



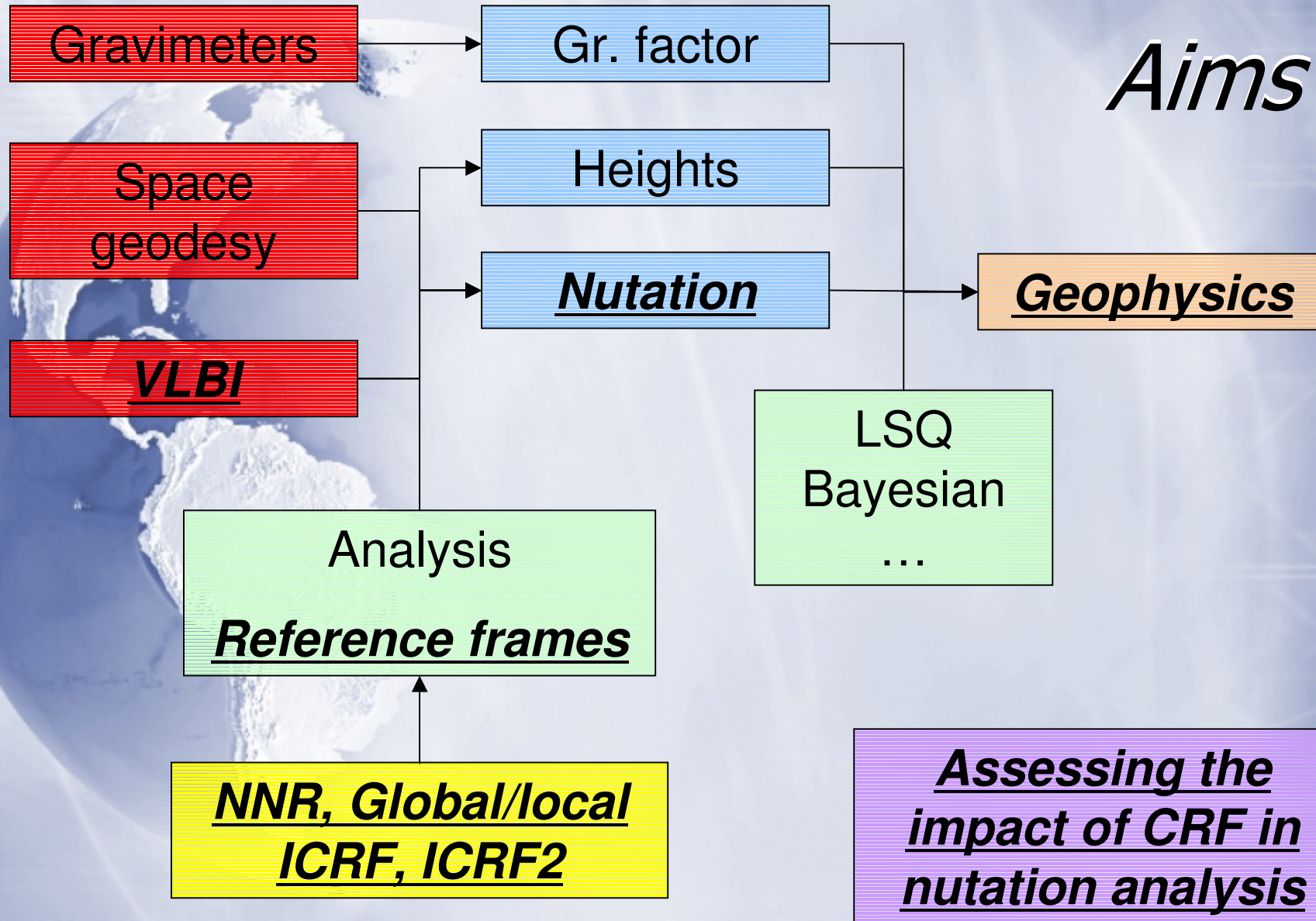
*Earth's interior with VLBI:  
... and the celestial frame?*

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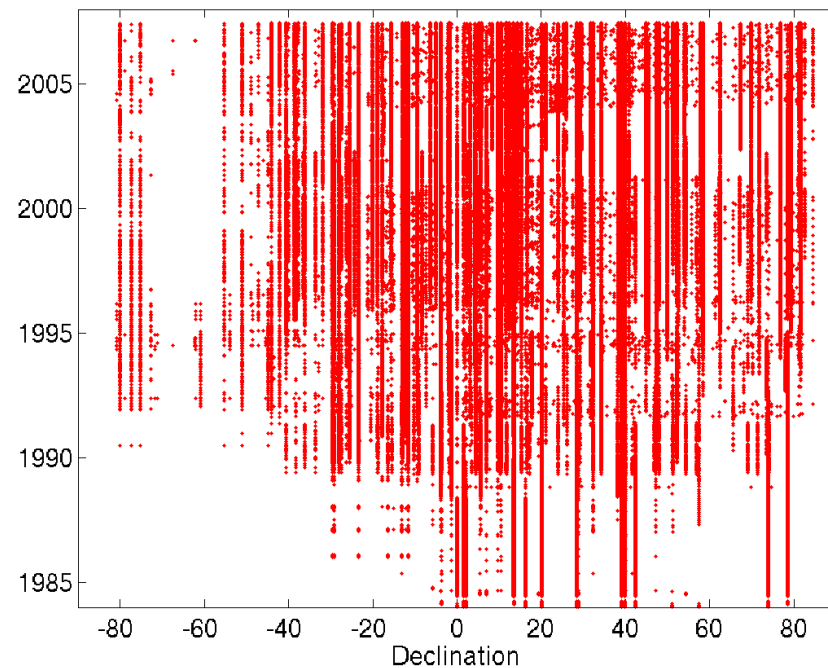
*<sup>(2)</sup>Royal Observatory of Belgium*

# *Aims*



# *Observed sources*

- ✧ The observational history is not uniform
- ✧ For global VLBI analysis, attention has to be paid to the underlying celestial reference frame
- ✧ Several strategies can be adopted: we propose to try some of them







# *Solution characteristics*

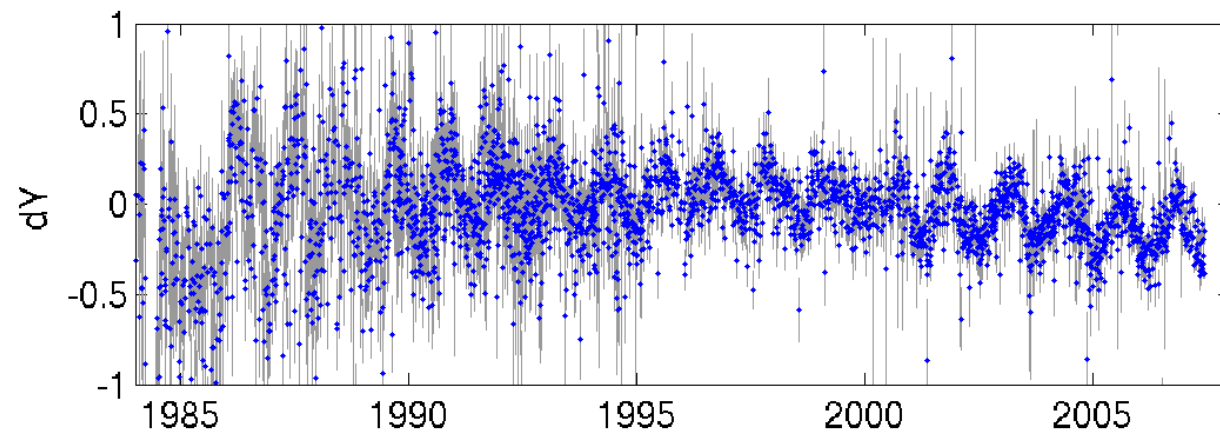
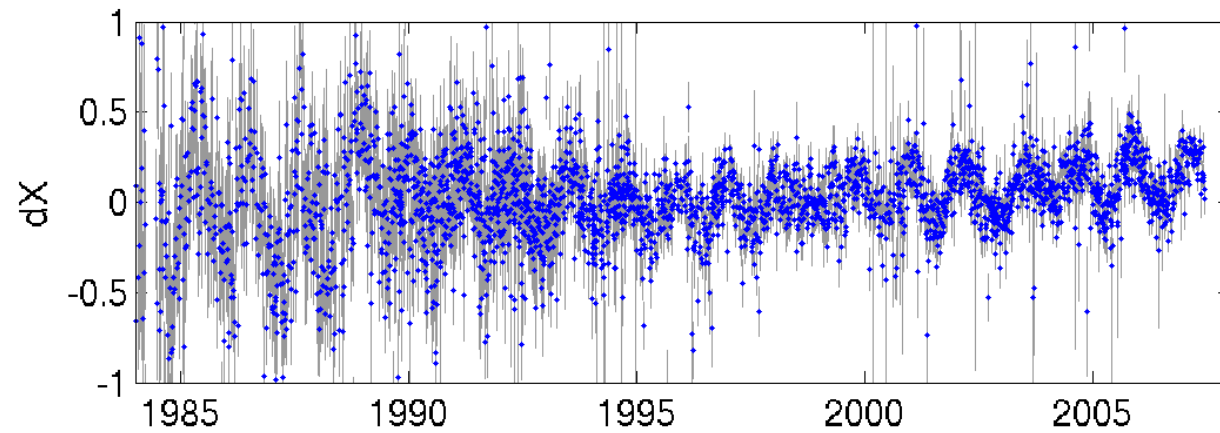
- ✧ Principle: several VLBI solutions
  - ✧ Identical...
  - ✧ ... except for the handling of radio source positions
- ✧ 2995 sessions ( $\sim 3.4$  Mdelays) spanning 1984.0-2007.5
- ✧ Terrestrial frame
  - ✧ Most station coordinates as global
  - ✧ Some modeled by splines (e.g., Fairbanks)
  - ✧ NNR/T on 26 sites wrt VTRF 2005
- ✧ EOPs
  - ✧ PM, UT1, nutation, PM rates, LOD as local
  - ✧ IAU 2000 resolutions, including the NRO-based transfo
- ✧ Atmosphere
  - ✧ NMF
  - ✧ Zenith delay every 20 mn
  - ✧ Gradients every 6 hr
- ✧ Solve 2006.06.08

# *Solution characteristics*

- ✧ Celestial frame
  - ✧ Elevation cut-off 6°
  - ✧ Split between global and local?
  - ✧ Subset for the NNR? ← ICRF defining sources, 247 stable of Feissel-Vernier et al. (2006)... Please make your choice

<b>Name</b>	<b>Fixed</b>	<b>Global</b>	<b>Local</b>	<b>NNR</b>	<b>Postfit rms (ps)</b>	<b>Rms dX/dY (μas)</b>
<b>A: Fixed</b>	816	0	0	-	24.0	165/167
<b>B: Glo 212</b>	0	816	0	ICRF 212	23.6	166/173
<b>C: Glo 247</b>	0	816	0	MFV 247	23.6	161/169
<b>D: Loc 212</b>	0	521	295 poorly observed	ICRF 212	23.6	162/170
<b>E: Loc 247</b>	0	521	295 poorly observed	MFV 247	23.6	161/169
<b>F: Uns 212</b>	0	653	163 unstable	ICRF 212	23.2	167/168
<b>G: Uns 247</b>	0	653	163 unstable	MFV 247	23.2	166/168

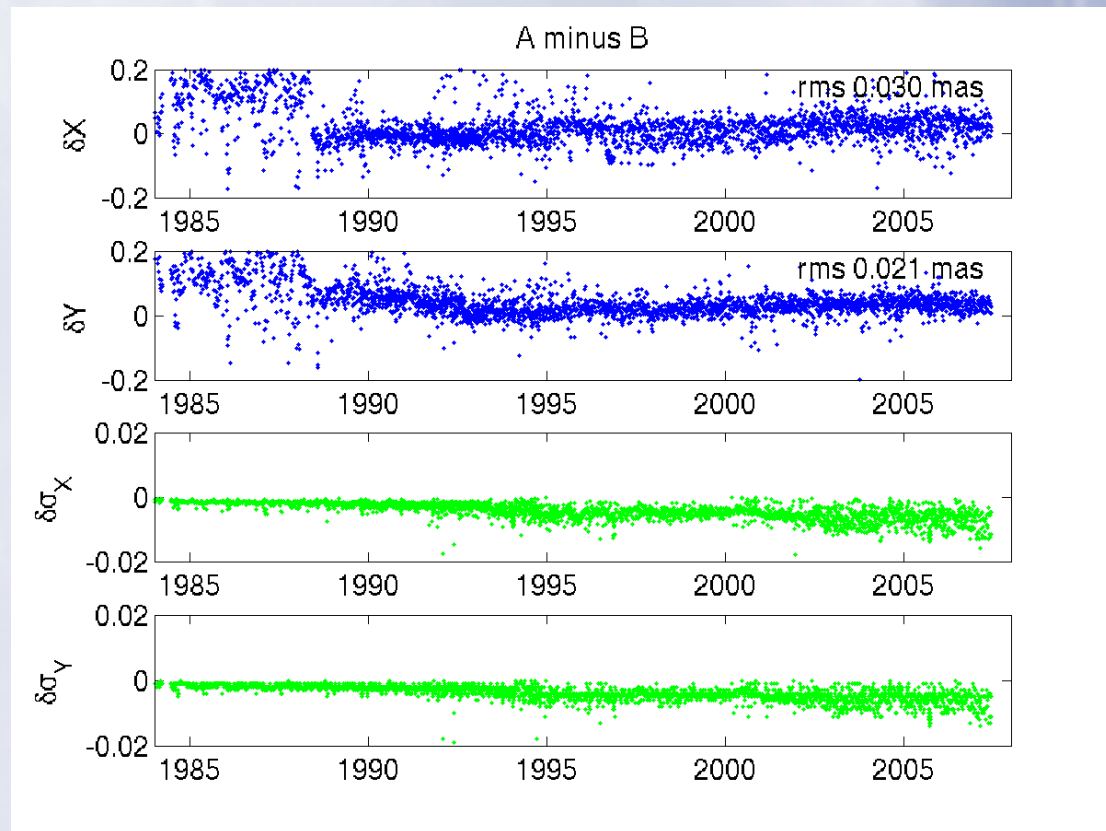
# *Nutation offsets at a glance*





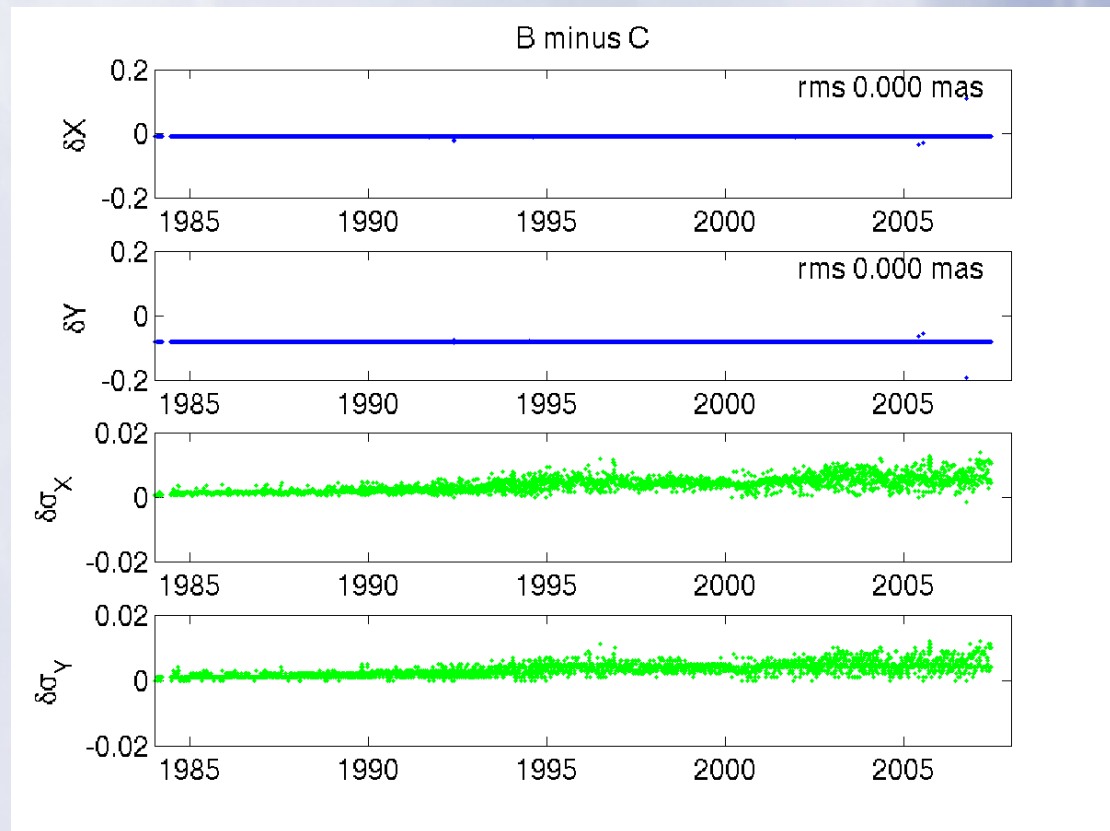
# *Differences between solutions*

✧ Fixed - Glo 212



# *Differences between solutions*

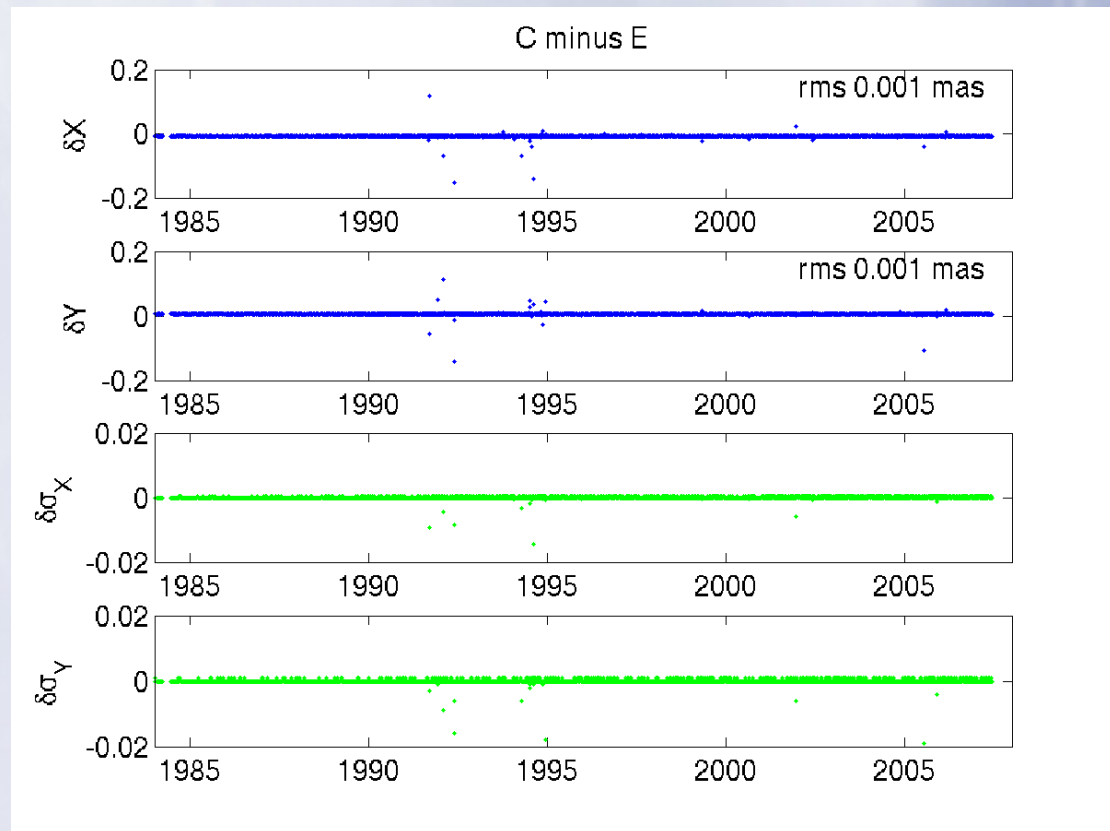
✧ Glo 212 - Glo 247





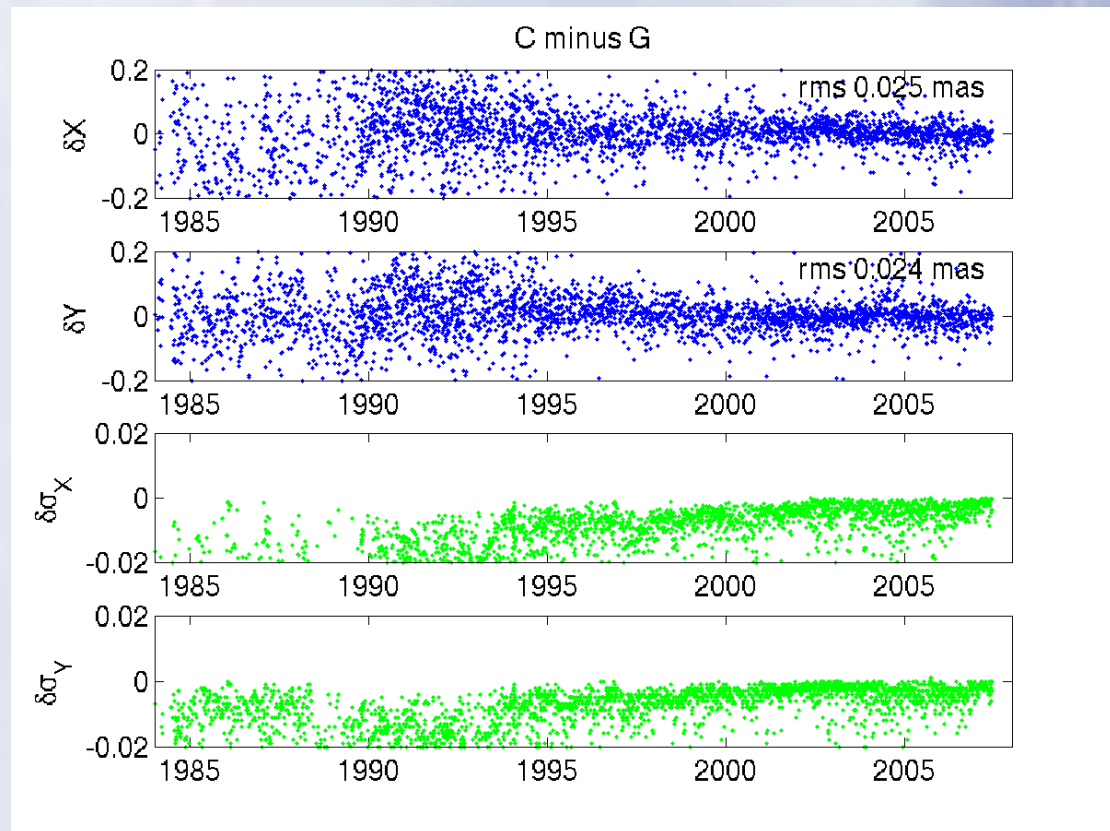
# *Differences between solutions*

✧ Glo 247 - Loc 247



# *Differences between solutions*

✧ Glo 247 - Uns 247



# *Forced nutations*

✧ LSQ fit of

$$\begin{aligned}\tilde{\eta}(t) &= \delta X(t) + i \delta Y(t), \\ &= (A^{\text{Re}} + i A^{\text{Im}}) e^{i\sigma t},\end{aligned}$$

✧ Corrections for non linear terms

✧ Erroneous values in MHB

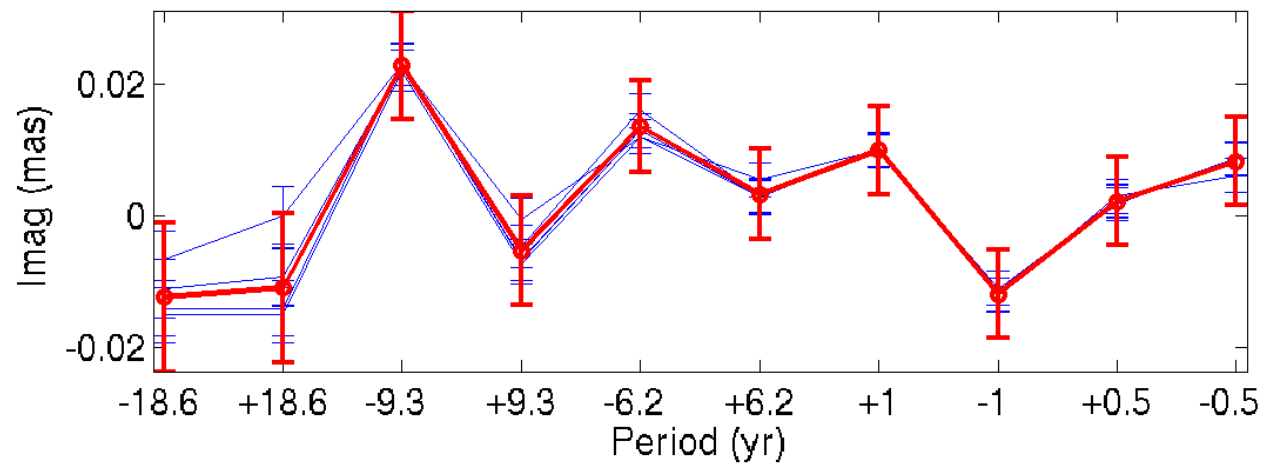
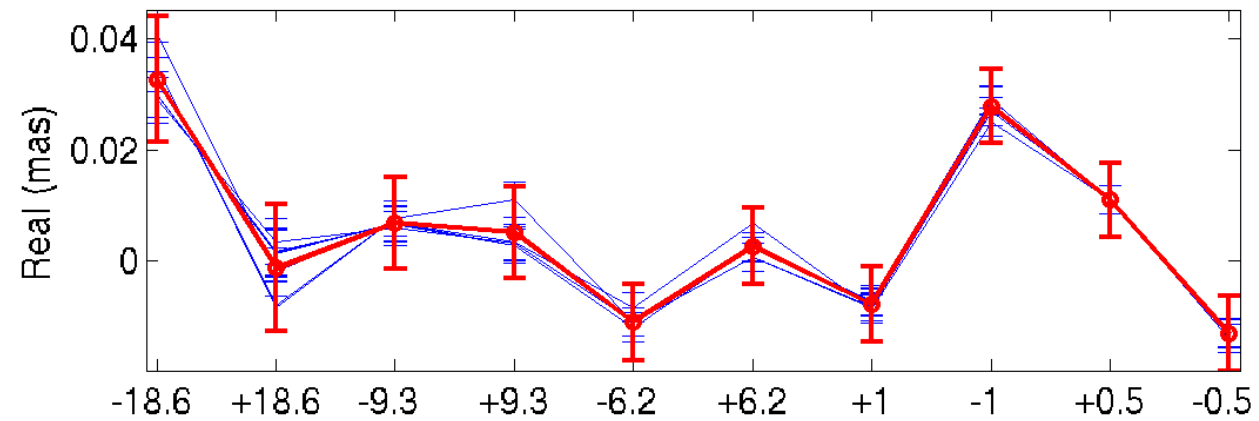
✧ Lambert & Mathews (2006)

✧ Only affect the 18.6-yr at  $\sim 30 \mu\text{as}$

✧ Pro and retro for 18.6-yr, 9.3-yr, 6.2-yr, annual, semi-, tri-, monthly, semi-, RFCN



# *Fit of forced nutations*



# *Down to the Earth's interior*

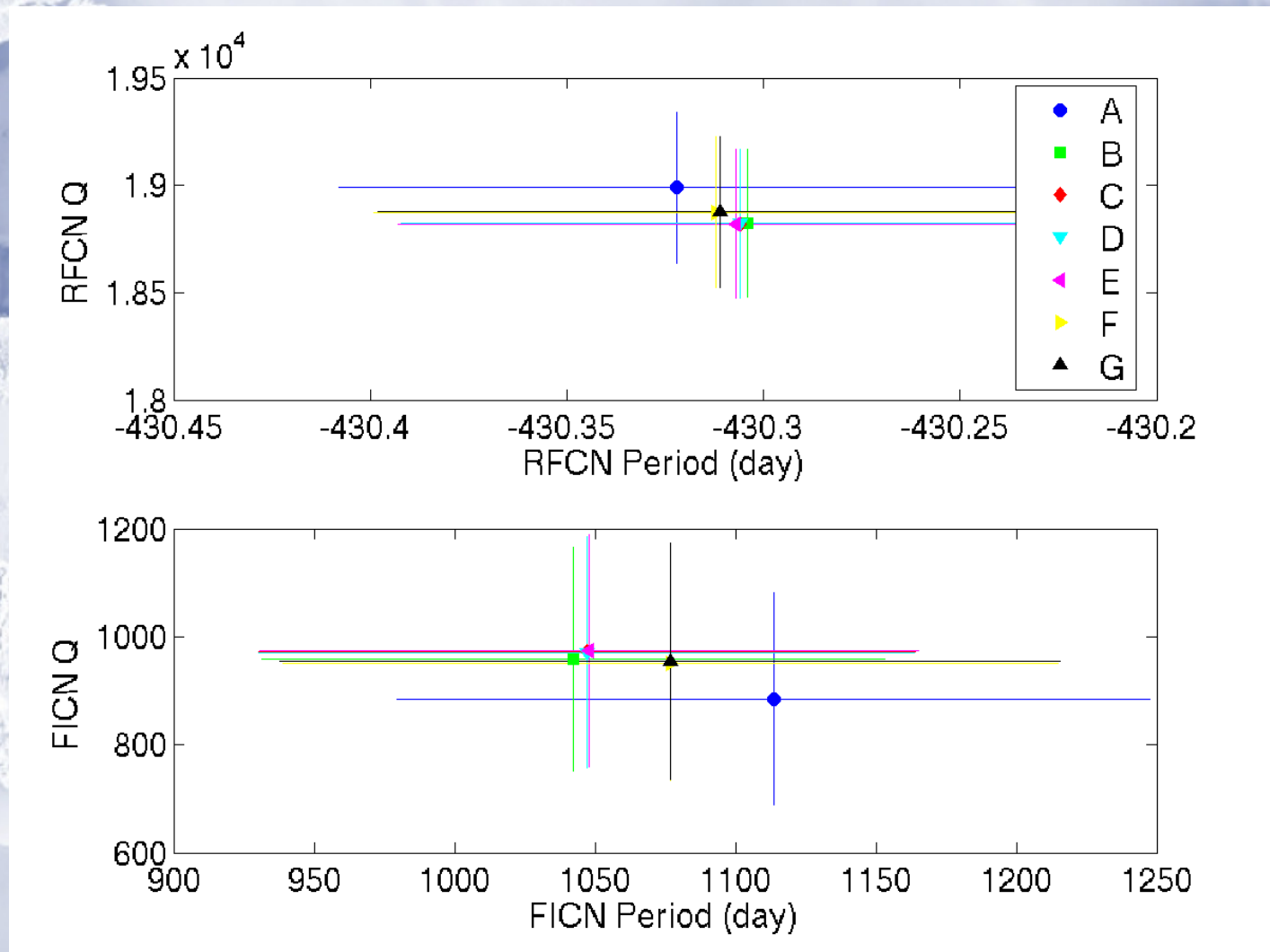
- ✧ Resonances associated w/ layers (mantle, core, inner core)
- ✧ Nutation frequency band: res. of outer and inner cores
- ✧ Resonance formula (Mathews et al. 2002)

$$\tilde{\eta}_{NR}(\sigma) = T(\sigma; e|e_R) \tilde{\eta}_R(\sigma),$$

$$T(\sigma; e|e_R) = \frac{e_R - \sigma}{e_R + 1} N_0 \left[ 1 + (1 + \sigma) \left( Q_0 + \sum_{\alpha=1}^4 \frac{Q_\alpha}{\sigma - s_\alpha} \right) \right],$$

- ✧ Periods and damping of outer and inner cores

# *Outer and inner cores resonance periods and Q*







# *Conclusion*

- ✧ Positional instability of radio sources + VLBI analysis strategy can move the nutation amplitudes by  $\sim 30 \mu\text{as}$  for the 18.6-yr,  $\sim 20 \mu\text{as}$  for the annual
- ✧ Earth's interior
  - ✧ Outer core  $\Delta P \sim 0.05$  day,  $\Delta Q \sim 1000$
  - ✧ Inner core  $\Delta P \sim 100$  days,  $\Delta Q \sim 300$
- ✧ Caution in using the different VLBI data sets since done using various analysis strategies
- ✧ Which strategy for the best internal accuracy of nutations? (MacMillan & Ma 2007)