	2.60837	31. Euphrosyne	3.12844	1. Ceres	39.82367		
49. Palas	1.34264	6.32408	0.74584	31.03000	10.74254	10.74254	19. Pallas	2.25191	22. Kalliope	2.01498	17. Thetis	32.21822		
52. Europa	6.43481	32.01968	1.32215	69.31637	342.59294	1296.32556	130. Elektra	1.70296	11. Parthenope	1.97254	11. Parthenope	19.60792		
65. Cybèle	7.49899	3.32215	76.66273	71.36757	325.03838	121. Hermione	1.47892	107. Camilla	1.92639	130. Elektra	16.72673			
87. Sylvia	1.35168	3.38215	3.47084	21.27100	43.60731	26.55712	121. Hermione	1.36178	15. Eunomia	1.87222	121. Hermione	14.81428		
88. Thise	4.08374	12.87875	6.41856	289.28267	119.82741	613.18380	9. Themis	0.96717	189. Phthia	1.37761	22. Kalliope	9.06023		
90. Antiope	0.49865	2.01203	0.04924	0.73147	15.30749	21.27100	107. Camilla	0.89007	21. Leucothea	0.83844	1. Irene	4.74760		
107. Camilla	1.59576	3.01203	5.00925	145.32703	188.82926	209.29009	21. Leucothea	0.56491	22. Kalliope	0.77833	42. Leucothea	4.95955		
111. Até	144.29700	397.33242	171.68672	425.37702	5169.29002	3158.32571	122. Kalliope	0.55424	17. Thetis	0.55975	107. Camilla	2.31888		
121. Hermione	0.56530	1.99023	0.94894	52.20917	20.41814	138.98333	12. Ceres	0.52483	283. Emma	0.52277	90. Antiope	1.03507		
130. Elektra	2.82965	10.37911	10.78122	110.47445	115.07393	132.25069	107. Camilla	0.47895	243. Ida	0.52822	253. Mathilde	0.87222		
165. Loreley	6.50281	27.76875	28.27952	465.46322	345.34894	313.09628	1. Ceres	0.47851	287. Sylvia	0.20355	87. Sylvia	0.50908		
189. Phthia	0.02475	0.43275	0.02475	0.06574	0.31220	0.45678	2. Palas	0.14139	253. Mathilde	0.18701	283. Emma	0.14407		
243. Ida	0.02475	0.43275	0.02475	0.06574	0.31220	0.45678	1. Ceres	0.09770	90. Antiope	0.07711	243. Ida	0.07351		
253. Mathilde	0.11630	0.19786	0.12028	4.48861	6.20194	8.82521								
283. Emma	0.35806	1.11882	1.04706	16.94125	9.49715	9.82105								

Table 2: Influence of each asteroid on each Mars orbital parameter: peak-to-peak amplitude for a 100 yr time interval

Table 3: Uncertainty of the EMB-Mercury, EMB-Venus and EMB-Mars distance due to the asteroid masses inaccuracy

3. REFERENCES

- J.G. Williams, 1984, Determining Asteroid Masses from Perturbations on Mars, ICARUS, 57, 1-13
E.M. Standish, Newhall, Williams, J.G., Folkner, W.M., 1995, JPL Planetary and lunar ephemerides, DE403/LE403. Interoffice Memorandum, IOM 314.10-127
E.V. Pitjeva, 2001, Modern numerical ephemerides of planets and the importance of ranging observations for their creation, Celestial mechanics and dynamical astronomy, 80, 249271
A. Fienga, Laskar, J., Morley, T., Manche, H., Kuchynka, P., Le Poncin-Lafitte, C., Budnik, F., Gastineau, M., and Somenzi, L., 2009, INPOP08, a 4-D planetary ephemeris: From asteroid and time-scale computations to ESA Mars Express and Venus Express contributions, A&A, 507, 1675-1686
S. Mouret, Simon, J.L., Mignard, F., and Hestroffer, D., 2009, The list of asteroids perturbing the Mars orbit to be seen during future space missions, A&A, 508, 479-489