

Long term solutions for the insolation quantities of the Earth

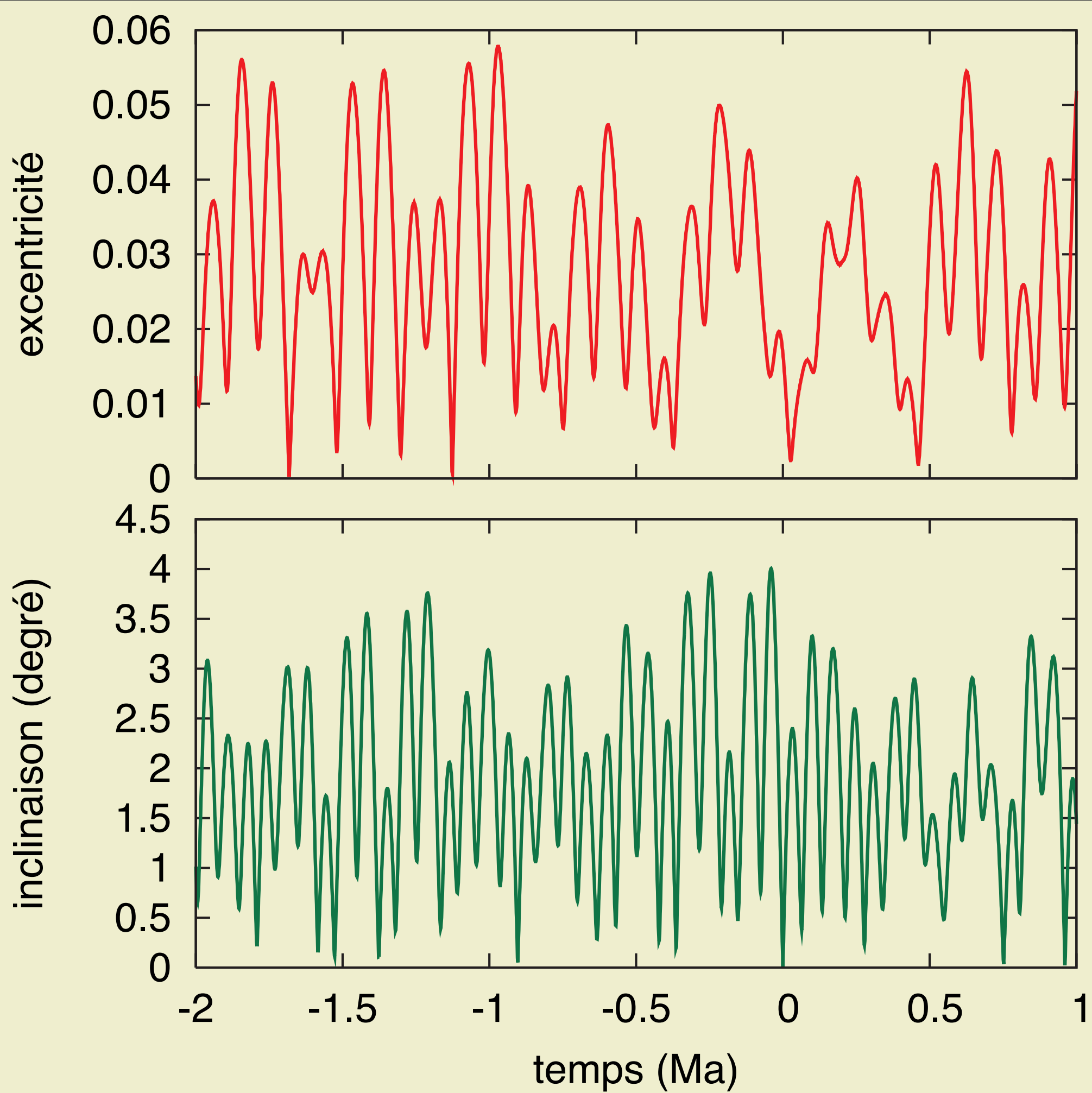
Jacques Laskar

IMCCE/CNRS, Observatoire de Paris

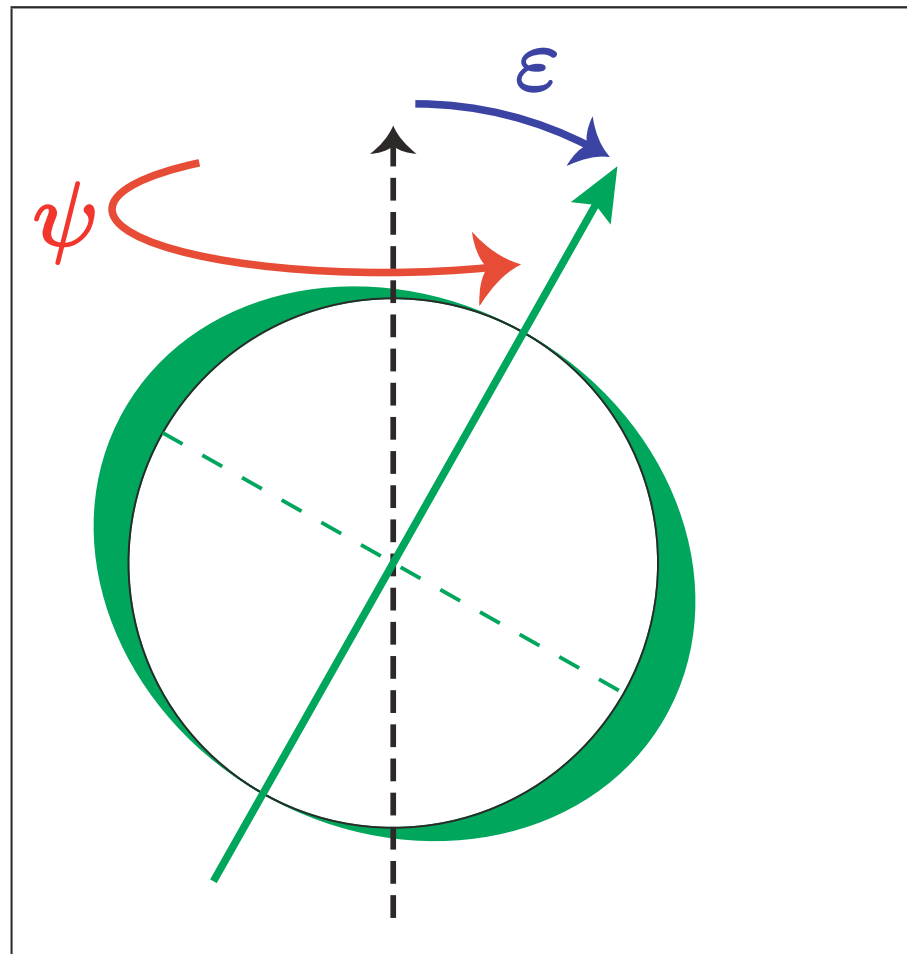
A.C.M Correia, A. Fienga, M. Gastineau, F. Joutel,
B. Levrard, H. Manche, P. Robutel

Astronomical theory of paleoclimates

- Adh mar, 1845
- Croll, 1890
- Pilgrim, 1904
- Milankovitch, 1920, 1941
- Hays, Imbrie, Shackleton, 1976



Precession (ψ) and obliquity ($X = \cos \varepsilon$)



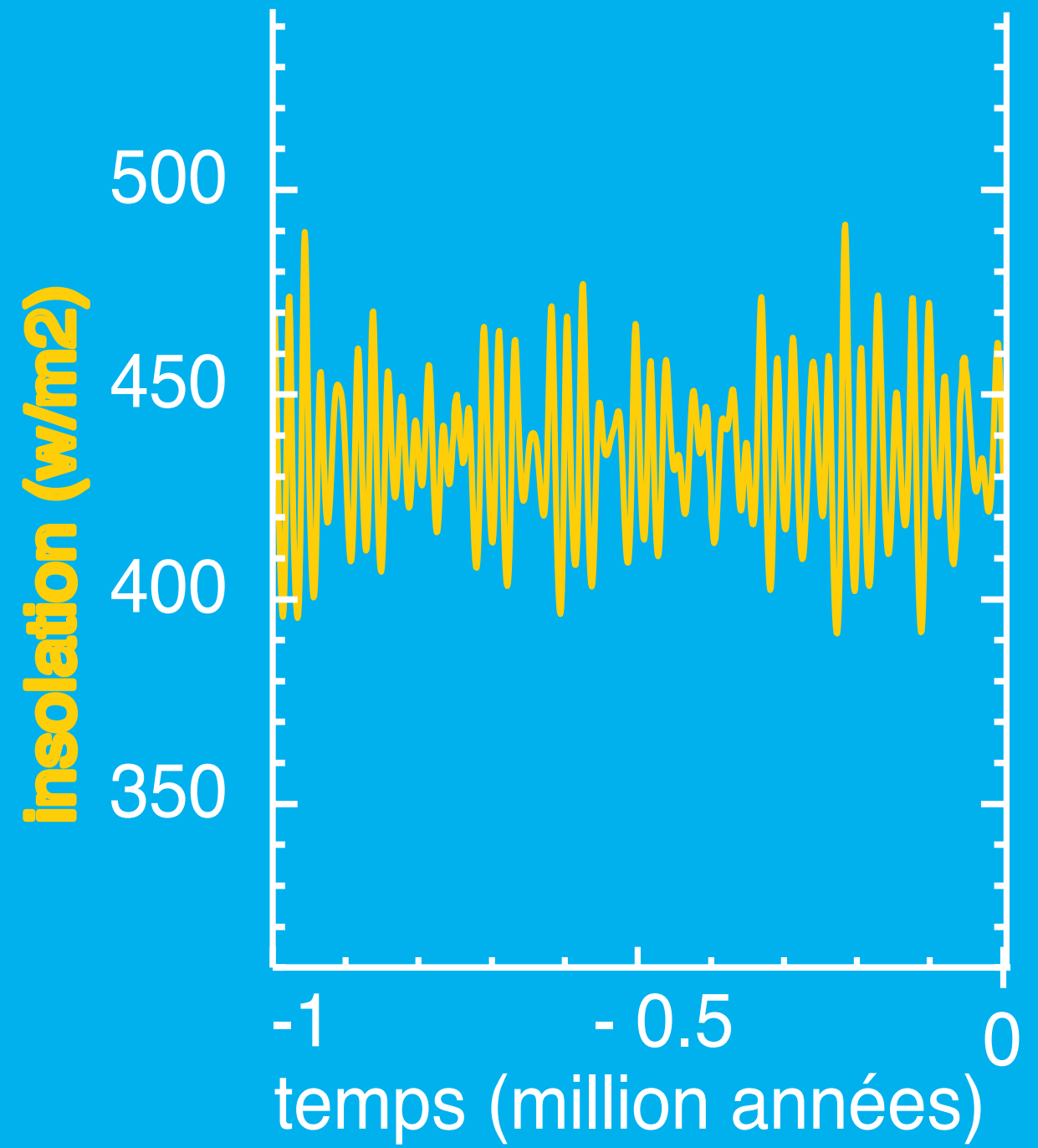
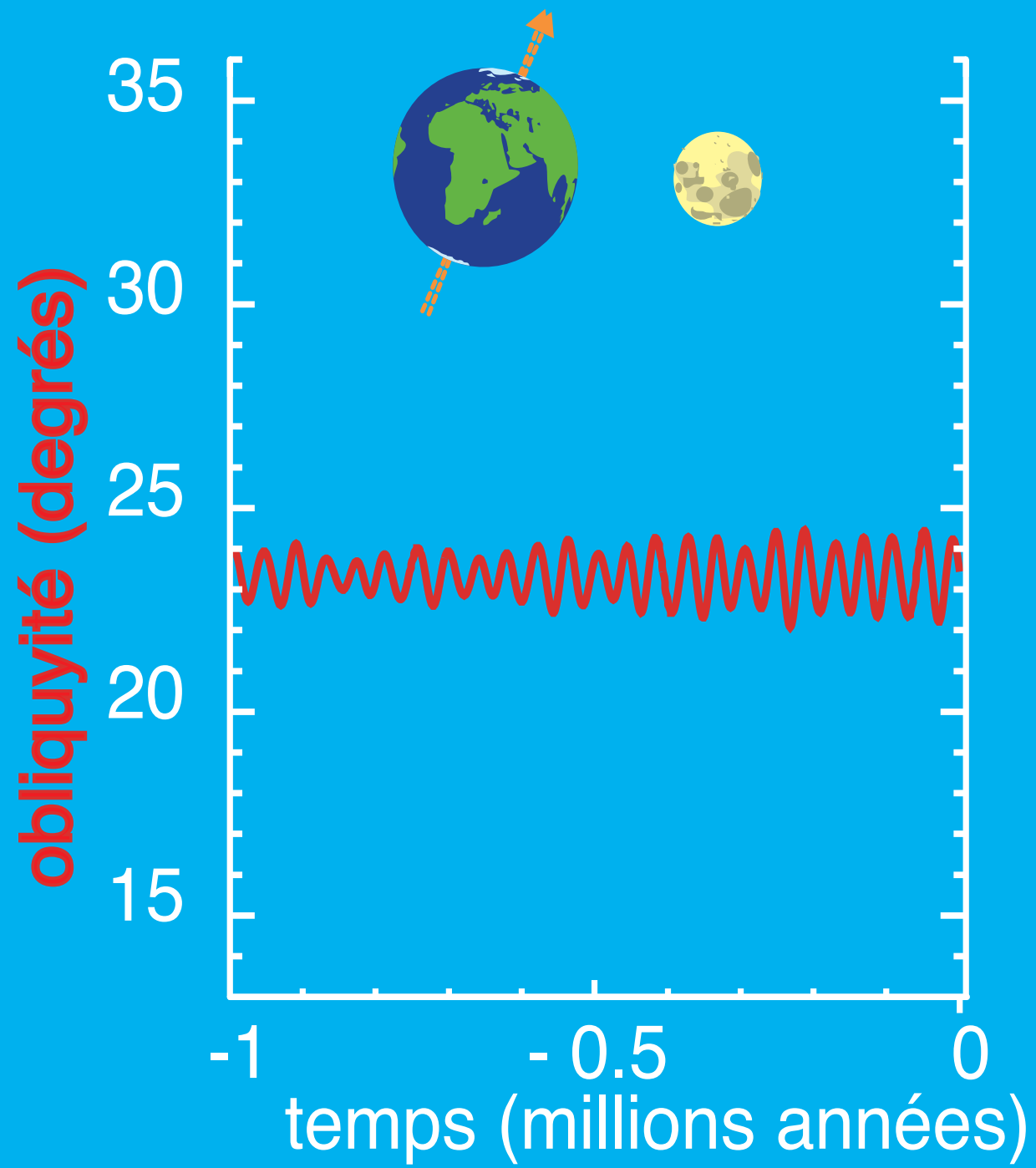
$$H = \frac{1}{2} \alpha X^2$$

$$\begin{cases} \frac{d\psi}{dt} = \frac{\partial H}{\partial X} = \alpha X_0 \\ \frac{dX}{dt} = -\frac{\partial H}{\partial \psi} = 0 \end{cases}$$

$$\alpha = \frac{3k^2 C - A}{2\nu C} \left[\frac{m_M}{a_M^3} + \frac{m_\odot}{a_\odot^3} \right]$$

Planetary perturbations

$$H(\psi, X) = \frac{1}{2} \alpha X^2 + \sqrt{1 - X^2} \sum_k \alpha_k \sin(\nu_k t + \phi_k + \psi)$$



LEG
1
7
7

SITE
1
0
9
0

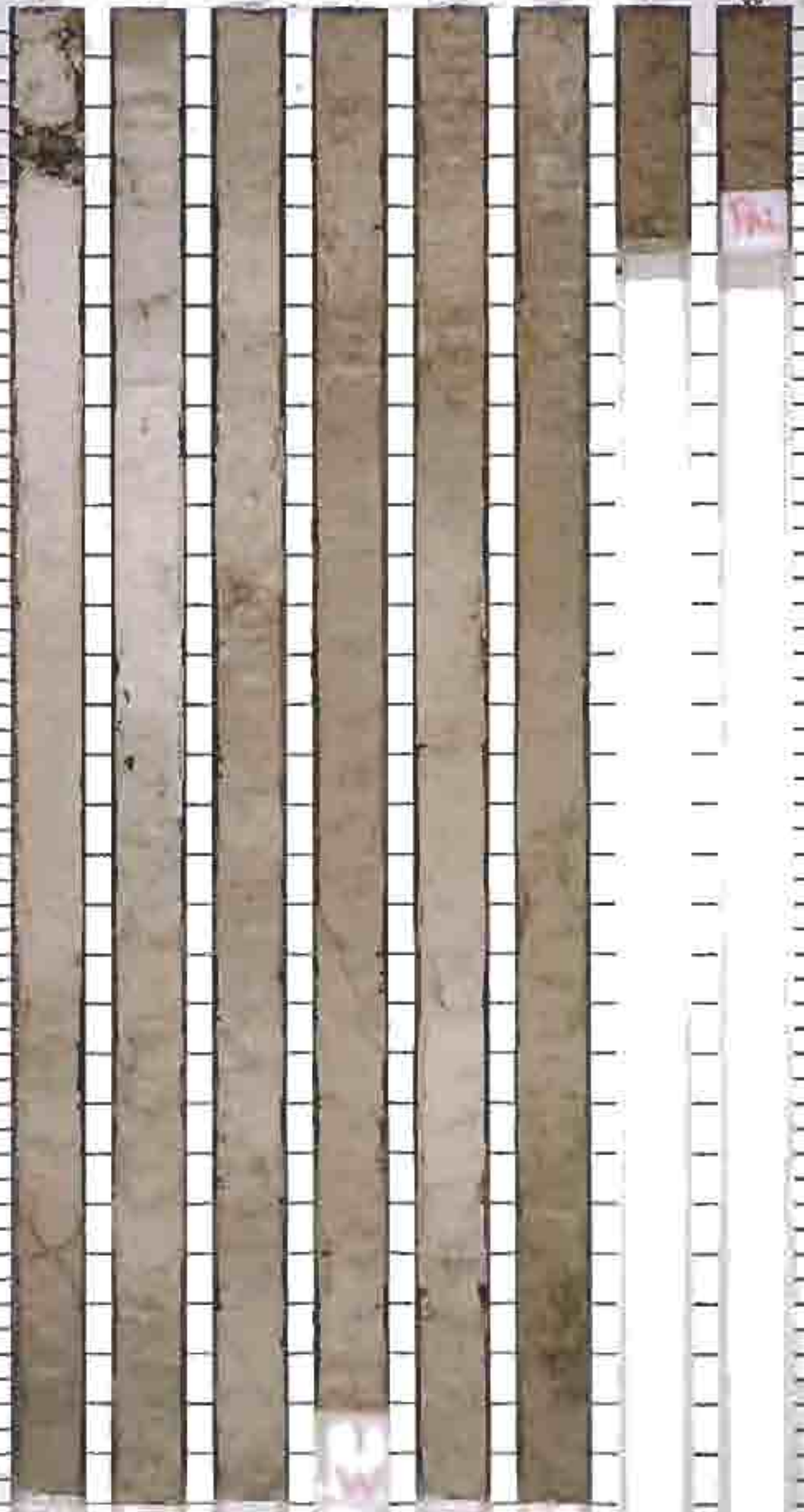


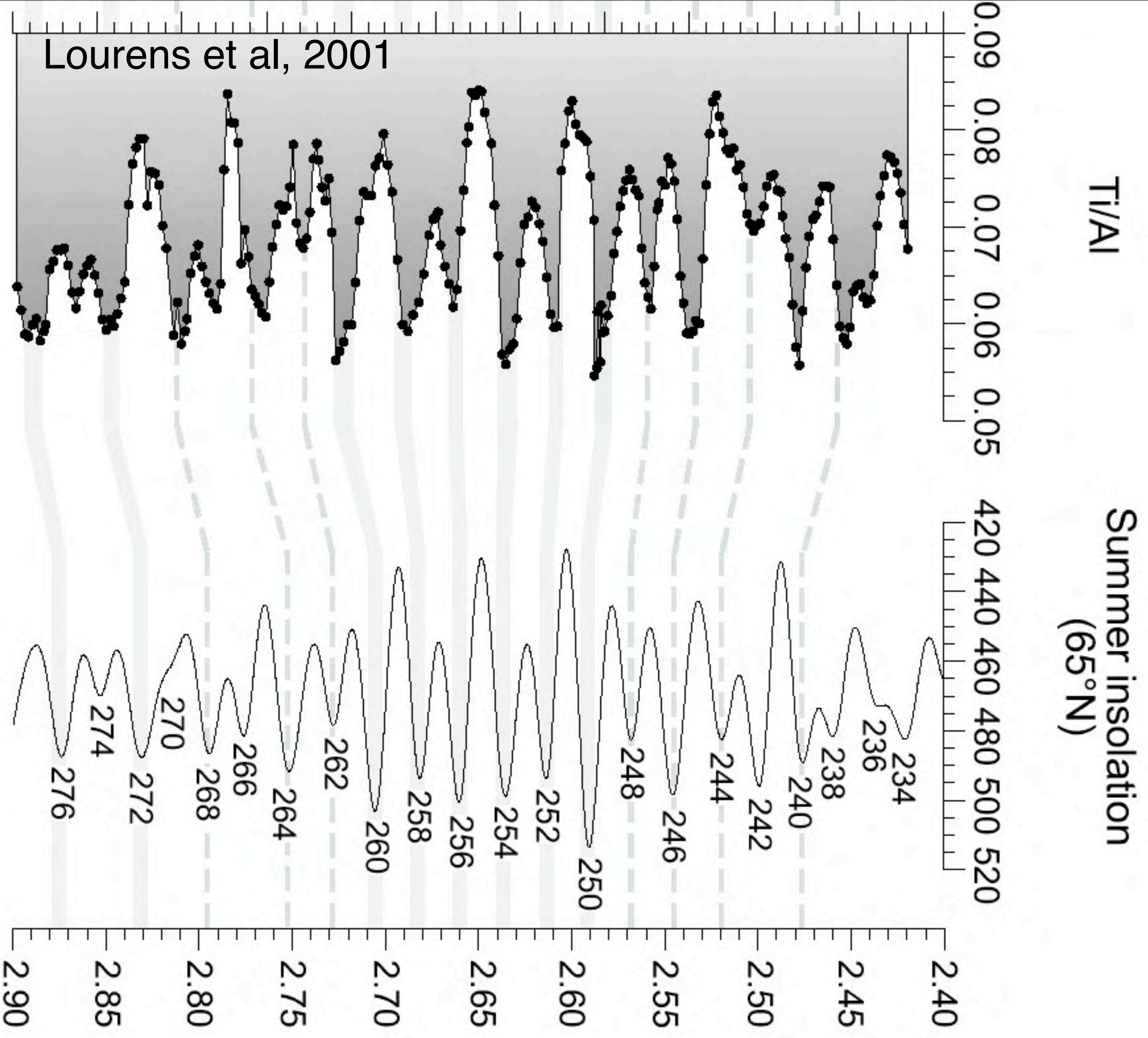
HOLE

B

CORE
2
1
X

5
10
15
20
25
30
35
40
45
50
55
60
65
70
75
80
85
90
95
100
105
110
115
120
125
130
135
140
145
150







INTERNATIONAL STRATIGRAPHIC CHART

International Commission on Stratigraphy



Eonothem Eon	Erathem Era	System Period	Series Epoch	Stage Age	Age Ma	GSSP	
Phanerozoic	Cenozoic	Neogene	Holocene				
			Pleistocene	Upper	0.0115		
				Middle	0.126		
				Lower	0.781		
			Pliocene	Gelasian	1.806		
				Piacenzian	2.588		
				Zanclean	3.600		
			Miocene	Messinian	5.332		
				Tortonian	7.246		
				Serravallian	11.608		
				Langhian	13.65		
				Burdigalian	15.97		
				Aquitanian	20.43		
					23.03		
	Paleogene	Oligocene	Chattian	28.4 ±0.1			
			Rupelian	33.9 ±0.1			
		Eocene	Priabonian	37.2 ±0.1			
			Bartonian	40.4 ±0.2			
			Lutetian	48.6 ±0.2			
			Ypresian	55.8 ±0.2			
			Paleocene	Thanetian	58.7 ±0.2		
		Selandian		61.7 ±0.2			
		Danian		65.5 ±0.3			
		Mesozoic	Cretaceous	Upper	Maastrichtian	70.6 ±0.6	
					Campanian	83.5 ±0.7	
					Santonian	85.8 ±0.7	
					Coniacian	89.3 ±1.0	
					Turonian	93.5 ±0.8	
Cenomanian	99.6 ±0.9						
Lower	Albian				112.0 ±1.0		
	Aptian			125.0 ±1.0			
	Barremian			130.0 ±1.5			
	Hauterivian			136.4 ±2.0			
	Valanginian			140.2 ±3.0			
	Berriasian			145.5 ±4.0			

Eonothem Eon	Erathem Era	System Period	Series Epoch	Stage Age	Age Ma	GSSP
Phanerozoic	Mesozoic	Jurassic	Upper	Tithonian	145.5 ±4.0	
				Kimmeridgian	150.8 ±4.0	
				Oxfordian	155.7 ±4.0	
			Middle	Callovian	161.2 ±4.0	
				Bathonian	164.7 ±4.0	
				Bajocian	167.7 ±3.5	
		Lower	Aalenian		171.6 ±3.0	
			Toarcian		175.6 ±2.0	
			Pliensbachian		183.0 ±1.5	
			Sinemurian		189.6 ±1.5	
	Triassic	Upper	Hettangian		196.5 ±1.0	
			Rhaetian		199.6 ±0.6	
			Norian		203.6 ±1.5	
		Middle	Carnian		216.5 ±2.0	
			Ladinian		228.0 ±2.0	
		Lower	Anisian		237.0 ±2.0	
			Olenekian		245.0 ±1.5	
			Induan		249.7 ±0.7	
	Paleozoic	Permian	Lopingian	Changhsingian	251.0 ±0.4	
				Wuchiapingian	253.8 ±0.7	
			Guadalupian	Capitanian	260.4 ±0.7	
				Wordian	265.8 ±0.7	
		Carboniferous	Roadian		268.0 ±0.7	
			Kungurian		270.6 ±0.7	
			Artinskian		275.6 ±0.7	
			Sakmarian		284.4 ±0.7	
			Asselian		294.6 ±0.8	
			Gzhelian		299.0 ±0.8	
		Pennsylvanian	Kasimovian		303.9 ±0.9	
			Moscovian		306.5 ±1.0	
			Bashkirian		311.7 ±1.1	
			Serpukhovian		318.1 ±1.3	
	Mississippian	Upper	Visean		326.4 ±1.6	
			Tournaisian		345.3 ±2.1	
		Lower			359.2 ±2.5	

Eonothem Eon	Erathem Era	System Period	Series Epoch	Stage Age	Age Ma	GSSP
Phanerozoic	Paleozoic	Devonian	Upper	Famennian	359.2 ±2.5	
				Frasnian	374.5 ±2.6	
				Givetian	385.3 ±2.6	
		Middle	Eifelian		391.8 ±2.7	
			Emsian		397.5 ±2.7	
			Pragian		407.0 ±2.8	
		Lower	Lochkovian		411.2 ±2.8	
			Pridoli		416.0 ±2.8	
			Ludlow	Ludfordian	418.7 ±2.7	
				Gorstian	421.3 ±2.6	
	Silurian	Wenlock	Homerian		422.9 ±2.5	
			Sheinwoodian		426.2 ±2.4	
			Telychian		428.2 ±2.3	
		Llandovery	Aeronian		436.0 ±1.9	
			Rhuddanian		439.0 ±1.8	
		Ordovician	Hirnantian		443.7 ±1.5	
					445.6 ±1.5	
					455.8 ±1.6	
		Cambrian	Darriwilian		460.9 ±1.6	
					468.1 ±1.6	
					471.8 ±1.6	
			Tremadocian		478.6 ±1.7	
Phanerozoic	Paleozoic	Cambrian			488.3 ±1.7	
			Furongian	Paibian	501.0 ±2.0	
					513.0 ±2.0	
					542.0 ±1.0	

Eonothem Eon	Erathem Era	System Period	Age Ma	GSSP GSSA
Precambrian	Proterozoic	Ediacaran	542	
			~630	
			850	
		Meso-proterozoic	1000	
			1200	
			1400	
		Paleo-proterozoic	1600	
			1800	
			2050	
			2300	
	Archean	Neoarchean	2500	
		Mesoarchean	2800	
		Paleoarchean	3200	
		Lower limit is not defined	3600	

Subdivisions of the global geologic record are formally defined by their lower boundary. Each unit of the Phanerozoic interval (~542 Ma to Present) and the base of the Ediacaran is defined by a Global Standard Section and Point (GSSP) at its base, whereas the Precambrian Interval is formally subdivided by absolute age, Global Standard Stratigraphic Age (GSSA).

This chart gives an overview of the international chronostratigraphic units, their rank, their names and formal status. These units are approved by the International Commission on Stratigraphy (ICS) and ratified by the International Union of Geological Sciences (IUGS).

The Guidelines of ICS (Remane et al., 1996, Episodes, 19: 77-81) regulate the selection and

definition of the international units of geologic time. Many GSSP's actually have a 'golden' spike () and Stage and/or System name plaque mounted at the boundary level in the boundary stratotype section, whereas a GSSA is an abstract age without reference to a specific level in a rock section on Earth. Updated descriptions of each GSSP and GSSA are posted on the ICS website (www.stratigraphy.org).

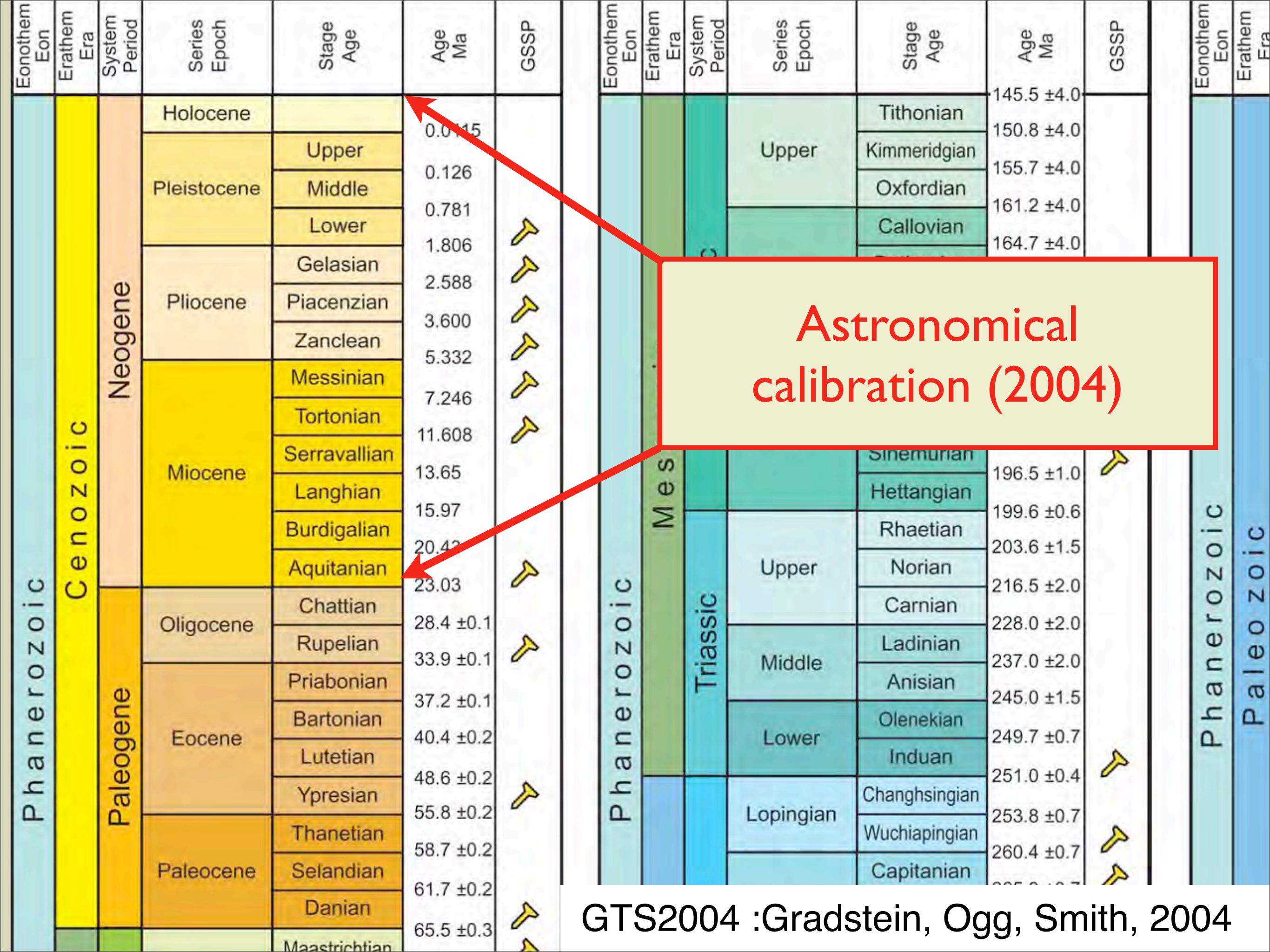
Some stages within the Ordovician and Cambrian will be formally named upon international agreement on their GSSP limits. Most intra-stage boundaries (e.g., Middle and Upper Aptian) are not formally defined. Numerical ages of the unit boundaries in the Phanerozoic are subject to revision. Colors are according to the United States Geological Survey (USGS). The listed numerical ages are from 'A Geologic Time Scale 2004', by F.M. Gradstein, J.G. Ogg, A.G. Smith, et al. (2004) with Cambridge University Press.

This chart was drafted and printed with funding generously provided for the GTS Project 2004 by ExxonMobil, Statoil Norway, ChevronTexaco and BP. The chart was produced by Gabi Ogg.

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GTS2004 :Gradstein, Ogg, Smith, 2004



Astronomical calibration (2004)

GTS2004 :Gradstein, Ogg, Smith, 2004

Astronomical
calibration (in project)

GTS2004 :Gradstein, Ogg, Smith, 2004

Insolation solutions for the Earth

analytical

Orbital Solution

Precession equations

Le Verrier, 1856

Pilgrim, 1904
Milankovitch, 1920, 1941

+Hill, 1897
Brouwer & Van Woerkom, 1950

Sharaf & Budnikova, 1967
Vernekar, 1972

Bretagnon, 1974

Berger, 1978

Laskar, 1988, 1990

Laskar et al, 1993
(-20 Ma -> + 10 Ma)

numerical

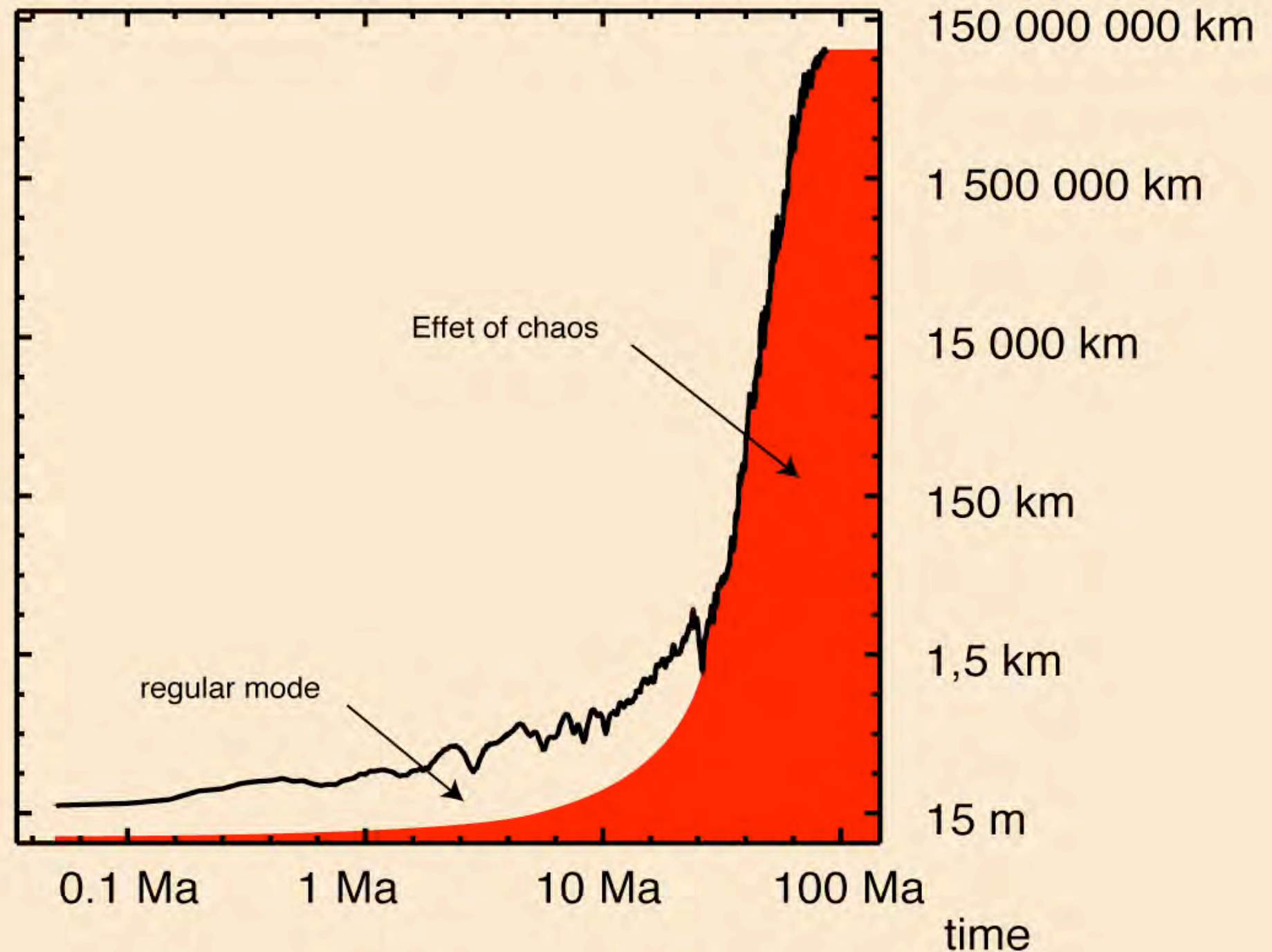
Quinn, Tremaine, Duncan, 1991
(-3Ma -> + 3Ma)

Laskar et al, 2004
(-250 Ma -> + 250 Ma orb: ~ 40 Ma pre : ~ 20 Ma)

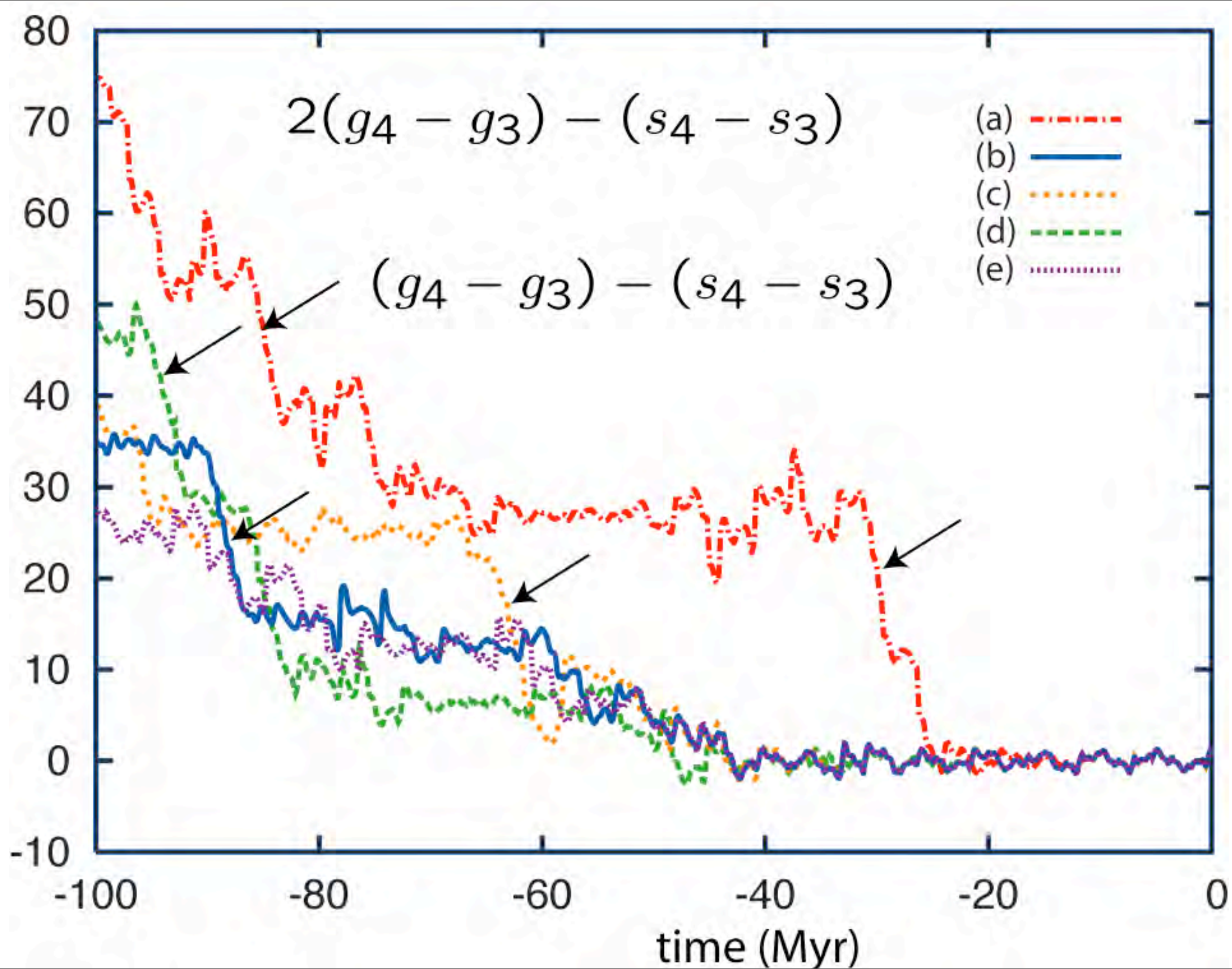
Chaotic motion of the Solar System

Secular equations : 200 Ma : Laskar (1989,1990)

Direct integration : 100 Ma : Sussman and Wisdom (1992)



$$d(T) \approx d_0 10^{T/10}$$



New Challenge :
Orbital solution over ~ 60 Myr



Improve the accuracy
by
2 orders of magnitude !

Development of a New planetary ephemeris

INPOP

(see Agnès Fienga's presentation)

Interactions in La2004, DE405, INPOP

- Newtonian (planets \leftrightarrow planets, planets \leftrightarrow asteroids (300), asteroids \leftrightarrow asteroids (5))
- Relativistic corrections (planets, asteroids)
- Non-spherical body \leftrightarrow point mass
- Sun (J2) \leftrightarrow Planets
- Earth (J2) \leftrightarrow (Moon, Sun, Venus, Jupiter)
- Moon (J2, C21, C22, S21, S22) \leftrightarrow (Earth, Sun, Venus, Jupiter)
- Deformation of extended bodies (tides) \leftrightarrow point mass
- Earth (Sun, Moon) \leftrightarrow (Moon, Sun, Venus, Jupiter)
- Moon (Spin, Earth, Sun) \leftrightarrow (Earth, Sun, Venus, Jupiter)
- Earth Shape \leftrightarrow Moon shape (torque exerted by the Moon)

eccentricity of the Earth (La2004)

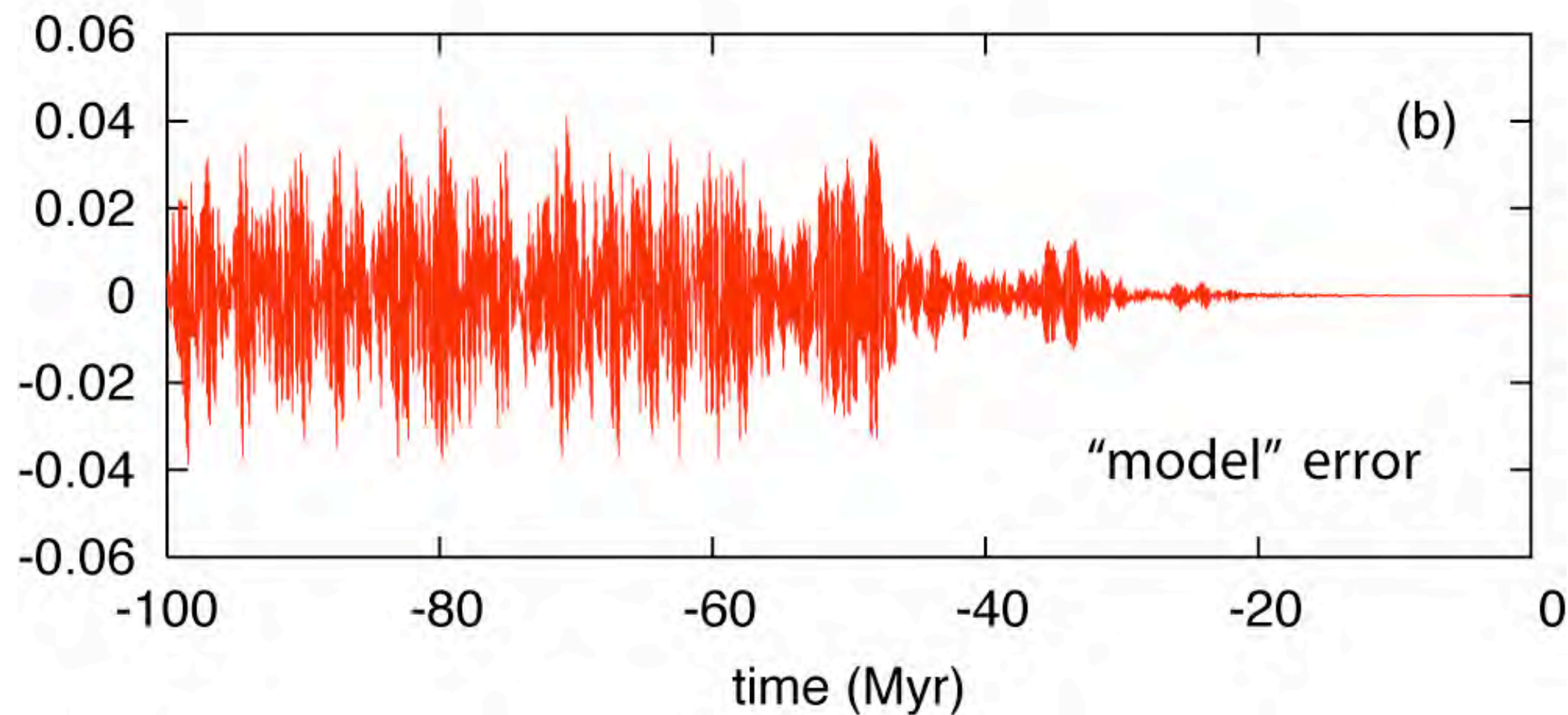
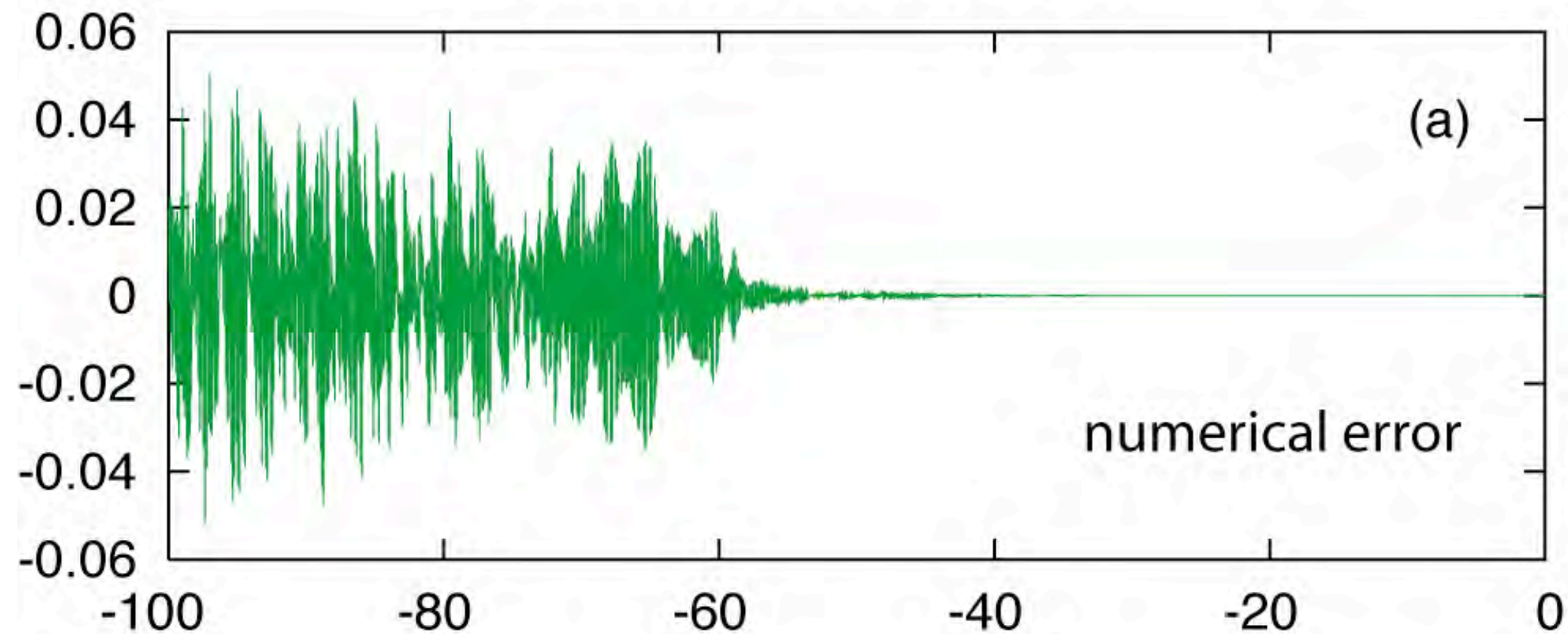
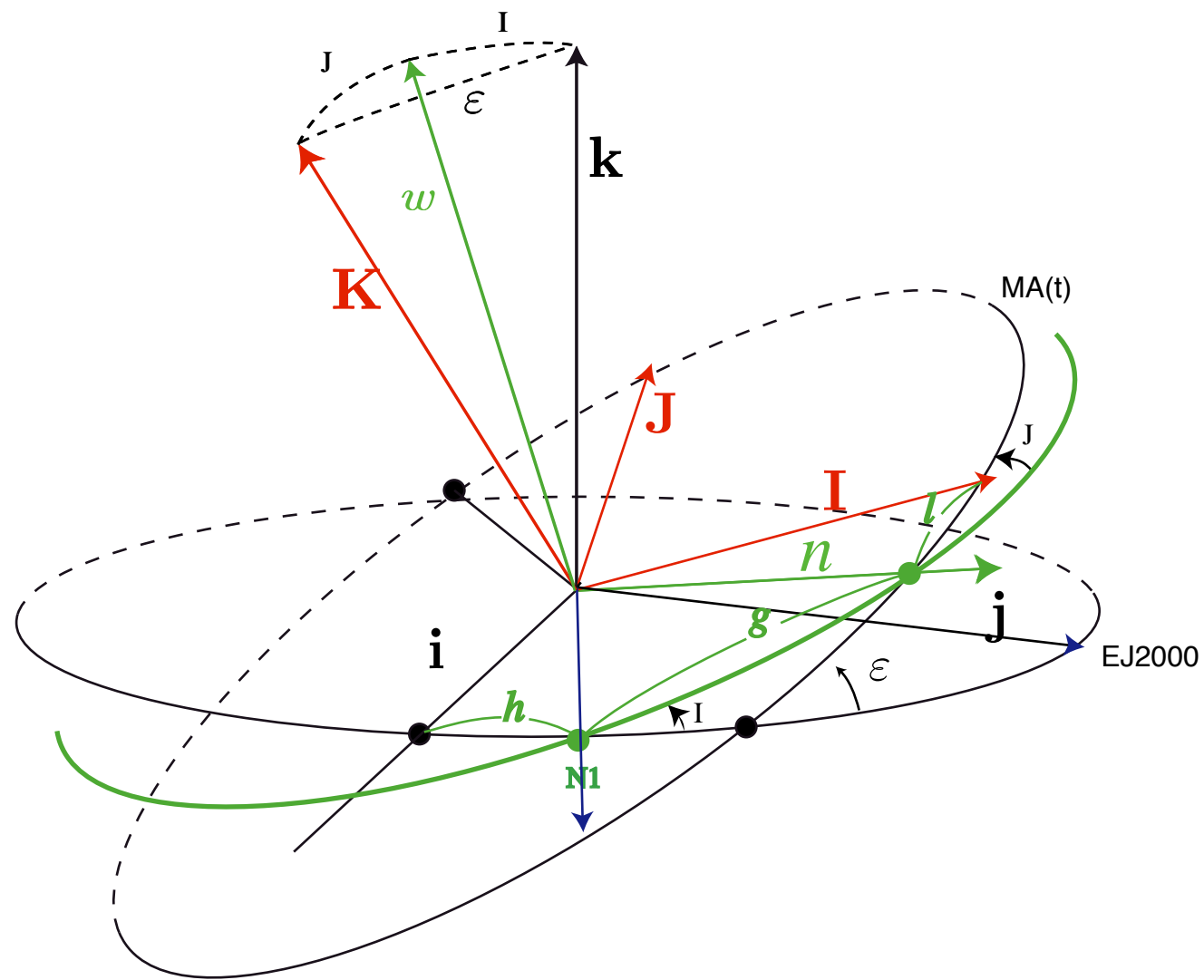


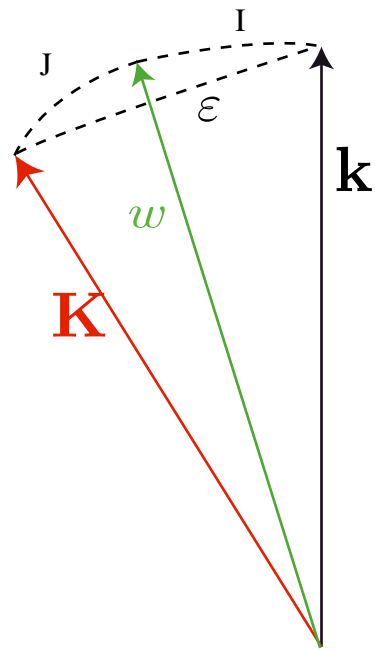
Table 2. Maximum difference between La2004 and DE406 over the whole time interval of DE406 (−5000 yr to +1000 yr with origin at J2000); Col. 1: −100 to +100 yr; Col. 2:−1000 to +1000 yr; Col. 3: −5000 to +1000 yr. EMB is the Earth–Moon barycenter.

		$a \text{ (UA} \times 10^{10} \text{)}$	(15 m)
Mercury	4	43	58
Venus	13	23	27
EMB	29	62	62
Mars	27	76	76
Jupiter	268	470	565
Saturn	743	1073	1168
Uranus	1608	3379	3552
Neptune	3315	5786	6458
Pluto	4251	10 603	10 603
Moon	23	204	520
Earth	487	3918	10 238

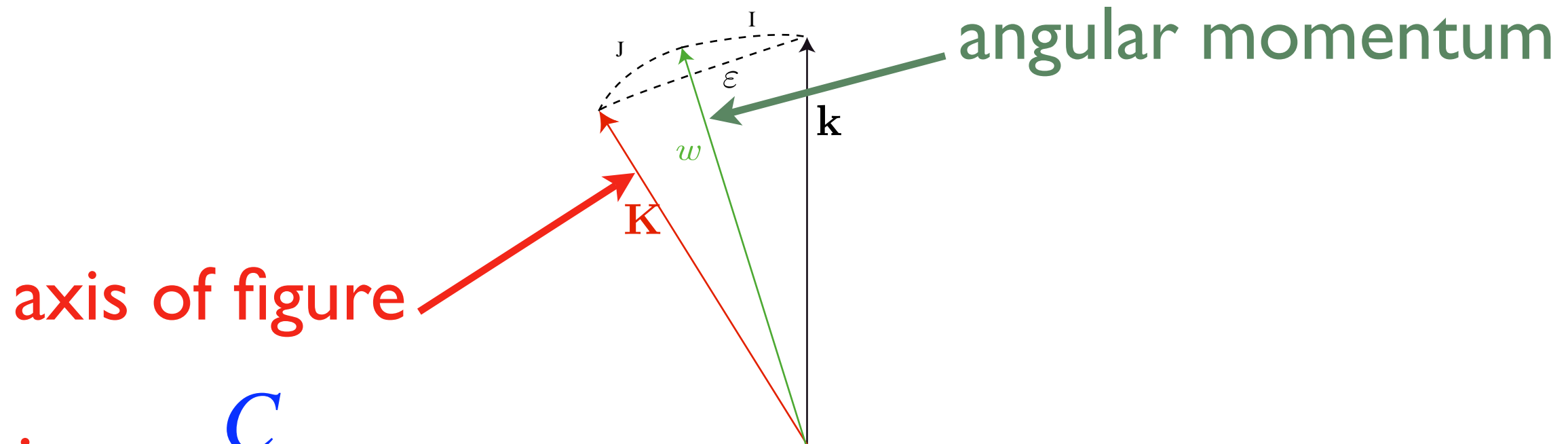
Remarks on Precession Equations



Remarks on Precession Equations



Remarks on Precession Equations

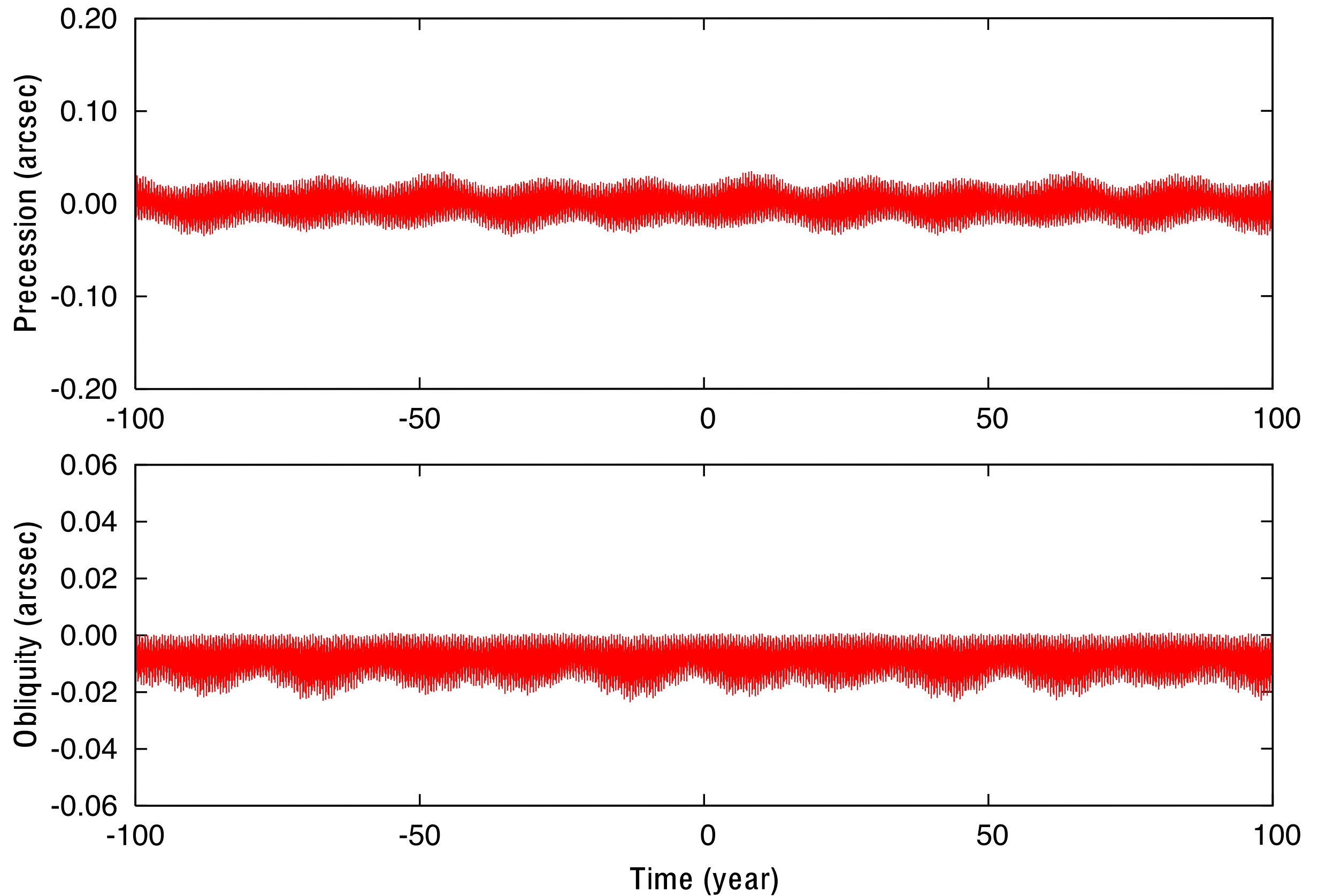


$$\dot{\mathbf{K}} = \omega \frac{C}{A} \mathbf{w} \wedge \mathbf{K}$$

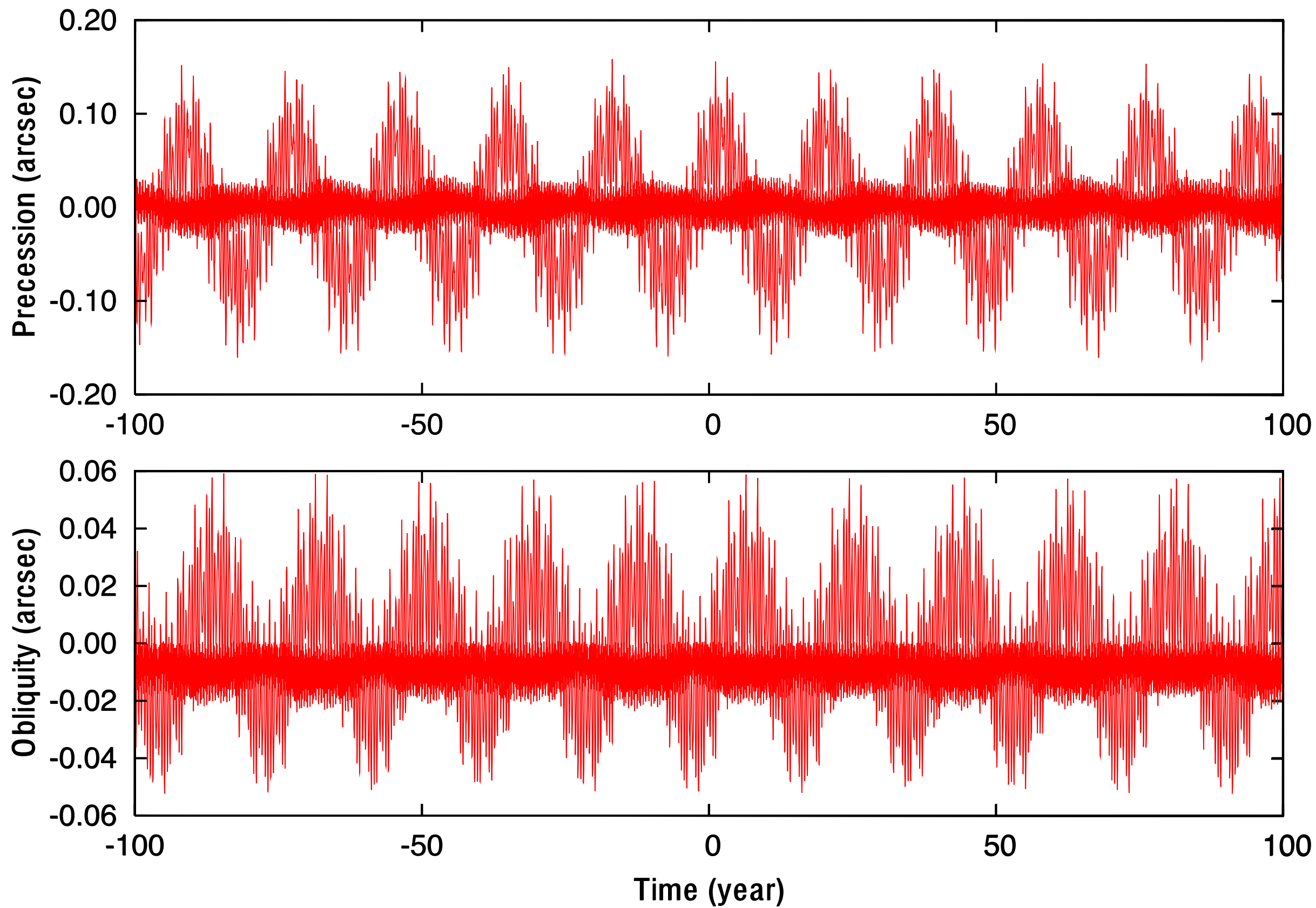
nearly diurnal motion around \mathbf{w}

$$\langle \mathbf{K} \rangle = \cos J \mathbf{w} = \mathbf{w} + O(J^2) \quad J < 0.2''$$

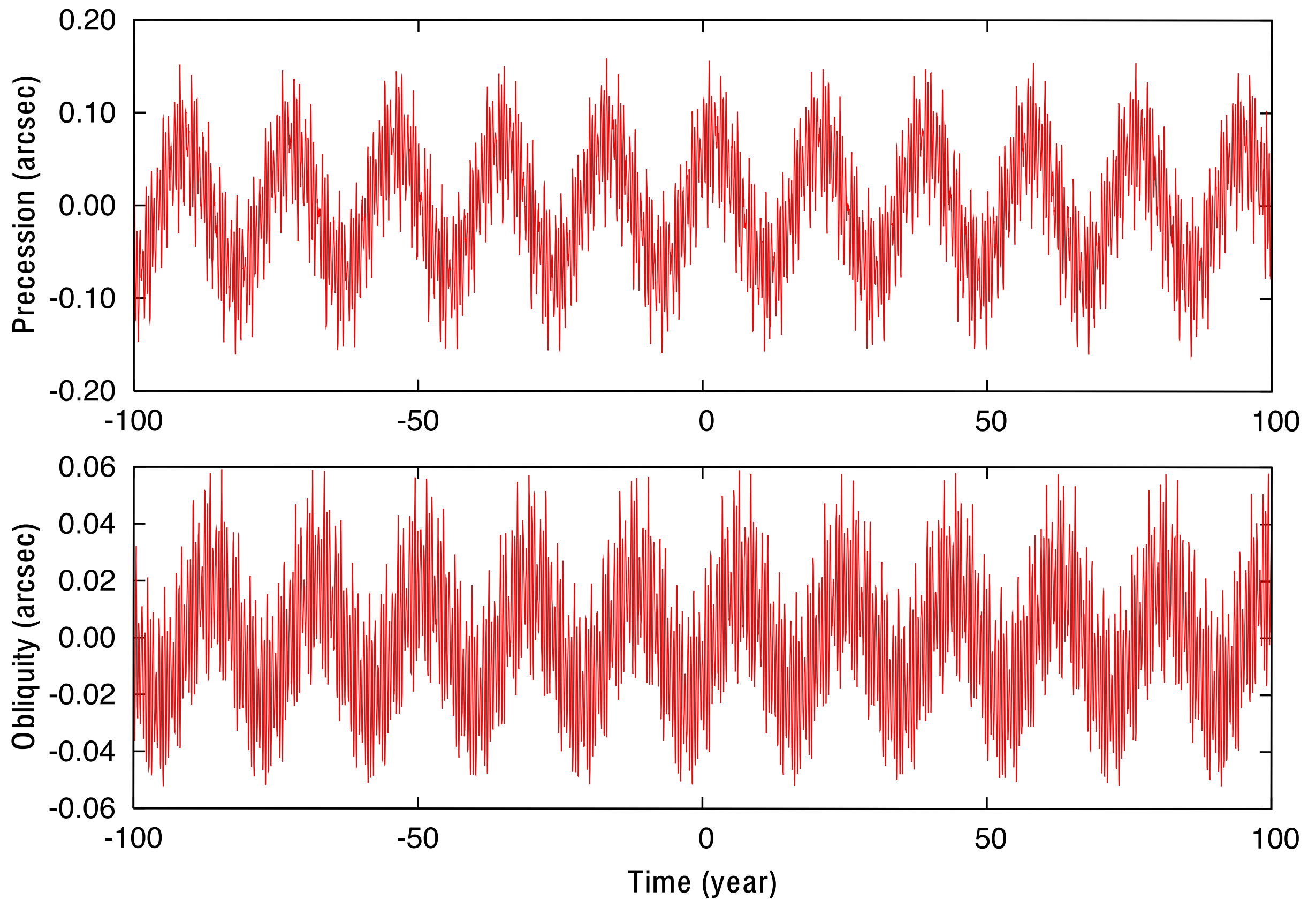
INPOP K - w



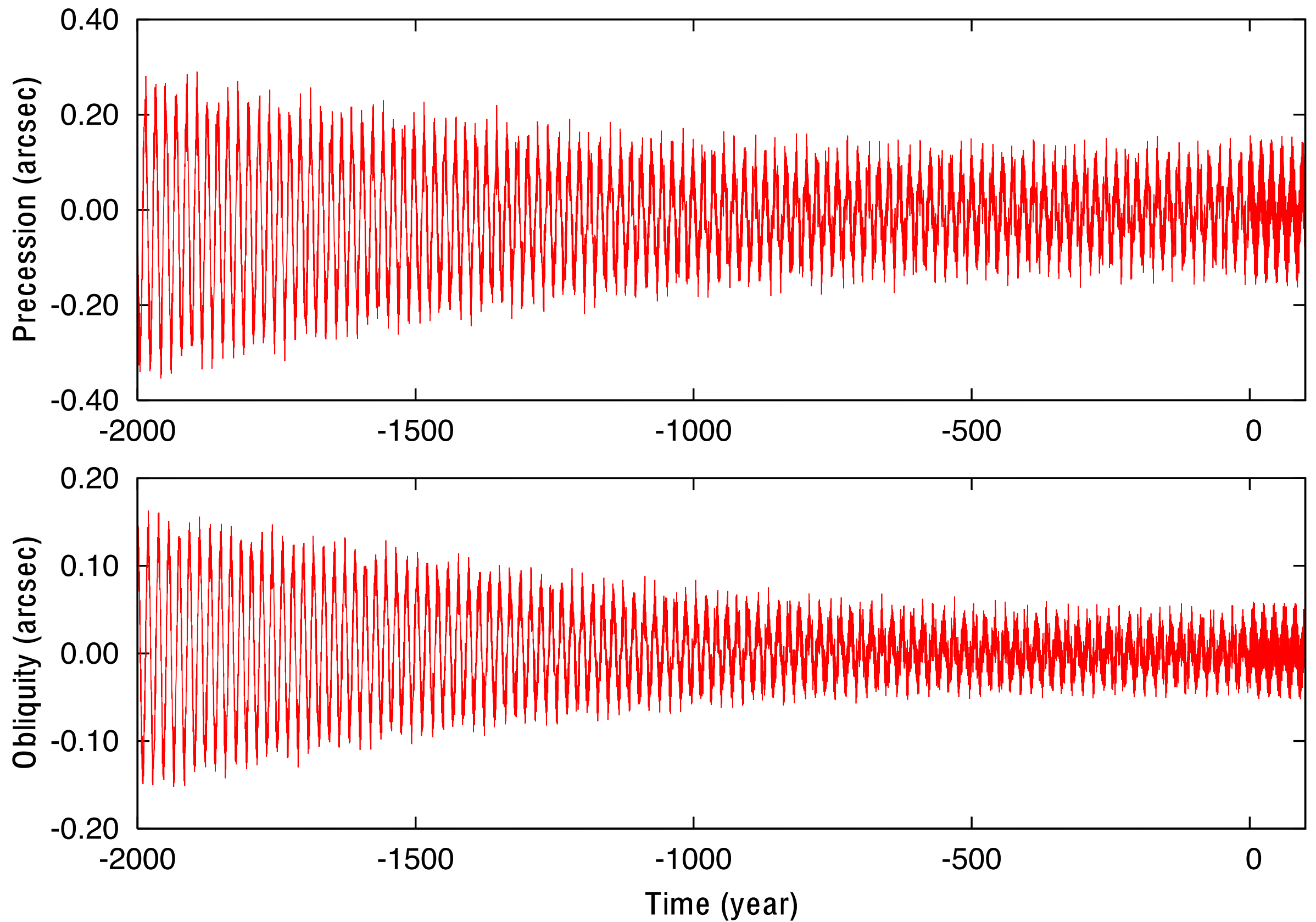
CIP (P03) - INPOP (w)



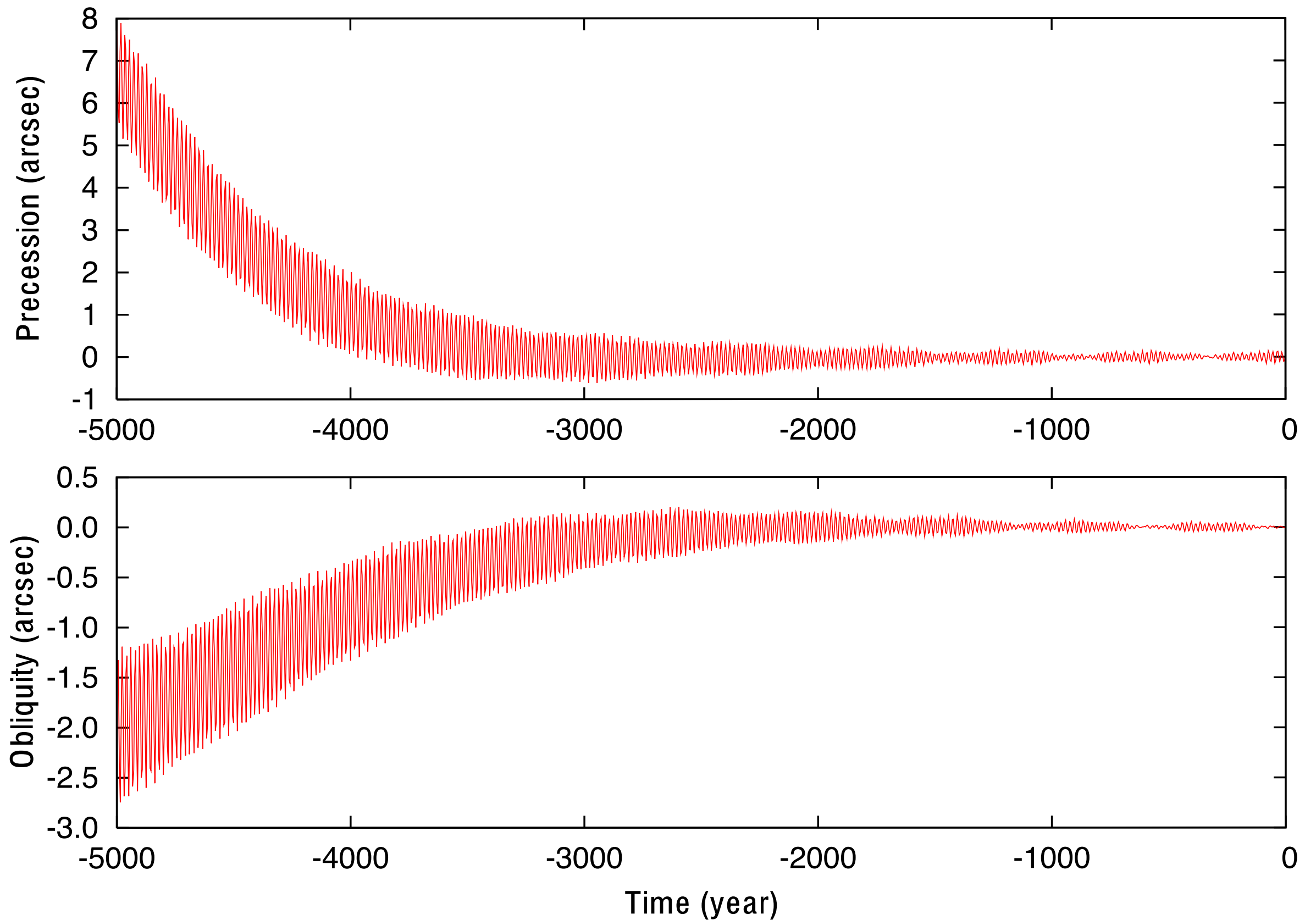
CIP (P03) - INPOP (w)



CIP (P03) - INPOP06 (same \dot{J}_2)



CIP (P03) - INPOP06 (same \dot{J}_2)



Integration of the angular momentum

- Averaging over the diurnal motion
(Boué and Laskar, 2006)
- Takes into account the tidal deformation of the Earth, or post glacial deformation
- Do not depend on the internal model of the Earth
- The precession motions of K and w are the same

Suggestion : Any rotational solution for the Earth should provide also its angular momentum solution !