

Long term solutions for the insolation quantities of the Earth

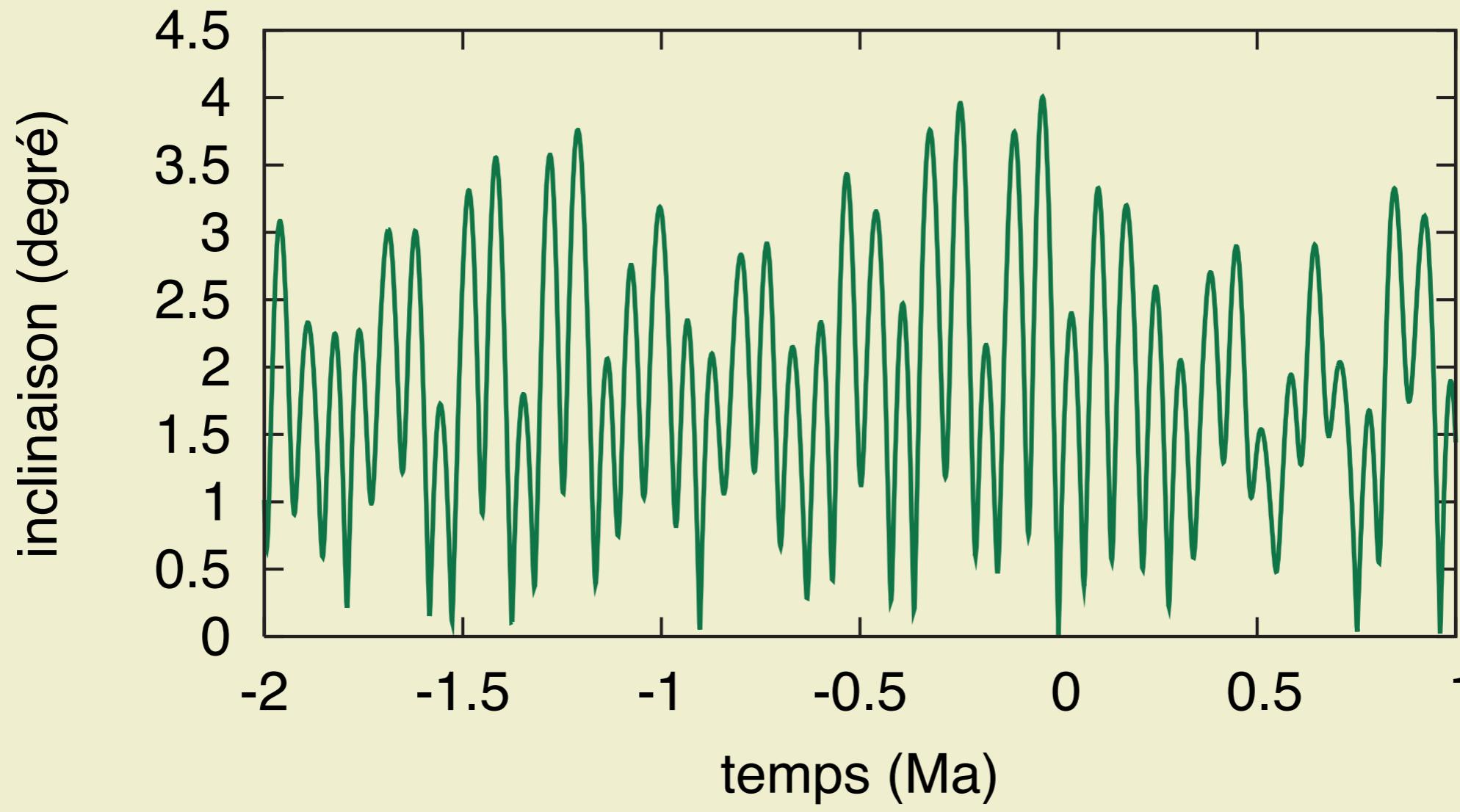
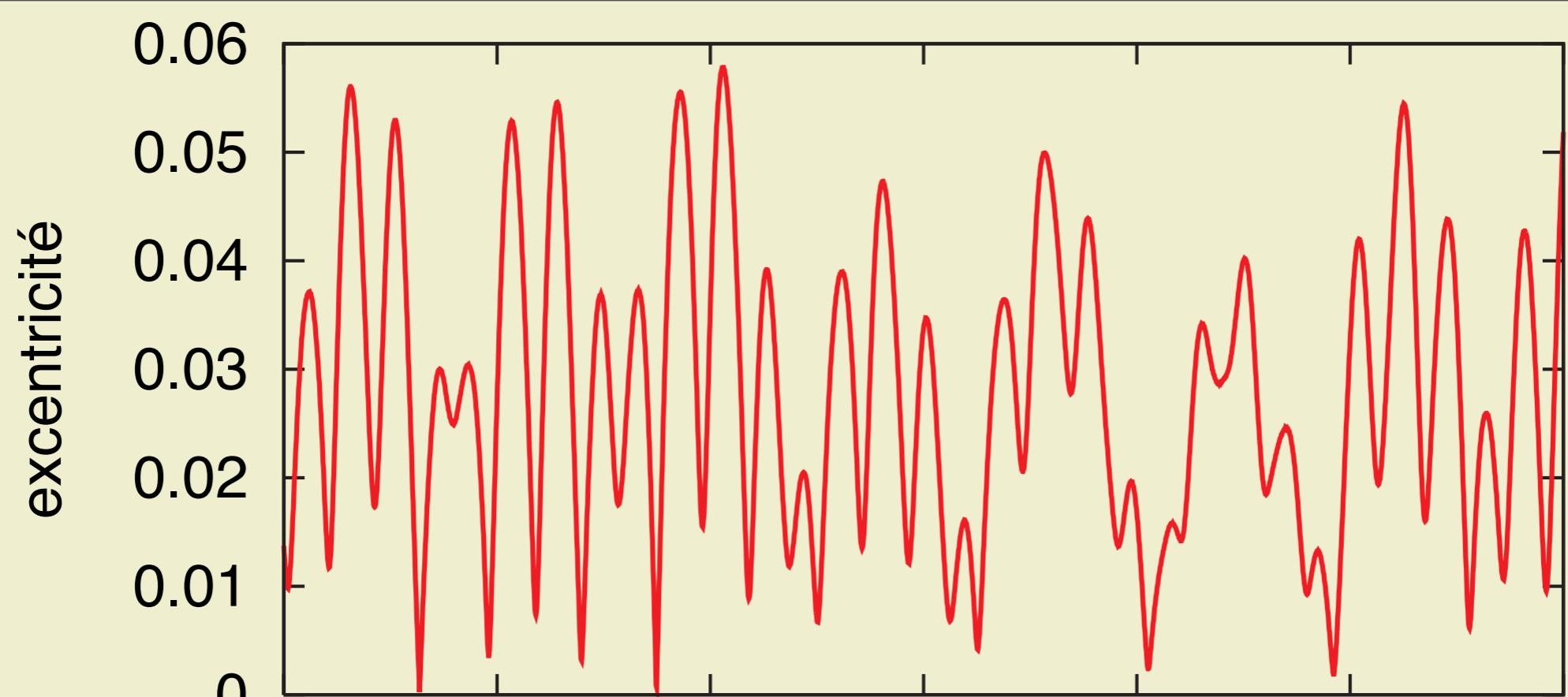
Jacques Laskar

IMCCE/CNRS, Observatoire de Paris

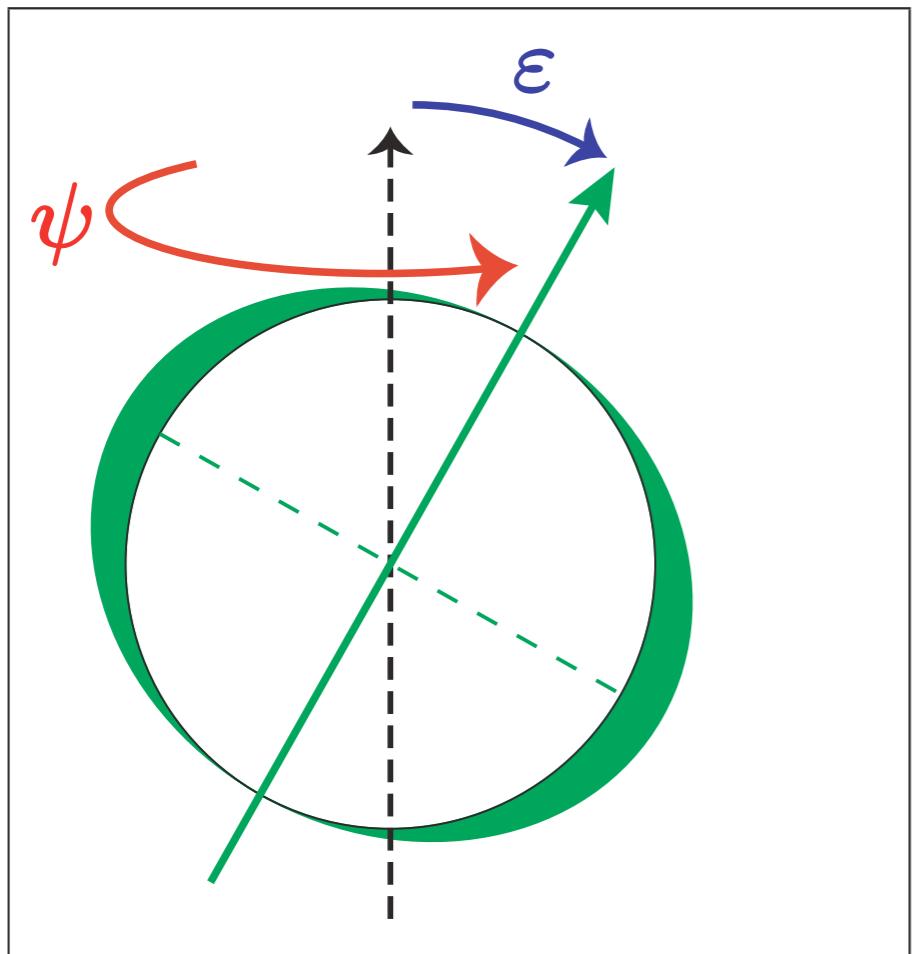
A.C.M Correia, A. Fienga, M. Gastineau, F. Joutel,
B. Levrard, H. Manche, P. Robutel

Astronomical theory of paleoclimates

- Adhémar, 1845
- Croll, 1890
- Pilgrim, 1904
- Milankovitch, 1920, 1941
- Hays, Imbrie, Shackleton, 1976



Precession (ψ) and obliquity ($X = \cos \varepsilon$)



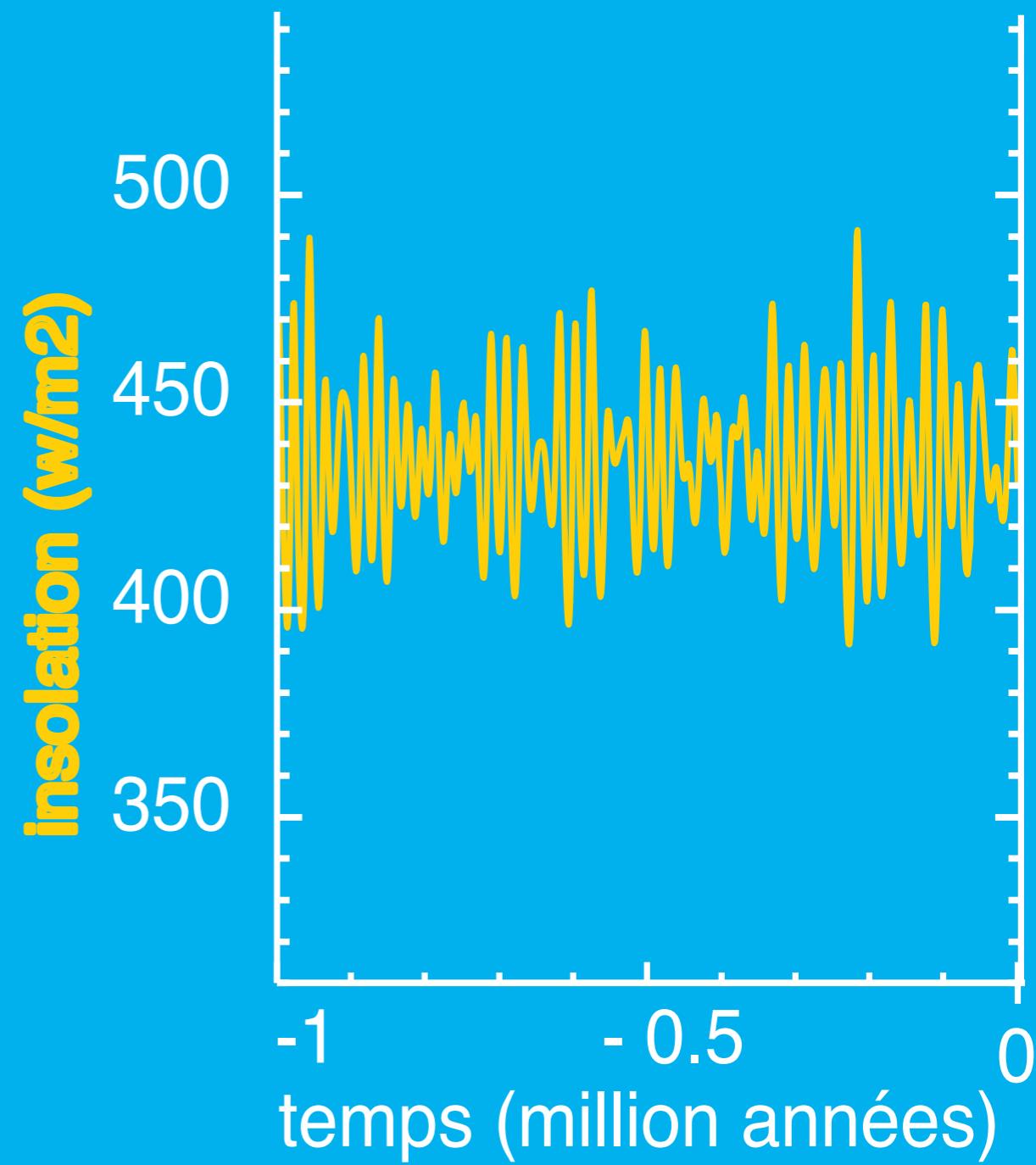
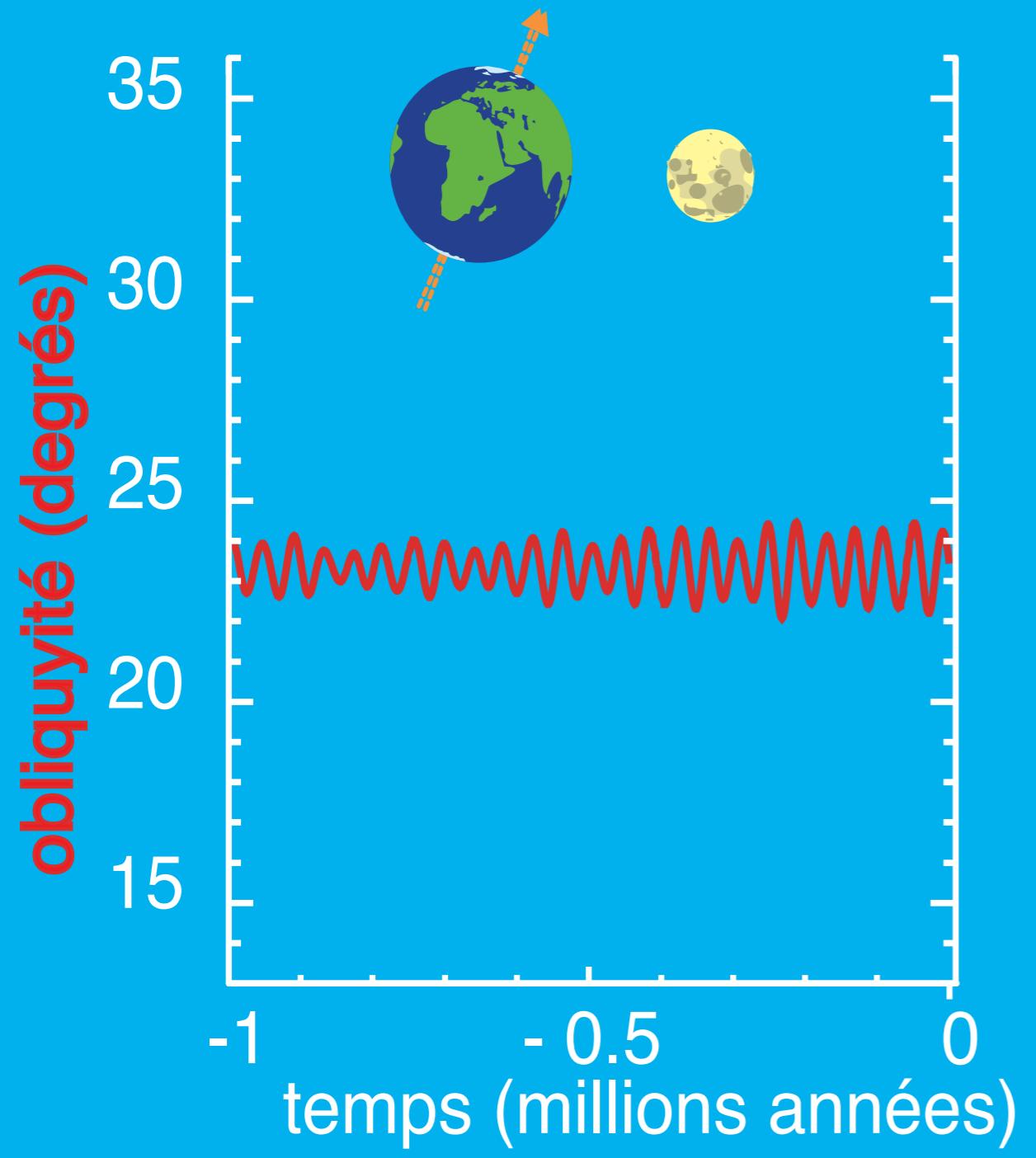
$$H = \frac{1}{2} \alpha X^2$$

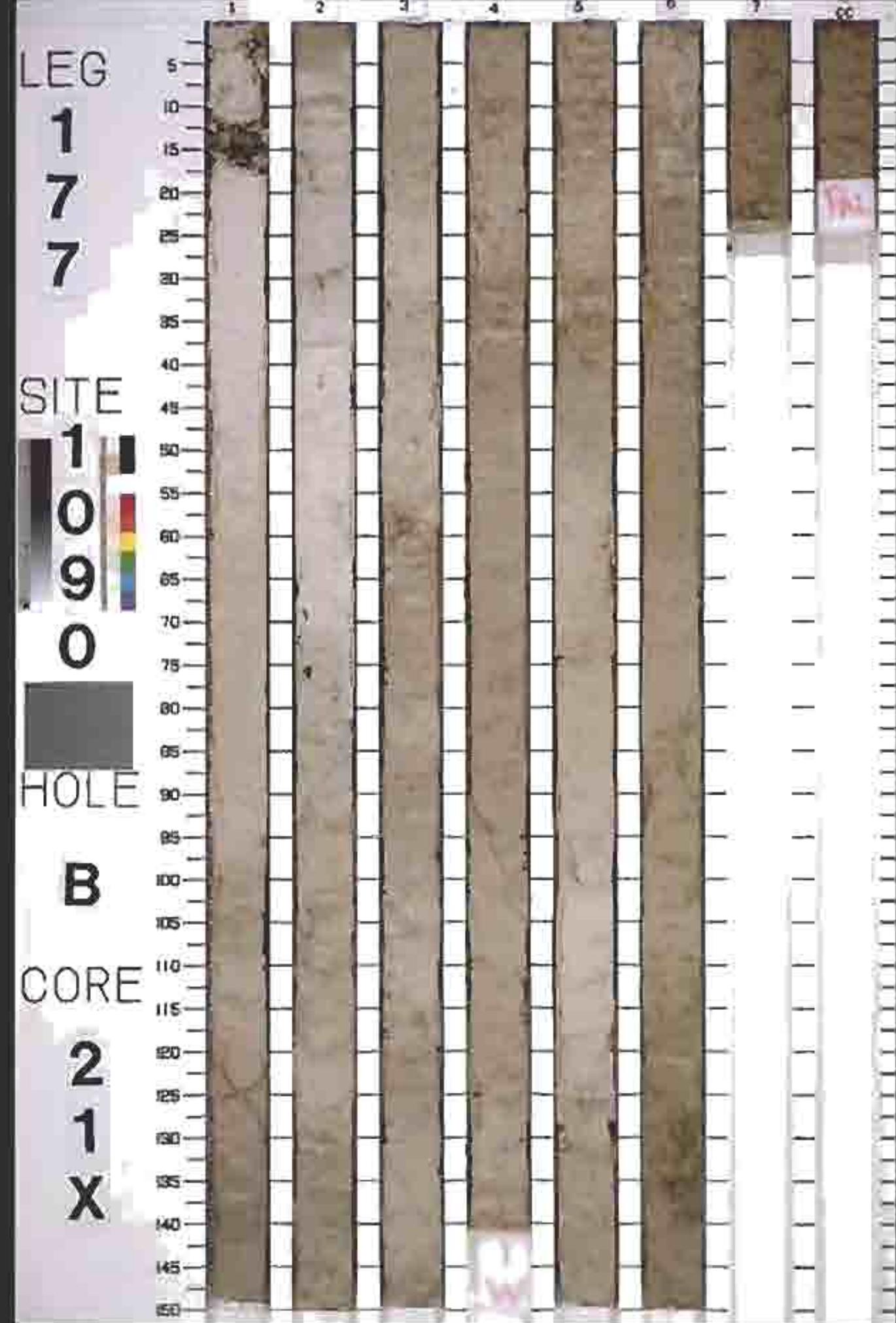
$$\left\{ \begin{array}{l} \frac{d\psi}{dt} = \frac{\partial H}{\partial X} = \alpha X_0 \\ \frac{dX}{dt} = -\frac{\partial H}{\partial \psi} = 0 \end{array} \right.$$

$$\alpha = \frac{3k^2}{2\nu} \frac{C - A}{C} \left[\frac{m_M}{a_M^3} + \frac{m_\odot}{a_\odot^3} \right]$$

Planetary perturbations

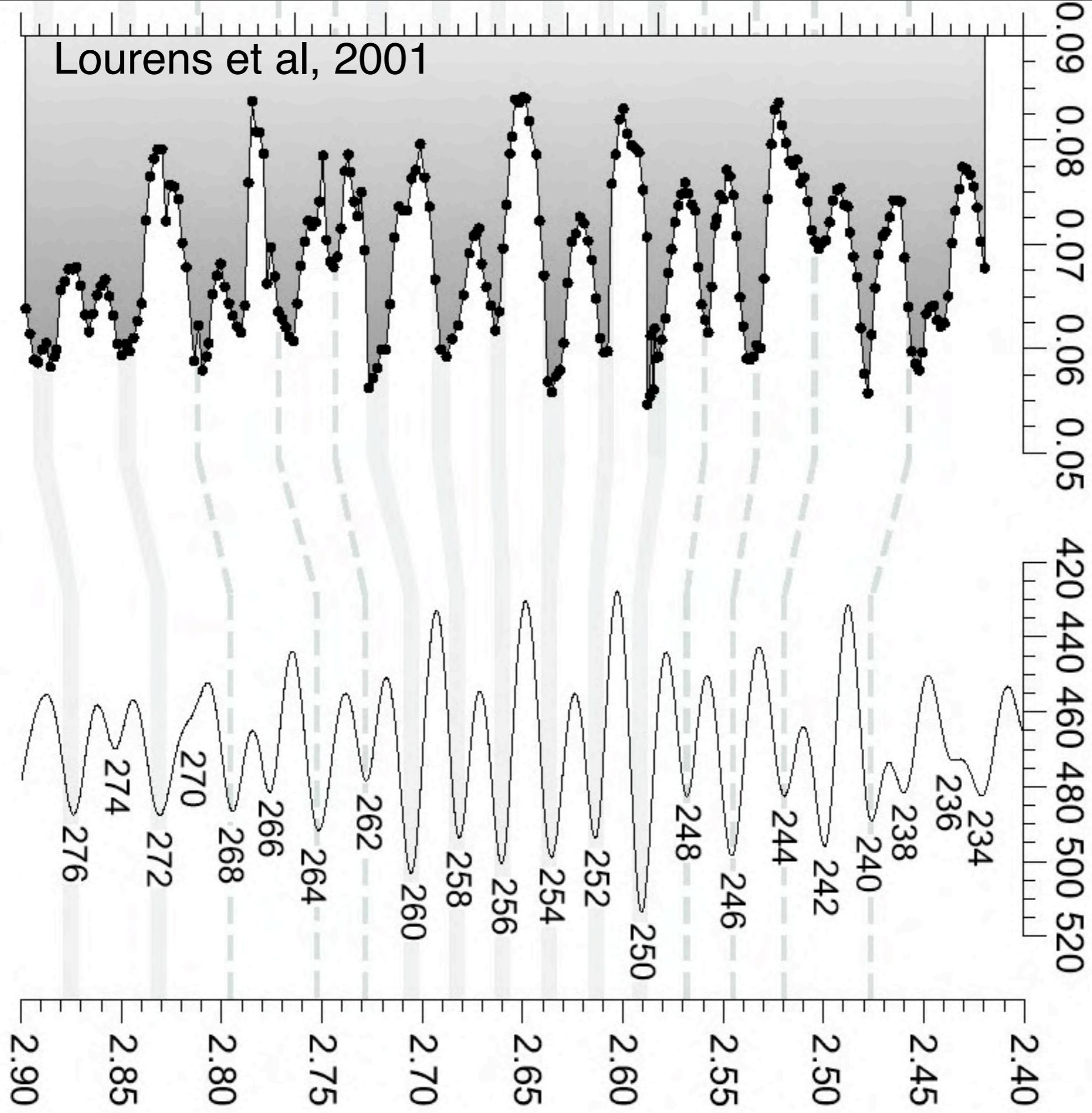
$$H(\psi, X) = \frac{1}{2} \alpha X^2 + \sqrt{1 - X^2} \sum_k \alpha_k \sin(\nu_k t + \phi_k + \psi)$$





Lourens et al, 2001

Ti/Al
Summer insolation
(65°N)



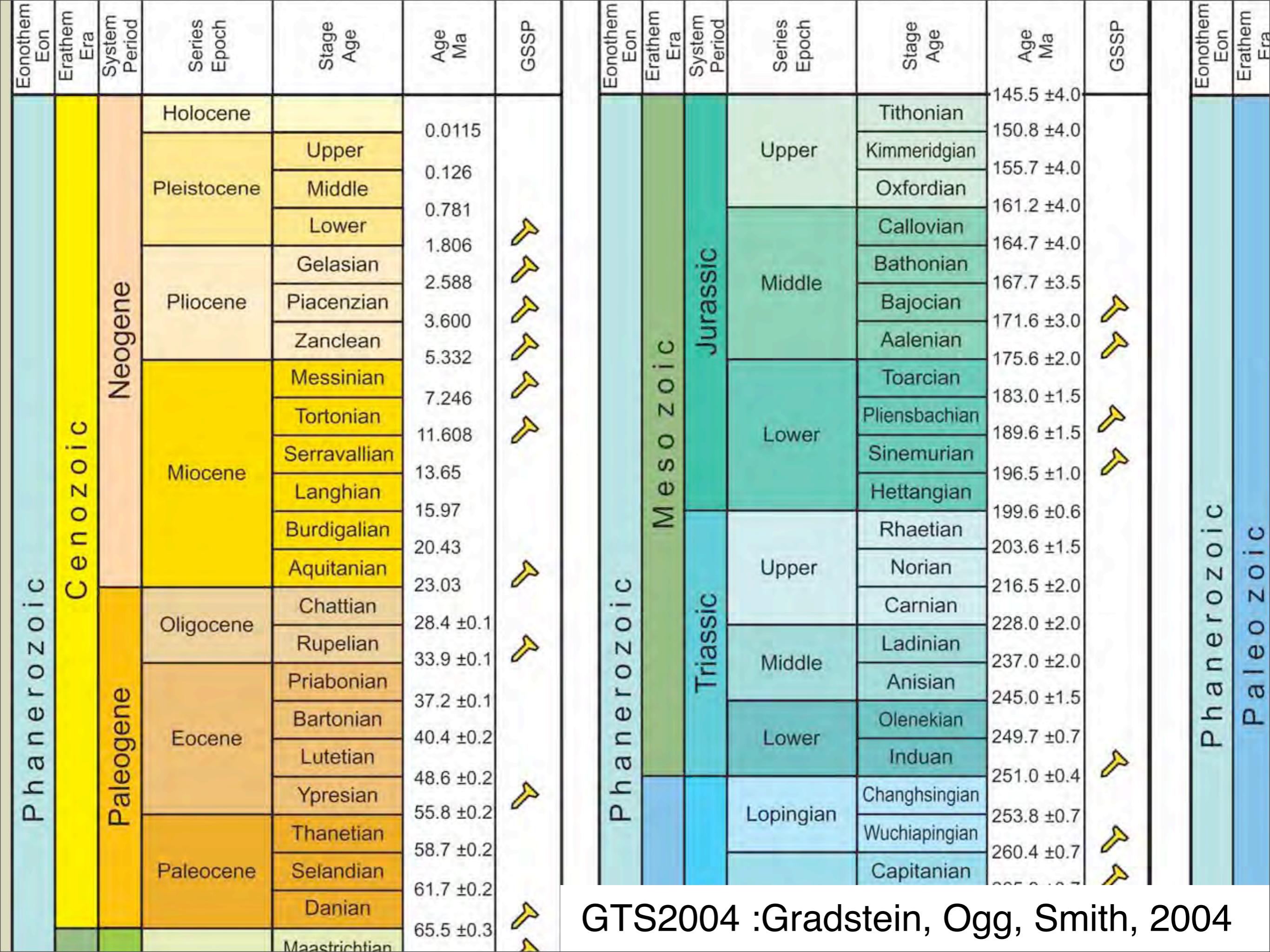
INTERNATIONAL STRATIGRAPHIC CHART

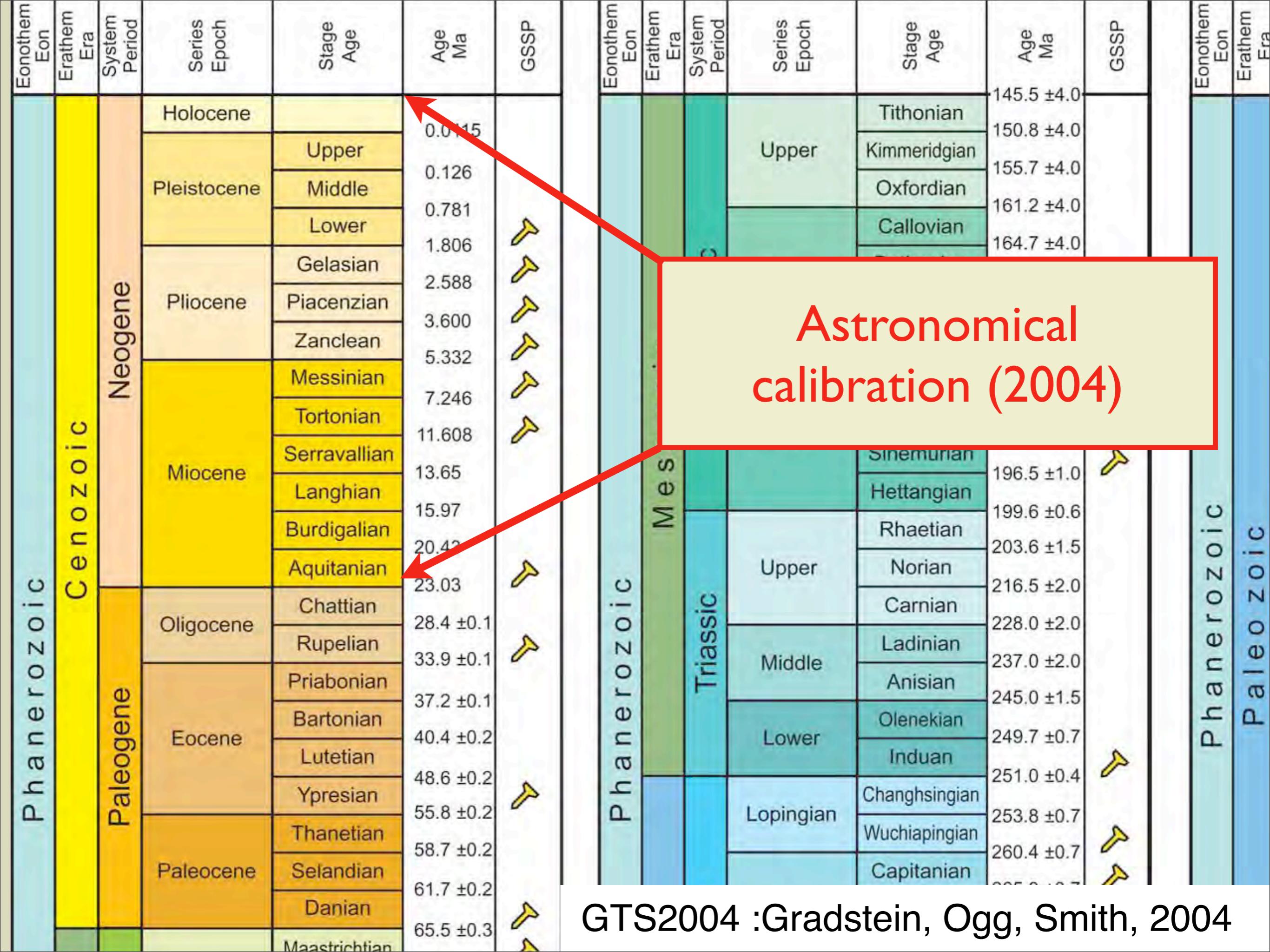
International Commission on Stratigraphy

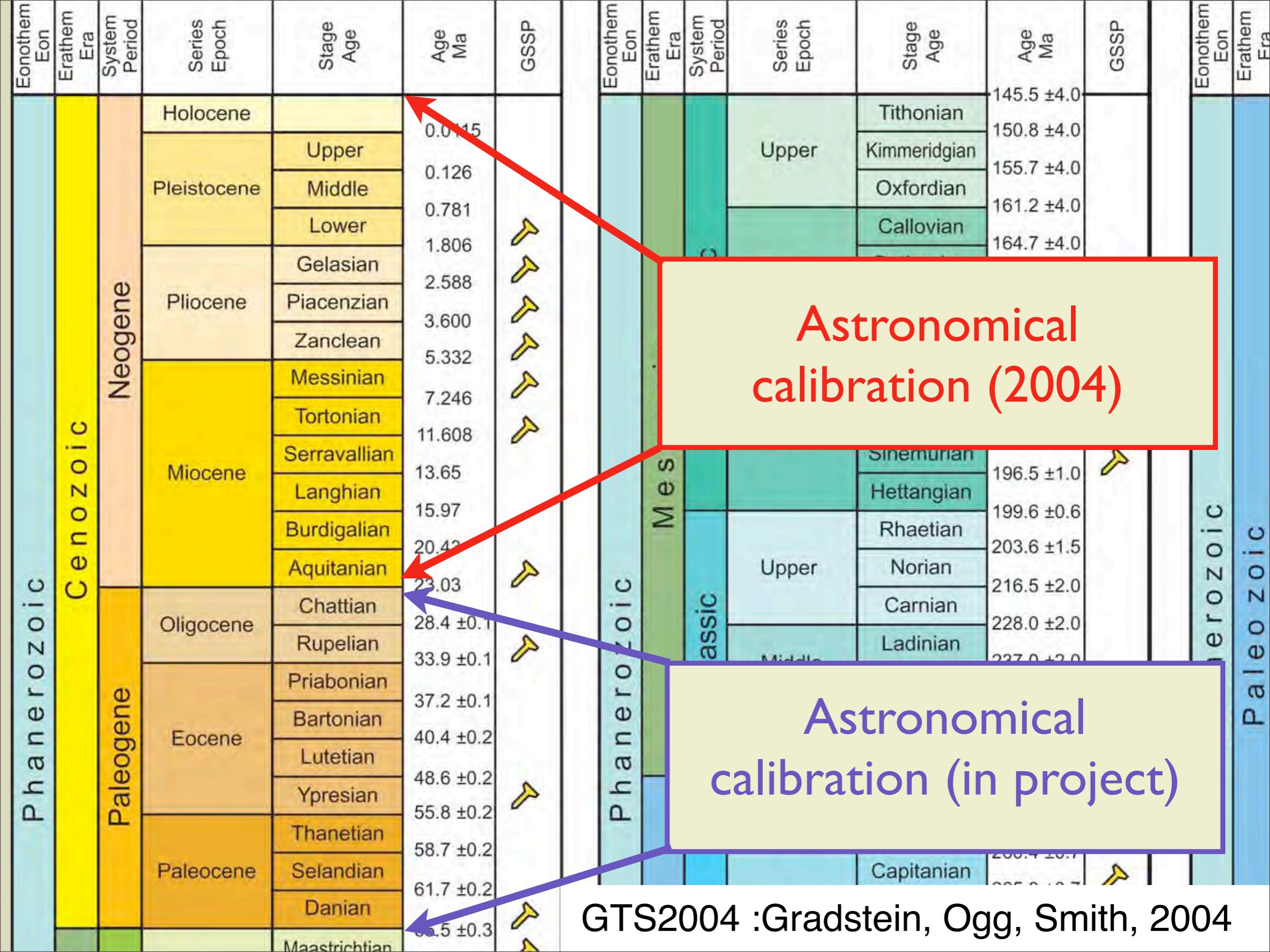
Eonothem	Eon	Erathem	Era	System	Period	Series	Epoch	Stage	Age	Age Ma	GSSP
Phanerozoic	Cenozoic	Paleogene	Neogene	Miocene	Pliocene	Holocene	Upper	0.0115	0.0115		
								Upper	0.126		
							Middle	0.781			
							Lower	1.806			
							Gelasian	2.588			
							Piacenzian	3.600			
							Zanclean	5.332			
							Messinian	7.246			
							Tortonian	11.608			
							Serravallian	13.65			
							Langhian	15.97			
							Burdigalian	20.43			
							Aquitanian	23.03			
							Oligocene	28.4 ± 0.1			
							Chattian	33.9 ± 0.1			
							Rupelian	37.2 ± 0.1			
							Priabonian	40.4 ± 0.2			
							Bartonian	48.6 ± 0.2			
							Lutetian	55.8 ± 0.2			
							Ypresian	58.7 ± 0.2			
							Thanetian	61.7 ± 0.2			
							Selandian	65.5 ± 0.3			
							Danian	70.6 ± 0.6			
							Maastrichtian	83.5 ± 0.7			
							Campanian	85.8 ± 0.7			
							Santonian	89.3 ± 1.0			
							Coniacian	93.5 ± 0.8			
							Turonian	99.6 ± 0.9			
							Cenomanian	112.0 ± 1.0			
							Albian	125.0 ± 1.0			
							Aptian	130.0 ± 1.5			
							Barremian	136.4 ± 2.0			
							Hauterivian	140.2 ± 3.0			
							Valanginian	145.5 ± 4.0			
							Berriasian				

Eonothem	Eon	Erathem	Era	System	Period	Series	Epoch	Stage	Age	Age Ma	GSSP
Phanerozoic	Palaeozoic	Carboniferous	Permian	Pennsylvanian	Upper	Tithonian	145.5 ± 4.0				
						Kimmeridgian	150.8 ± 4.0				
						Oxfordian	155.7 ± 4.0				
						Callovian	161.2 ± 4.0				
						Bathonian	164.7 ± 4.0				
						Bajocian	167.7 ± 3.5				
						Aalenian	171.6 ± 3.0				
						Toarcian	175.6 ± 2.0				
						Pliensbachian	183.0 ± 1.5				
						Sinemurian	189.6 ± 1.5				
						Hettangian	196.5 ± 1.0				
						Rhaetian	199.6 ± 0.6				
						Norian	203.6 ± 1.5				
						Carnian	216.5 ± 2.0				
						Ladinian	228.0 ± 2.0				
						Anisian	237.0 ± 2.0				
						Olenekian	245.0 ± 1.5				
						Induan	249.7 ± 0.7				
						Changhsingian	251.0 ± 0.4				
						Wuchiapingian	253.8 ± 0.7				
						Capitanian	260.4 ± 0.7				
						Wordian	265.8 ± 0.7				
						Roadian	268.0 ± 0.7				
						Kungurian	270.6 ± 0.7				
						Artinskian	275.6 ± 0.7				
						Sakmarian	284.4 ± 0.7				
						Asselian	294.6 ± 0.8				
						Gzhelian	299.0 ± 0.8				
						Kasimovian	303.9 ± 0.9				
						Moscovian	306.5 ± 1.0				
						Bashkirian	311.7 ± 1.1				
						Serpukhovian	318.1 ± 1.3				
						Visean	326.4 ± 1.6				
						Toumaisian	345.3 ± 2.1				
							359.2 ± 2.5				

Eonothem	Eon	Erathem	Era	System	Period	Series	Epoch	Stage	Age	Age Ma	GSSP
Phanerozoic	Palaeozoic	Cisuralian	Ordovician	Lower	Cambrian	Upper	Upper	Famennian	359.2 ± 2.5		
							Middle	Frasnian	374.5 ± 2.6		
							Lower	Givetian	385.3 ± 2.6		
								Eifelian	391.8 ± 2.7		
								Emsian	397.5 ± 2.7		
								Pragian	407.0 ± 2.8		
								Lochkovian	411.2 ± 2.8		
								Pridoli	416.0 ± 2.8		
								Ludlow	418.7 ± 2.7		
								Gorstian	421.3 ± 2.6		
	</										







Insolation solutions for the Earth

numerical

Orbital Solution

Le Verrier, 1856
+Hill, 1897
Brouwer & Van Woerkom, 1950

Bretagnon, 1974

Laskar, 1988, 1990

Quinn, Tremaine, Duncan, 1991
(-3Ma -> + 3Ma)

Laskar et al, 2004
(-250 Ma -> + 250 Ma orb: ~ 40 Ma pre : ~ 20 Ma)

Precession equations

Pilgrim, 1904
Milankovitch, 1920, 1941
Sharaf & Budnikova, 1967
Vernekar, 1972

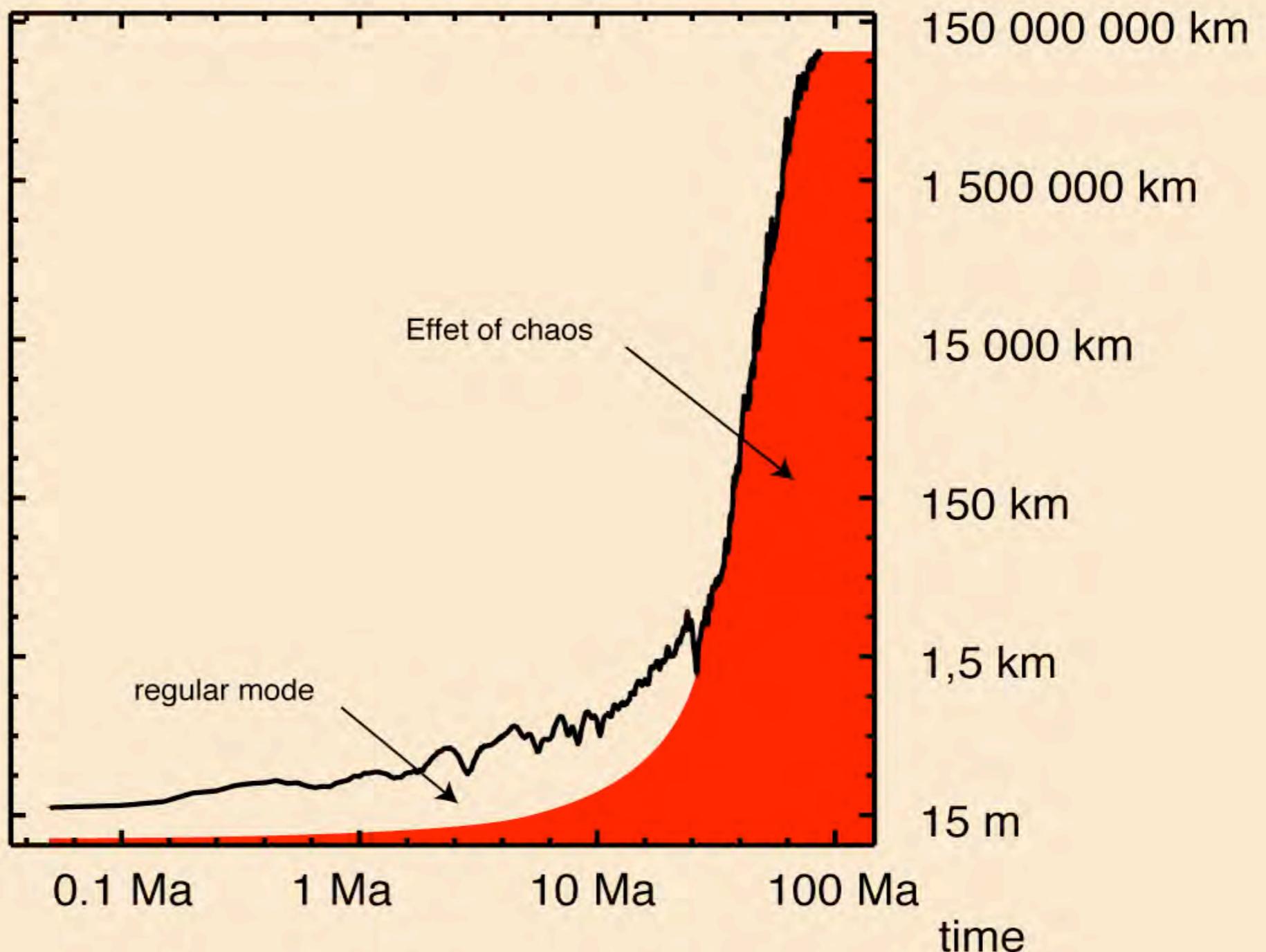
Berger, 1978

Laskar et al, 1993
(-20 Ma -> + 10 Ma)

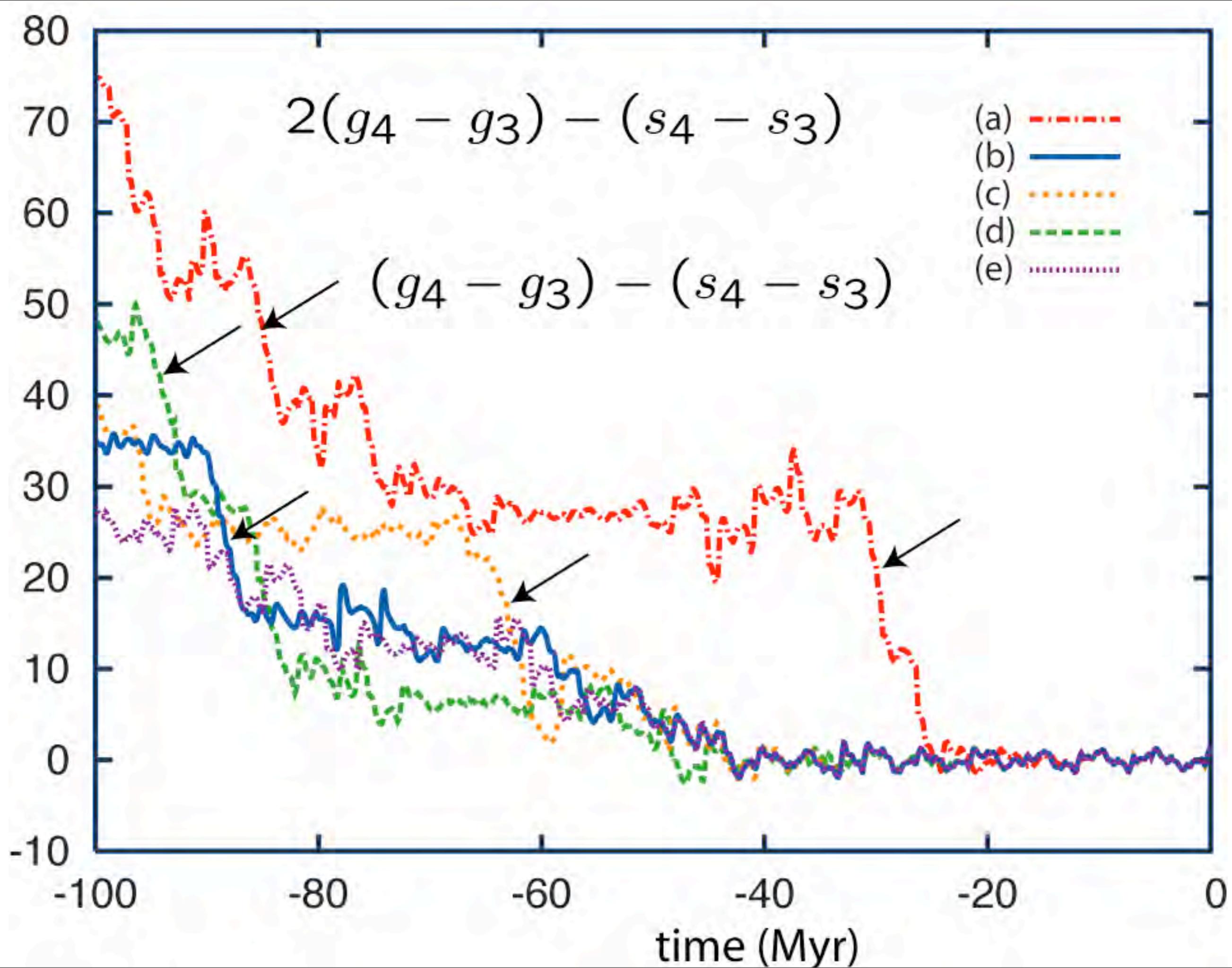
Chaotic motion of the Solar System

Secular equations : 200 Ma : Laskar (1989,1990)

Direct integration : 100 Ma : Sussman and Wisdom (1992)



$$d(T) \approx d_0 10^{T/10}$$



New Challenge :
Orbital solution over ~ 60 Myr



Improve the accuracy
by
2 orders of magnitude !

Development of a New planetary ephemeris

INPOP

(see Agnès Fienga's presentation)

Interactions in La2004, DE405, INPOP

- Newtonian (planets \leftrightarrow planets, planets \leftrightarrow asteroïds (300), asteroïds \leftrightarrow asteroïds (5))
- Relativistic corrections (planets, asteroïds)
- Non-spherical body \leftrightarrow point mass
- Sun (J2) \leftrightarrow Planets
- Earth (J2) \leftrightarrow (Moon, Sun, Venus, Jupiter)
- Moon (J2, C21, C22, S21, S22) \leftrightarrow (Earth, Sun, Venus, Jupiter)
- Deformation of extended bodies (tides) \leftrightarrow point mass
- Earth (Sun, Moon) \leftrightarrow (Moon, Sun, Venus, Jupiter)
- Moon (Spin, Earth, Sun) \leftrightarrow (Earth, Sun, Venus, Jupiter)
- Earth Shape \leftrightarrow Moon shape (torque exerted by the Moon)

eccentricity of the Earth (La2004)

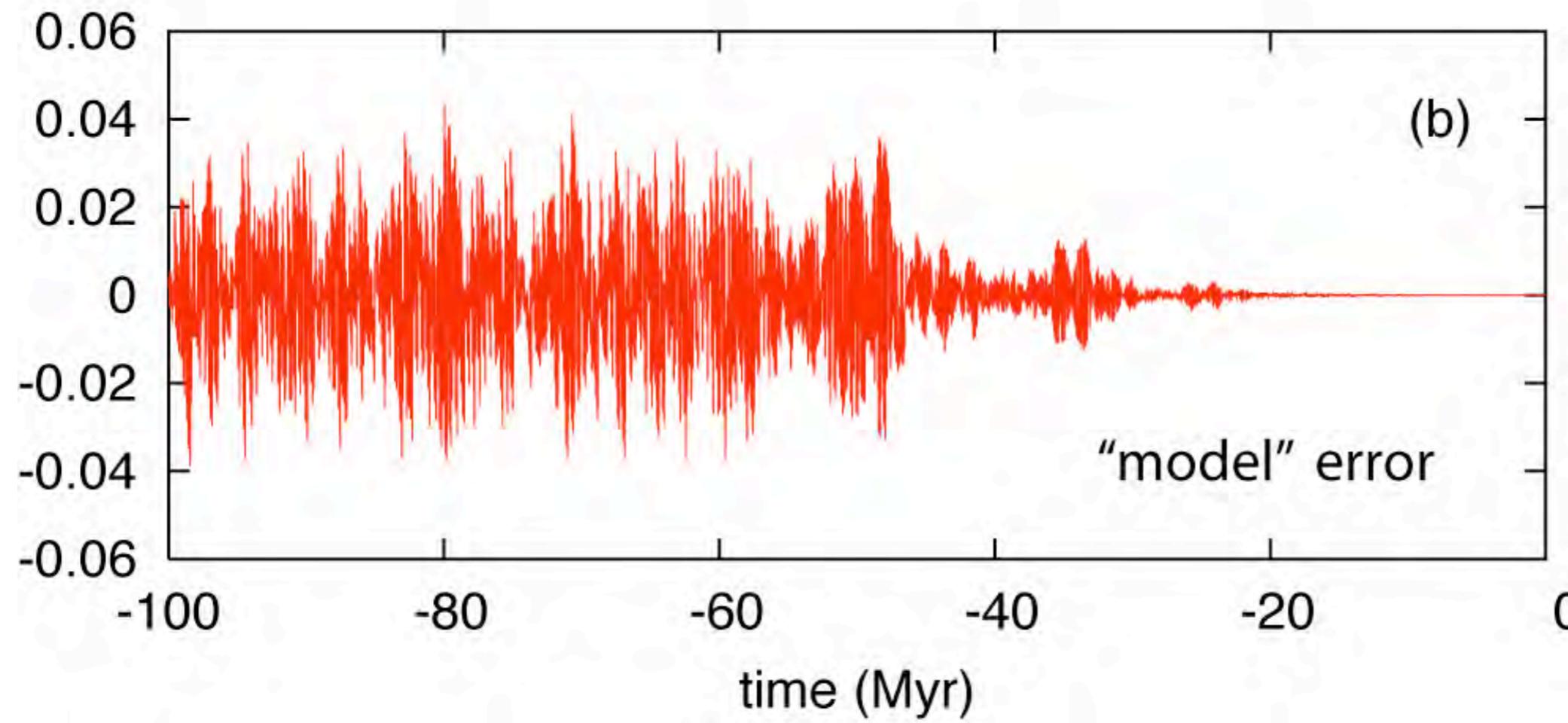
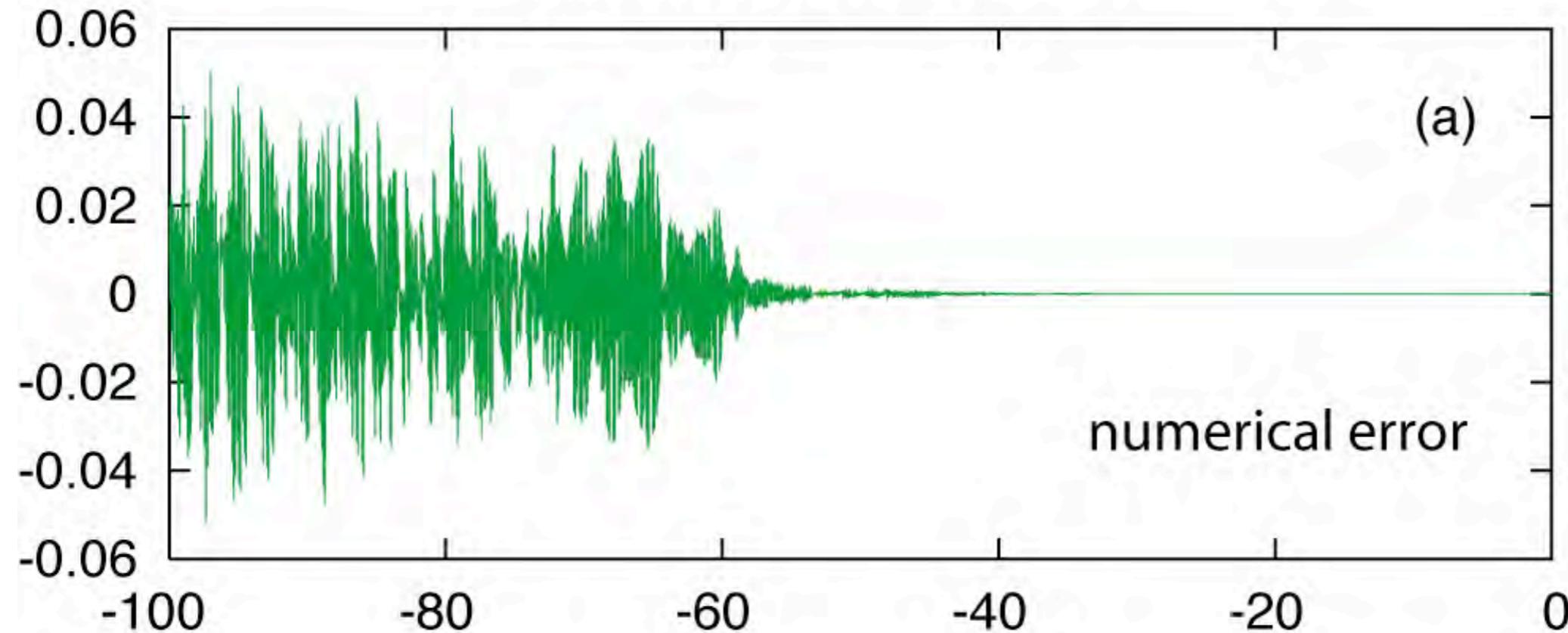
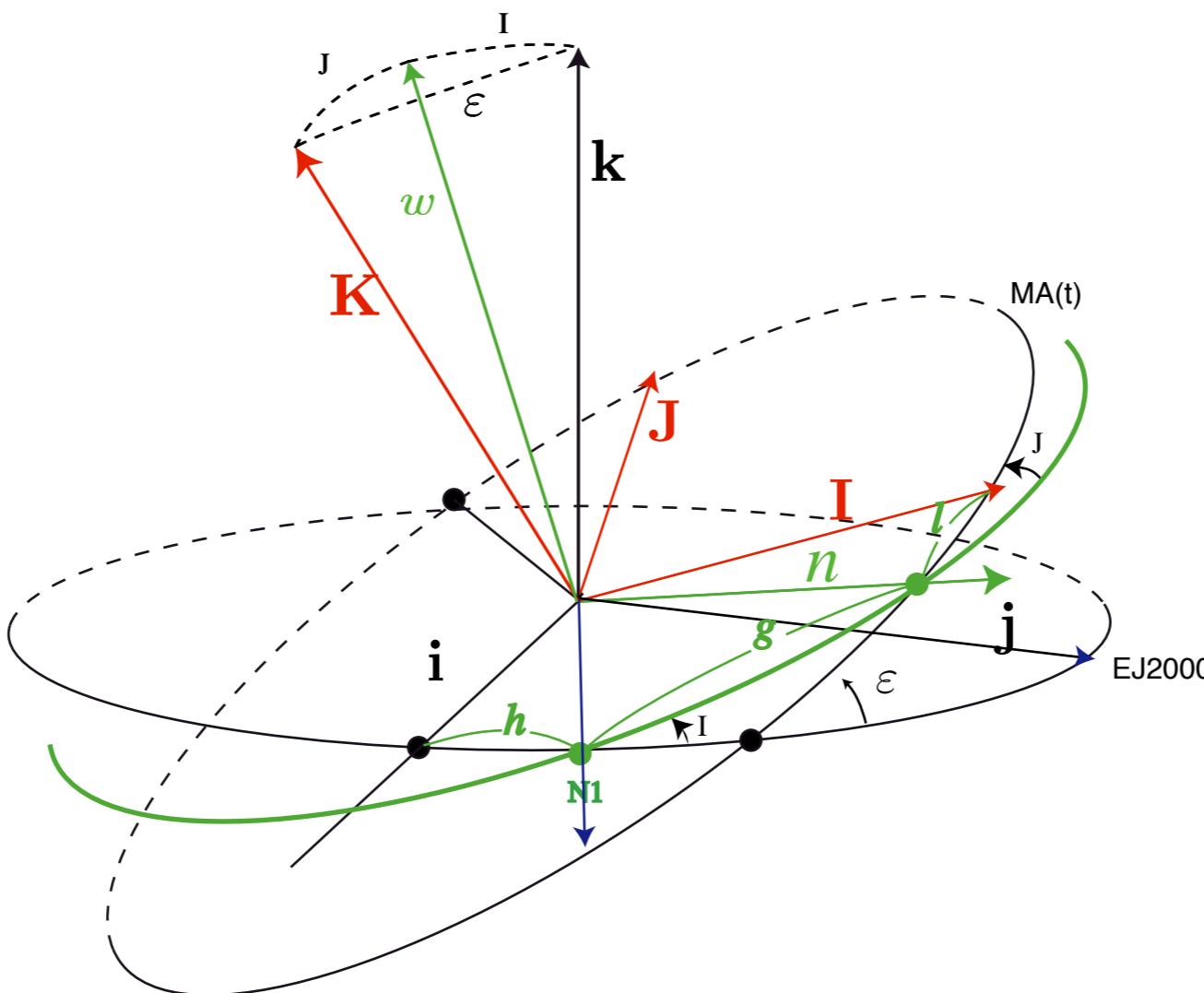


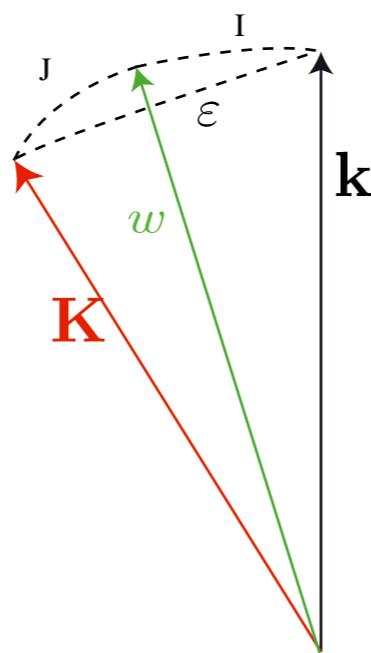
Table 2. Maximum difference between La2004 and DE406 over the whole time interval of DE406 (−5000 yr to +1000 yr with origin at J2000); Col. 1: −100 to +100 yr; Col. 2: −1000 to +1000 yr; Col. 3: −5000 to +1000 yr. EMB is the Earth–Moon barycenter.

	a (UA $\times 10^{10}$)	(15 m)
Mercury	4	43
Venus	13	23
EMB	29	62
Mars	27	76
Jupiter	268	470
Saturn	743	1073
Uranus	1608	3379
Neptune	3315	5786
Pluto	4251	10 603
Moon	23	204
Earth	487	3918
		10 238

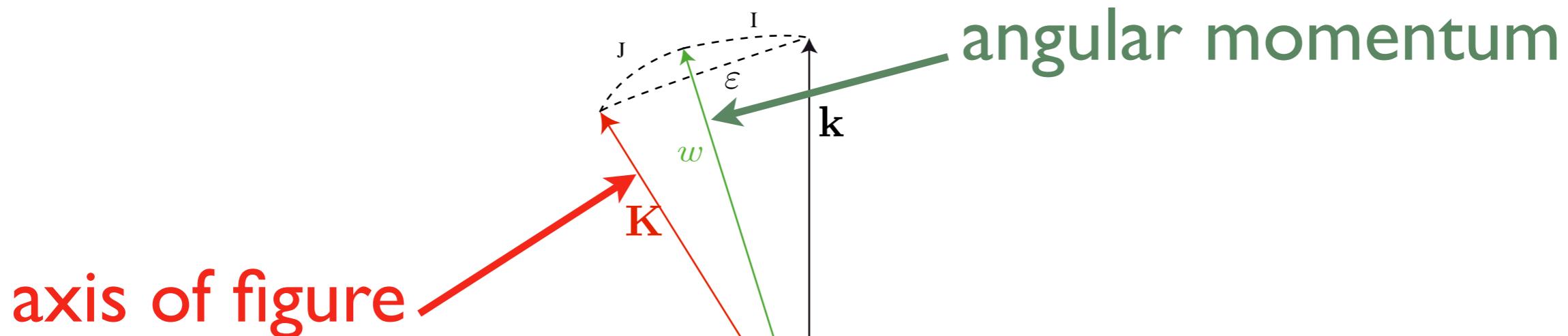
Remarks on Precession Equations



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Remarks on Precession Equations

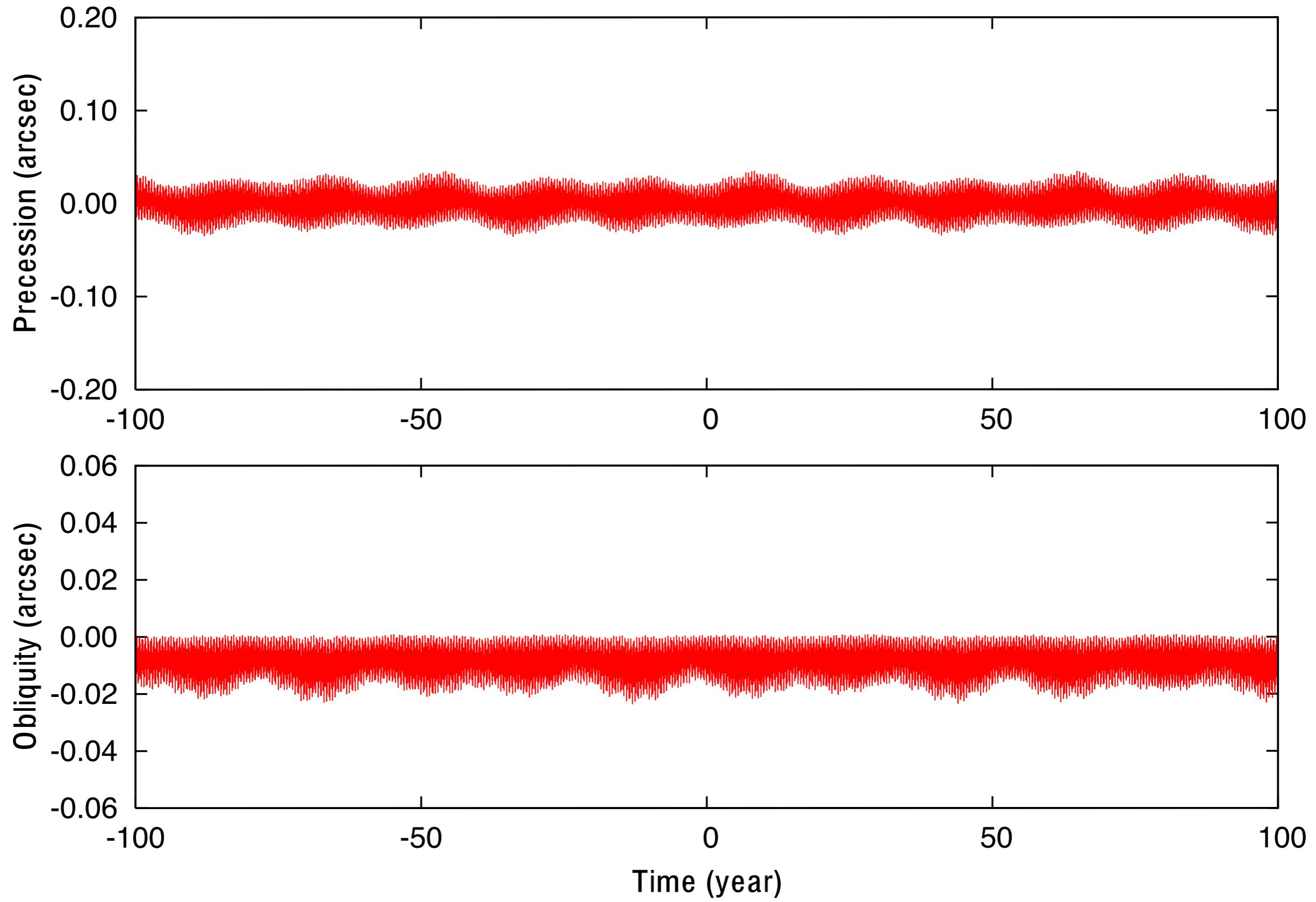


$$\dot{K} = \omega \frac{C}{A} w \wedge K$$

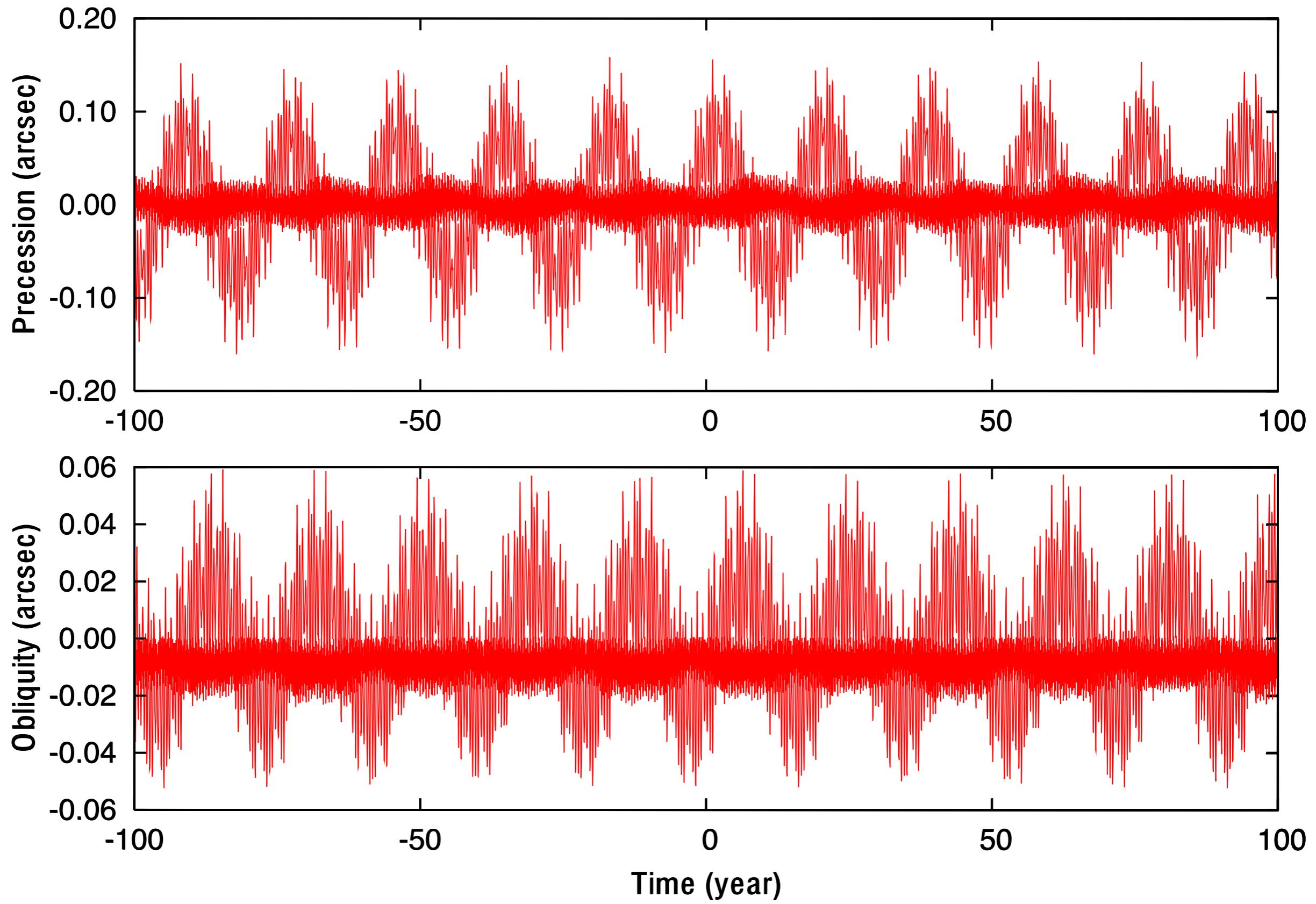
nearly diurnal motion around w

$$\langle K \rangle = \cos J \ w = w + O(J^2) \quad J < 0.2''$$

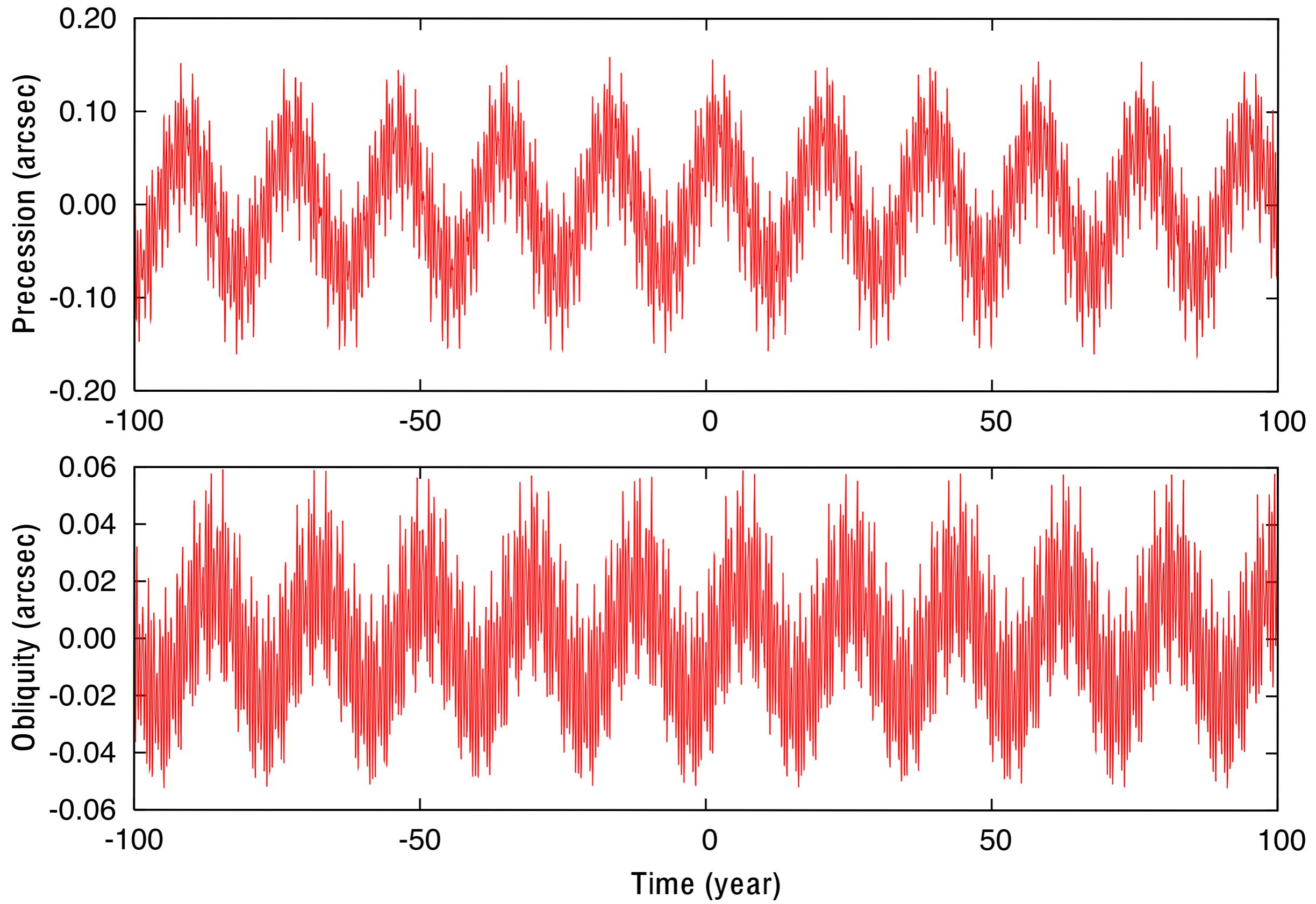
INPOP K - w



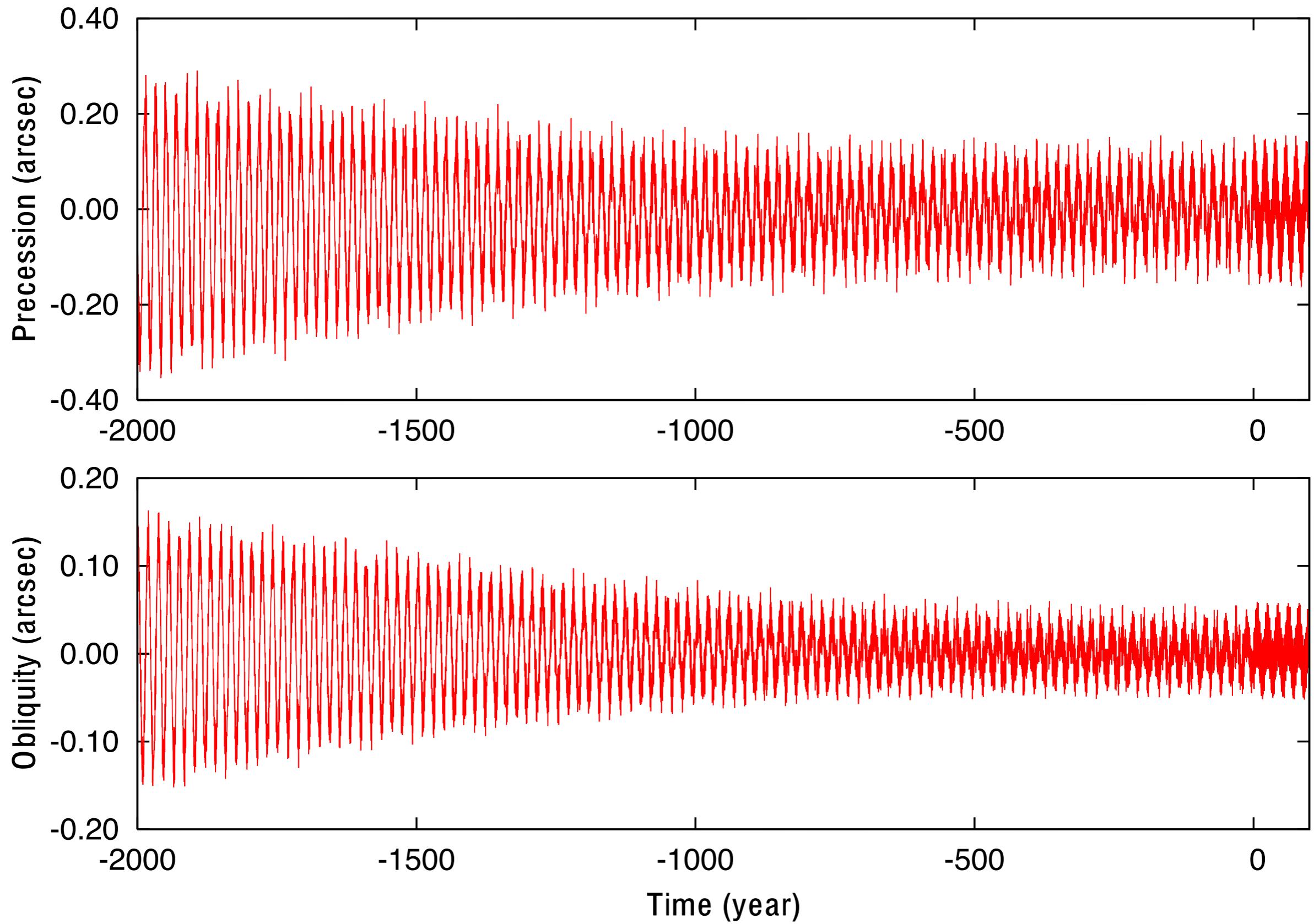
CIP (P03) - INPOP (w)



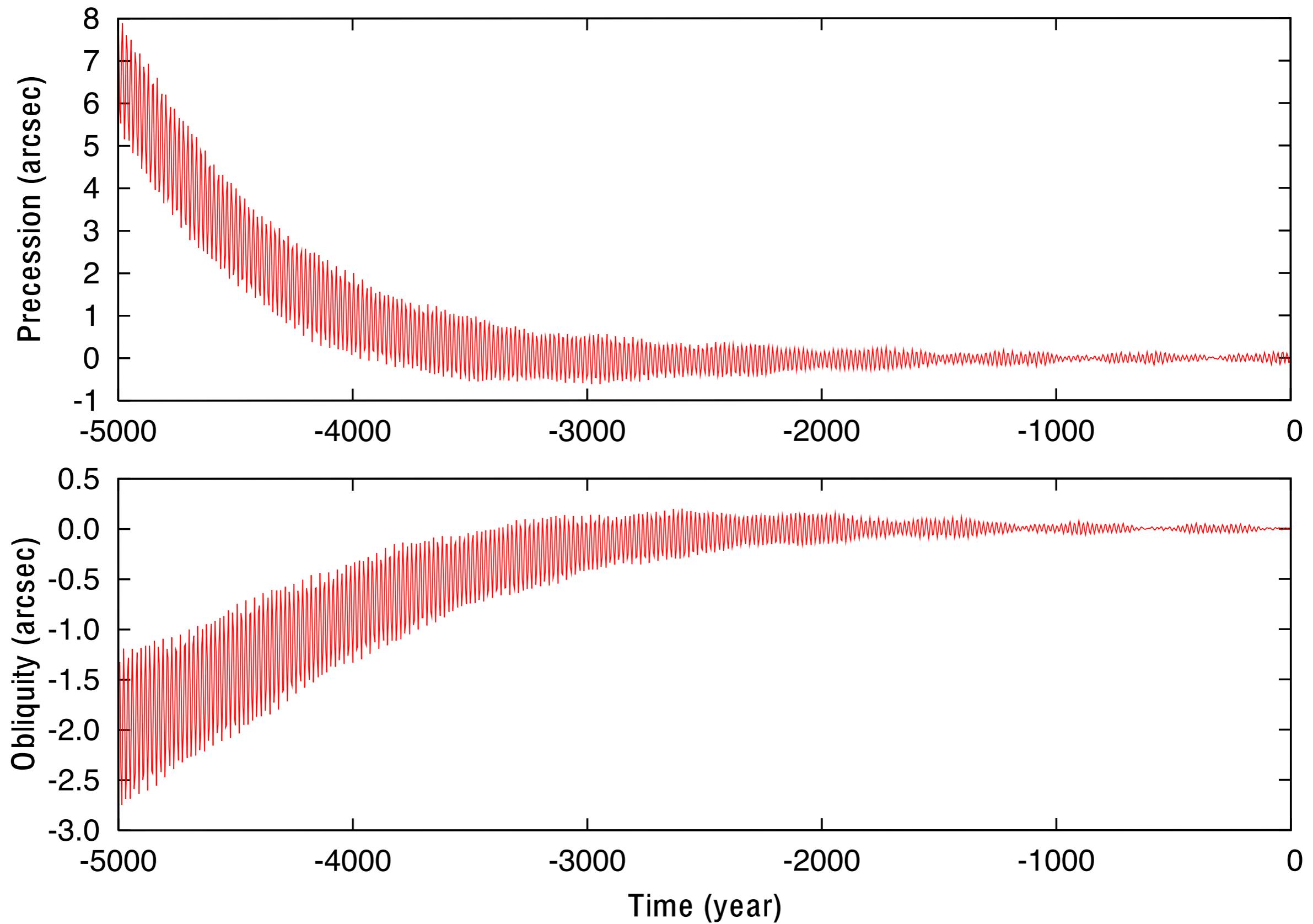
CIP (P03) - INPOP (w)



CIP (P03) - INPOP06 (same \dot{J}_2)



CIP (P03) - INPOP06 (same \dot{J}_2)



Integration of the angular momentum

- Averaging over the diurnal motion
(Boué and Laskar, 2006)
- Takes into account the tidal deformation of the Earth, or post glacial deformation
- Do not depend on the internal model of the Earth
- The precession motions of K and w are the same

Suggestion : Any rotational solution for the Earth should provide also its angular momentum solution !