

# NEW COMBINATION SOLUTION FOR RS POSITIONS RSC(GAOU UA)05 C 03

Y.S. YATSKIV<sup>1</sup>, A.N. KUR'YANOVA, S.O. LYTVYN<sup>2</sup>

Main Astronomical Observatory National Academy of Sciences of Ukraine

27 Akademika Zabolotnoho St, 03680 Kiev, Ukraine

e-mail: <sup>1</sup>yatskiv@mao.kiev.ua, <sup>2</sup>slytvyn@mao.kiev.ua

**ABSTRACT.** Main Astronomical Observatory of the NAS of Ukraine is involved in the work of a redetermination of the ICRF, that was initiated by the IVS. We have carried out several combination runs based on individual catalogues provided by 8 IVS Analysis Centers and have selected the final version RSC(GAO UA)05 C 03. In comparison with the ICRF and other combined catalogues, orientation parameters and uncertainty of this new combined catalogue are given.

## 1. INITIAL CATALOGUES OF RS POSITIONS

In 2004 the IVS began a program to redetermine the ICRF. The first step in February 2005 was the generation of RS catalogues in a configuration similar to the 1995 ICRF analysis. There are currently 8 initial catalogues provided by the IVS Analysis Centers (see Fig.1).

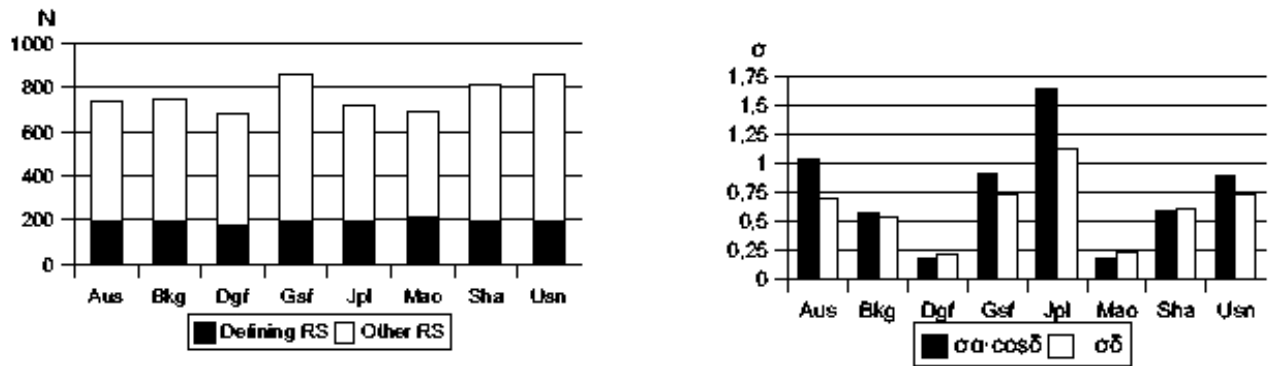


Figure 1: Number of RS (left side) and internal uncertainties (right side) for initial catalogues

## 2. NEW VERSION OF THE COMBINED CATALOGUE RSC(GAO UA)05 C 03

The Kyiv arc length method proposed by Y.Yatskiv and A. Kur'yanova, 1990 was used for construction of new combined catalogue. For this purpose several runs of combination solution based on initial catalogues mentioned above have been conducted and the catalogue RSC(GAO UA)05 C 03 has been selected as a final solution (see Lytvyn S., 2006).

This catalogue contains the positions of 953 RS, including 212 defining sources. The average values of positional uncertainties of defining RS are about 0,03 and 0,04 mas for RA and DEC respectively. The relative orientation and deformation parameters between the catalogue and the ICRF-Ext.1 are the following (in  $\mu\text{as}$ )

$$\begin{aligned} A_1 &= -1 \pm 27 & A_2 &= 2 \pm 27 & A_3 &= -21 \pm 32. \\ D_\alpha &= -2 \pm 1 & D_\delta &= 0 \pm 1 & B_\delta &= 12 \pm 24. \end{aligned}$$

To check the quality of the RSC (GAO UA) 05 C 03 we have compared it with the ICRF-Ext.1 and the RSC(GAO UA)03 C 02 (see Yatskiv, Bolotin and Kur'yanova, 2004).

In Table 1 r.m.s. differences  $d_{ij}$ , coefficients of correlation  $r_{ij}$  between the frames, and estimated "external" uncertainties  $\sigma_i$  of the ICRF( $\sigma_1$ ), RSC(GAO UA)03 C 02( $\sigma_2$ ) and the RSC(GAO UA)05 C 03( $\sigma_3$ ) are given.

	r.m.s.differences $d_{ij}$ , mas			correlations $r_{ij}$			uncertainties, mas, on condition $r_{ij} \neq 0$		
	$d_{12}$	$d_{13}$	$d_{23}$	$r_{12}$	$r_{13}$	$r_{23}$	$\sigma_1$	$\sigma_2$	$\sigma_3$
All 584 common RS									
RA	0.54	0.44	0.39	-0.74	-0.45	-0.26	0.30	0.28	0.21
Decl	0.62	0.42	0.55	-0.73	-0.15	-0.56	0.31	0.36	0.25
211 common defining RS									
RA	0.32	0.35	0.26	-0.56	-0.69	-0.22	0.21	0.15	0.18
Decl	0.31	0.30	0.30	-0.52	-0.51	-0.47	0.18	0.17	0.17

Table 1: "External" estimations of uncertainties of the reference frames. R.M.S. differences  $d_{ij}$ , correlations  $r_{ij}$ , and uncertainties,  $\sigma_i$  calculated for all and for defining RSs. Indexes are used as follows: 1- ICRF, 2- RSC (GAO UA)03 C 02, 3- RSC(GAO UA)05 C 03

## 3. CONCLUSION

We have constructed a new combined catalogue which is internally more consistent as compared with other realisations of the ICRF.

## REFERENCES

- Lytvyn S., 2006. Comparison of various runs of combination solution for constructing the GAOUA combined catalogue of RS positions., Proc. Journées System de Reference Spation-Temporals, 2005. Warsaw( this volume)
- Yatskiv Ya., Kur'yanova A. and Bolotin S., 2004. ICRF consistency check by comparison of the ICRF-Ext.1 with the GAOUA series of RS catalogues., Proc. Journées 2003, Astrometry, Geodynamics and Solar SyStem Dynamics: from milliarcseconds to microarcseconds, St. Petersburg (Russia), pp. 39-46.
- Yatskiv Ya.S., Bolotin S.L., Kur'yanova A.N., 2004. GAOUA - realizations of celestial reference frame, Kinenatics and Phisics of Celestial Bodies, Vol. 20, No 4, pp. 291-299.