Next step in Earth interior modeling for nutation

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Nutation Series

- Longer time series
 - Better adjustment of long-period terms (e.g., 18.6-yr)
 - Improvement of the formal error
 - Can choose to drop data before 1995



Some Challenges for the free nutations

- The amplitudes can only be observed due to the poor knowledge of their excitation! Only FCN free mode observed. (but resonance)
- Explain the variations of the FCN amplitude/phase

The FCN

• Detect a signal related to the FICN (see work of Lambert et al.)

The nutation after removal of the FCN and main tidal terms



rigid Earth nutation

Earth

Forced Nutations





structure of the Earth's interior for its response

normal modes rotation axis of the mantle rotation axis of the core Change in the CAB rotation axis of the inner core interastic interaction interastic interaction interaction



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- calculate rigid nutations (precision better than observation precision) from celestial mechanics
- 2. Calculate response of planet (transfer function in frequency domain) from geophysics

• Earth: **amplifications** up to 30 mas

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Earth rotation changes due to the core; core-mantle coupling

 \rightarrow coupling mechanisms: topographic torque gravitational torque viscous torque electromagnetic torque this talk classically



Electromagnetic torque + viscous torque: dissipative Outer core electrical conductivity: known from

- Outer core electrical conductivity: known from laboratory experiments: 5 10⁵ S m⁻¹ (Stacey & Anderson 2001).
- Lowermost mantle electrical conductivity (~200 m layer at the base of the mantle): unknown but has to be lower than that of the core.

 σ_{m} = 10 S m⁻¹, 5 10⁴ S m⁻¹, 5 10⁵ S m⁻¹

- RMS of the radial magnetic field at the CMB: from surface magnetic field measurements: > 0.3 mT.
- Viscosity of the outer core fluid close to the CMB:
 - molecular viscosity: ~10⁻⁶ m² s⁻¹ (laboratory experiments and ab initio computations).
 - eddy viscosity: < 10^{-4} m² s⁻¹ (Buffett & Christensen 2007).

Constraints on the physical properties of the CMB Viscosity and Radial Uniform Magnetic Field at the CMB



Coupling model used: Buffet et al. 2002 for EM and Mathews & Guo 2005 for viscomagnetic

From Koot et al. 2010



Constraints on the physical properties of the CMB

- For EM coupling only: RMS of the radial magnetic field at the CMB: 0.7 mT or higher.
- Viscomagnetic coupling:
 - Allows for lower values of the magnetic field at the CMB.
 - Allows for lower values of mantle conductivity.
 - Outer core viscosity: ~10⁻² m² s⁻¹.
 - ➡ Very high value, unlikely to be realistic.

Electromagnetic coupling (if no topography) how to explain large RMS?

Earth rotation changes due to the core; core-mantle coupling





Laboratory: thermal conductivity of liquid iron under the conditions in Earth's core is several times higher than previous estimates (Pozzo et al. *Nature* 2012)

- \rightarrow increase of heat to be carried by conduction in this layer
- \rightarrow less heat to drive convection in the core
- decrease in electrical resistance

more generation than loss of magnetic field balance equation

Earth rotation changes due to the core; core-mantle coupling



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Topographic coupling how to explain CMB coupling constants?



Core Angular Momentum exchange due to **topographic** torque at CMB

pressure at CMB
 core-mantle boundary topography (<2km)
 Difficult, challenging, controversial
 but cannot be ruled out



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Topographic coupling



- Why only some of the topography coefficients are important?
- Related to resonance with inertial waves
 [when perturbing a rotating fluid, the particle motion is
 characterized by a low-frequency oscillation called inertial wave]
- Related to the geometry of the core and of the topography.

Analytical approach

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Research objective and strategy

- Aim at obtaining torque and associated effects on nutation
- Strategy:
 - Establish the motion equations and boundary conditions in the fluid;
 - Compute analytically/numerically the solutions;
 - Obtain the dynamic pressure as a function of the physical parameters;
 - Determine the topographic torque.
- Assessment: Comparison with Wu and Wahr (1997) who used a numerical technique



Differential equations and boundary conditions Linearized Navier-Stokes equation: The oscillations of a rotating fluid where the restoring force involves Coriolis force are inertial waves

$$\begin{cases} i \sigma_m \vec{q} + 2\vec{\hat{z}} \times \vec{q} + \nabla \Phi = 0 \\ \text{Coriolis} \\ \nabla \cdot \vec{q} = 0 \end{cases} \quad \text{where } \Phi = \frac{\phi}{\Omega^2 L^2} \text{ and } \phi = \frac{p}{\rho_f} + \chi.$$

Process for obtaining the solutions and the torque

- \vec{q} as a function of $\nabla \Phi$: $\vec{q} = \frac{-i\sigma_m}{4-\sigma_m^2} \left[\nabla \Phi \frac{2}{i\sigma_m} \vec{\hat{z}} \times \nabla \Phi \frac{4}{\sigma_m^2} (\vec{\hat{z}} \cdot \nabla \Phi) \vec{\hat{z}} \right]$
- Expression for Φ : $\Phi = \sum_{l=1}^{k} P_{lk}(\frac{\sigma_m}{2}) Y_l^k(\vartheta, \lambda).$
- Expression of \vec{v} in function of χ :

$$egin{array}{rcl} v_1+iv_2&=&-rac{i}{\Omega}\left(rac{\partial\chi}{\partial x}+irac{\partial\chi}{\partial y}
ight)\ v_3&=&rac{i}{\Omega}rac{d\chi}{dz} \end{array}$$

$$\vec{\Gamma}_{topo} = \vec{\Gamma}_{0} + \int \int_{CMB} \vec{r} \times \vec{n} \rho_{f}(\chi - \phi) \, dS = \vec{\Gamma}_{0} + \vec{\Gamma}_{topo}^{\chi} + \vec{\Gamma}_{topo}^{\phi}$$

$$\Gamma_{1}^{\phi} = -\frac{i}{2} \sum_{l=1}^{i_{max}} \sum_{k=-l}^{i} (-1)^{k} \left[\sqrt{l-k} \sqrt{l+k+1} \epsilon_{l}^{-k-1} + \sqrt{l-k+1} \sqrt{l+k} \epsilon_{l}^{-k+1} \right] P_{lk}(\frac{\sigma_{m}}{2} (a_{l}^{k}))$$

$$\Gamma_{2}^{\phi} = \frac{1}{2} \sum_{l=1}^{l_{max}} \sum_{k=-l}^{l} (-1)^{k} \left[-\sqrt{l+k+1} \sqrt{l-k} \epsilon_{l}^{-k-1} + \sqrt{l-k+1} \sqrt{l+k} \epsilon_{l}^{-k+1} \right] P_{lk}(\frac{\sigma_{m}}{2} (a_{l}^{k}))$$

$$\tau_{3}^{\phi} = i \sum_{l=1}^{l_{max}} \sum_{k=-l}^{l} (-1)^{k} k \epsilon_{l}^{-k} P_{lk}(\frac{\sigma_{m}}{2}) a_{l}^{k}$$

Final expressions for nutation







frequency values for getting determinant =0 → resonances

Final expressions for LOD

Equation for obtaining the analytical expression of a_{l}^{k} in function of the topography coefficients ε_{l}^{k}

$$\sum_{l,k} Y_l^k \left[k P_{lk} \left(\frac{\sigma_m}{2} \right) - \left(1 - \frac{\sigma_m^2}{4} \right) P_{lk}' \left(\frac{\sigma_m}{2} \right) \right] a_l^k - 2 \left(1 - \frac{\sigma_m^2}{4} \right) \sum_{n=1} m \epsilon_n^m Y_n^m m_3 = 0$$

functions of the frequency

This is very similar but simpler than for nutation.







Earth rotation changes due to the core; core-mantle coupling

+ Core stratification





Topography, stratification, and magnetism

• chemical interactions between the core and the mantle





Stratification and magnetism

silicate mantle

iron core CMB

motion almost parallel to constant density surfaces

little change in density and the resulting buoyancy forces are weak







Earth rotation changes due to the core; inner core-outer core coupling

 \rightarrow coupling mechanisms:

- topographic torque
- gravitational torque
- viscous torque
- electromagnetic torque



Constraints on the physical properties of the ICB



Constraints on the physical properties of the ICB

- No solution for a purely EM coupling.
- Outer core viscosity: ~ 10 m² s⁻¹: unrealistic!
- RMS of the mag. field at the ICB: 6-7 mT.

Another mechanism is required to explain the observed damping of the FICN mode ! Inner core viscous deformation?





Koot & Dumberry EPSL (2011)

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 Inner core viscous deformation?
 Koot & Dumberry EPSL (2011)
 For Inner core viscosity: ~2-7 10¹⁴ Pa s.
 RMS of the mag. field at the ICB: 4.5 6.5 mT



Modelling the Earth's rotation

- Interior Realistic Model Rotation
- Current model IAU2000 (Mathews et al. 2002)
 - interior properties summarised in a set of parameters
 - poorly known parameters are estimated, improving knowledge of the Earth's interior (Koot et al. 2010, 2011...)
 - other parameters are computed for a spherical Earth
- Former model IAU1980 (Wahr 1981)
 - full consideration of the polar flattening
 - disregarded non-hydrostaticity which affects the FCN period (Gwinn et al. 1986), eventually discarded
 - since then refined (e.g. Huang et al. 2011), now working on non-hydrostaticity and associated triaxiality



Non-hydrostaticity & Triaxiality



Work of Antony Trinh

- **Spectral** analysis of the equations of **continuum mechanics**
- Rotation perturbations modelled as infinitesimal toroidal degree-1 displacement
- ODE submatrix:
 - spherical, non-rotating
 - biaxial, rotating
 - triaxial, rotating, convecting





Better understanding of the Earth interior!

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