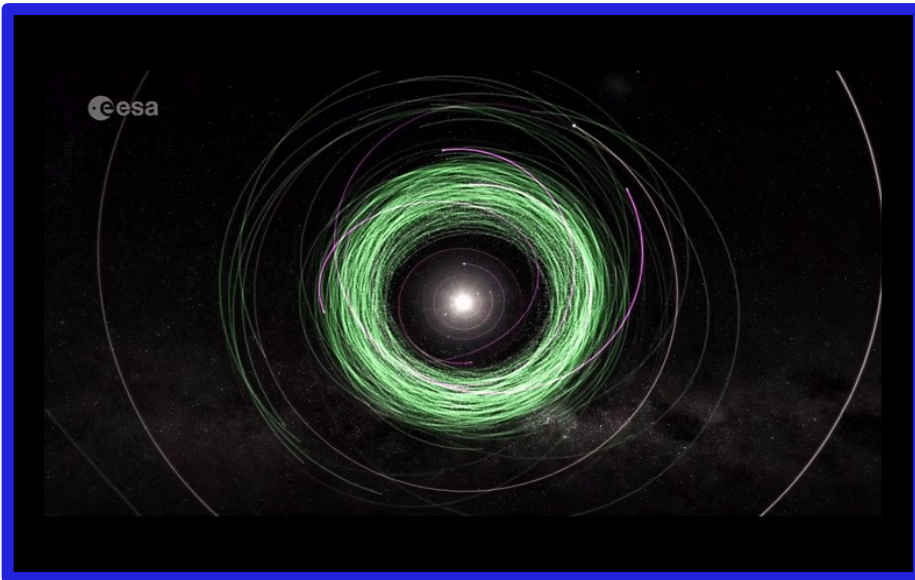


# Asteroid astrometry in the Gaia era

*How Gaia observations are changing  
our view of asteroid astrometry*

**F. Spoto<sup>1</sup>, P. Tanga<sup>1</sup>, B. Carry<sup>1,2</sup>, D. Michalik<sup>3</sup>**

Observatoire de la Côte d'Azur, Laboratoire Lagrange<sup>1</sup>  
Observatoire de Paris, IMCCE<sup>2</sup>  
ESA/ESTEC<sup>3</sup>

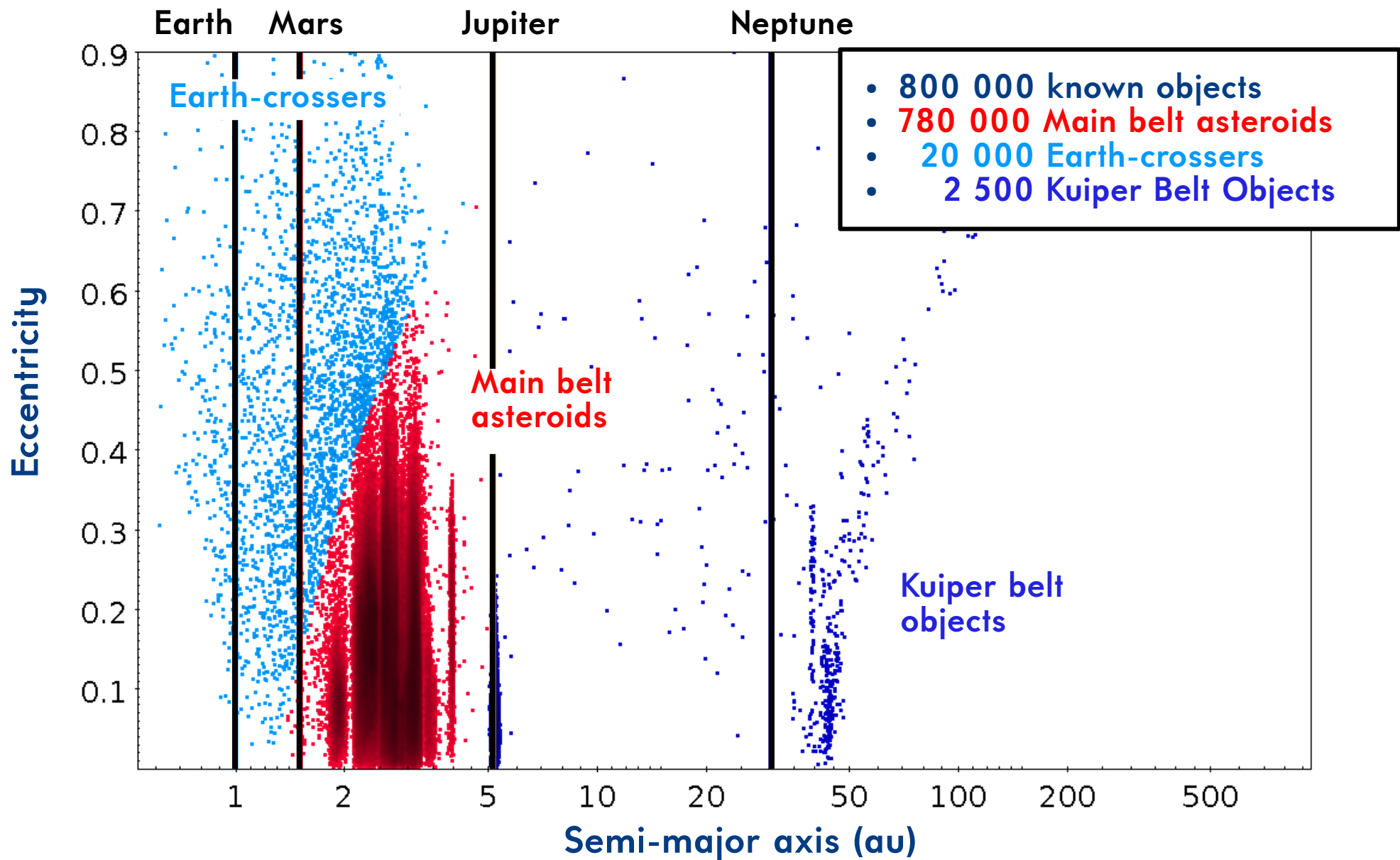


**Asteroid in our Solar  
System observed  
by Gaia**

Visualization: Gaia Sky

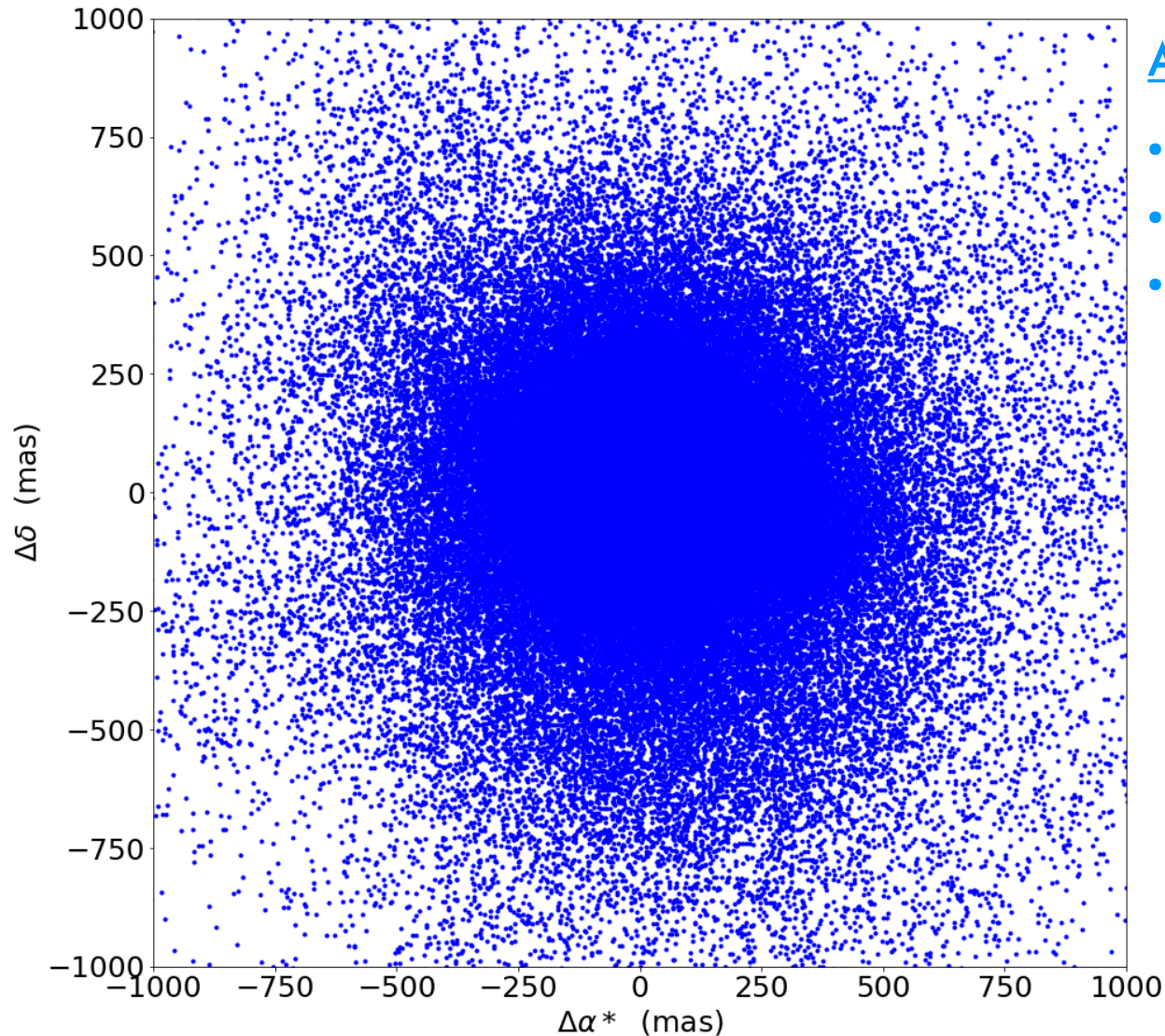
*“Journées 2019:  
Astrometry, Earth Rotation and Reference Systems in the Gaia era”  
Paris, 7/10/2019*

# Our knowledge of the asteroid population



# Typical asteroid observation residuals

Post-fit residuals on the sky : Observed - Computed

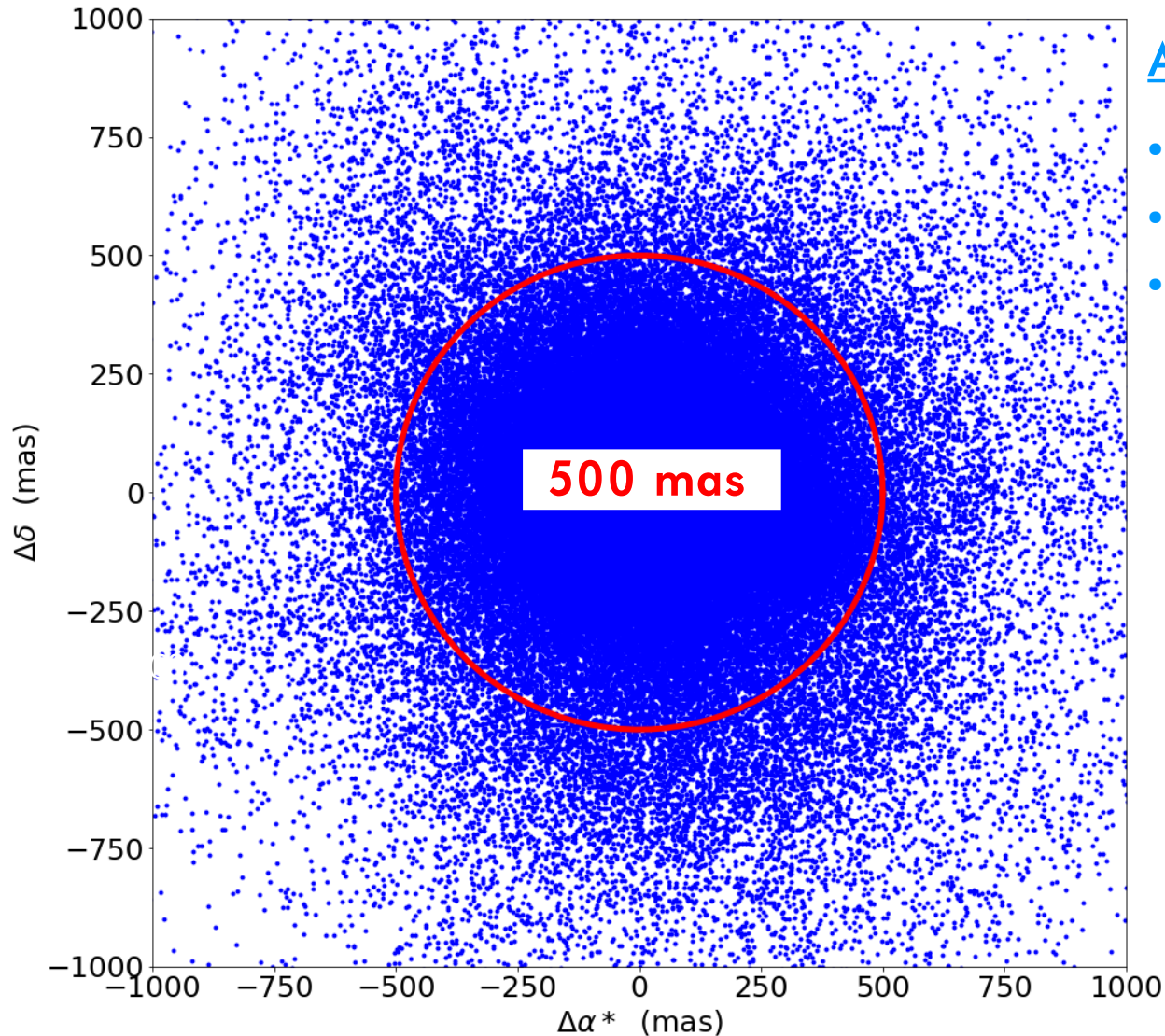


## Available ground-based astrometry

- 200 millions of observations
- Typical accuracy: **400 / 500 mas**
- 2 000 accurate observations (mostly radar)

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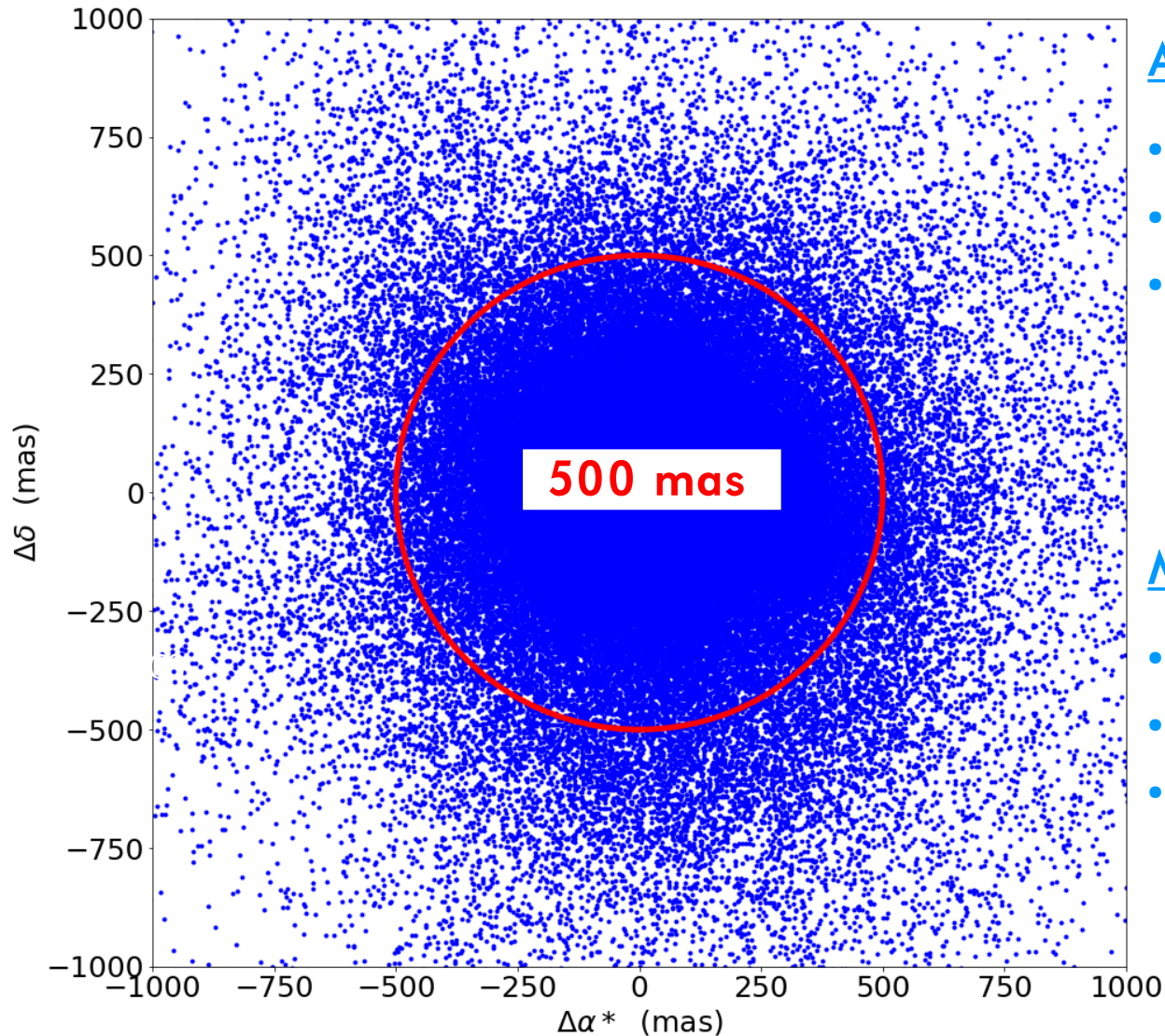
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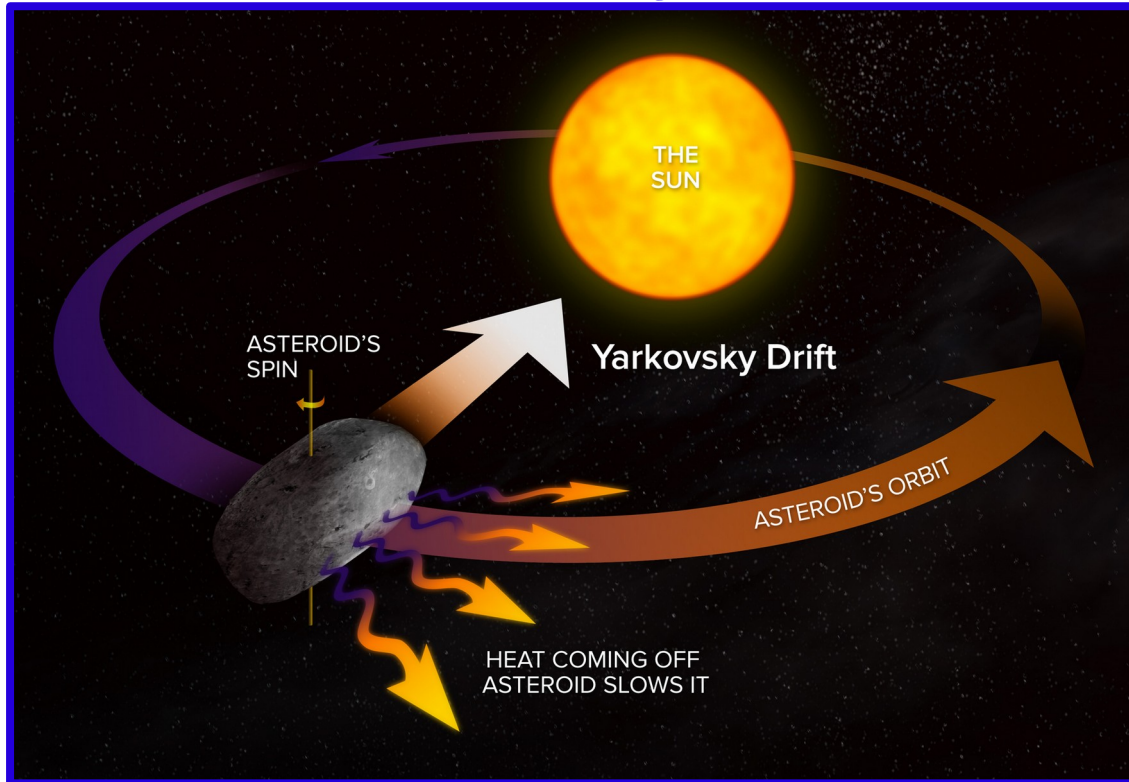
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## Main consequences

- Orbital elements : large uncertainties
- Poorly known orbits
- Observations focused on NEAs

# Why do we need good astrometry ?

## The Yarkovsky effect



### Description

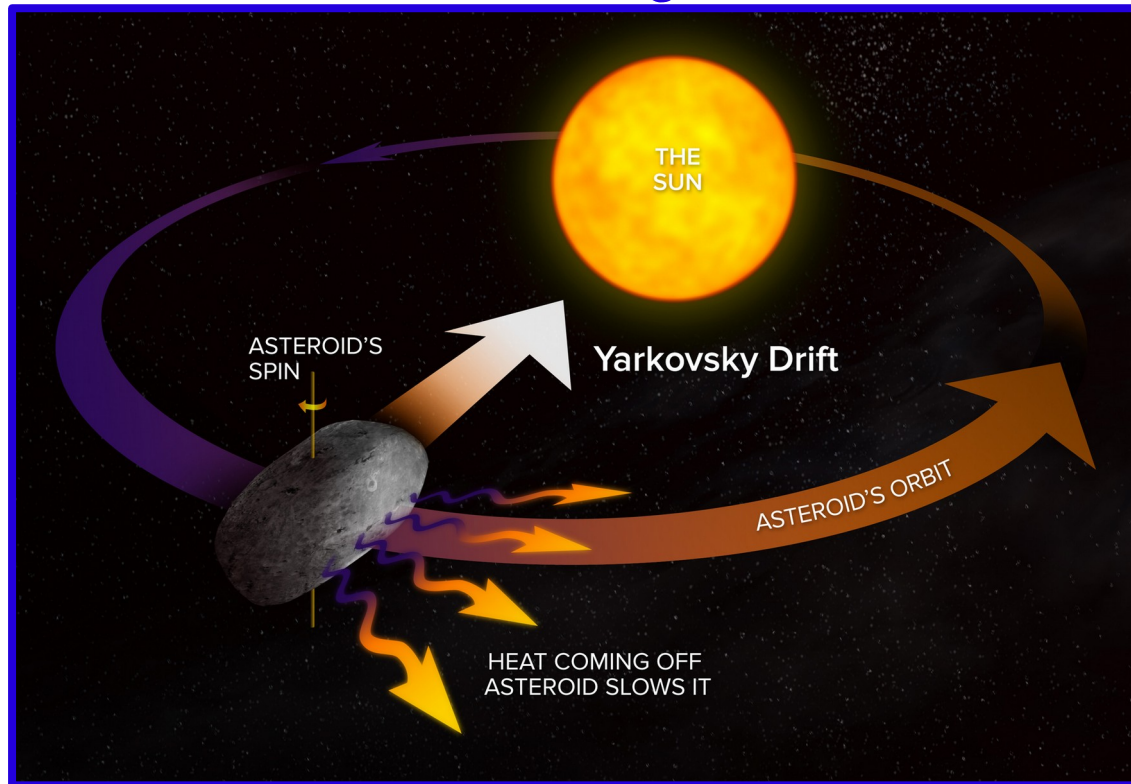
- Subtle **non-gravitational** perturbation
- Resulting from the **anisotropic thermal emission** of the solar radiation
- Dependence on **physical parameter** usually unknown

### Consequences

- **Secular semi-major axis drift**
- Necessary to **understand the evolution of our Solar System**
  - **Collisional history**
  - **Delivery of NEAs from the Main Belt**

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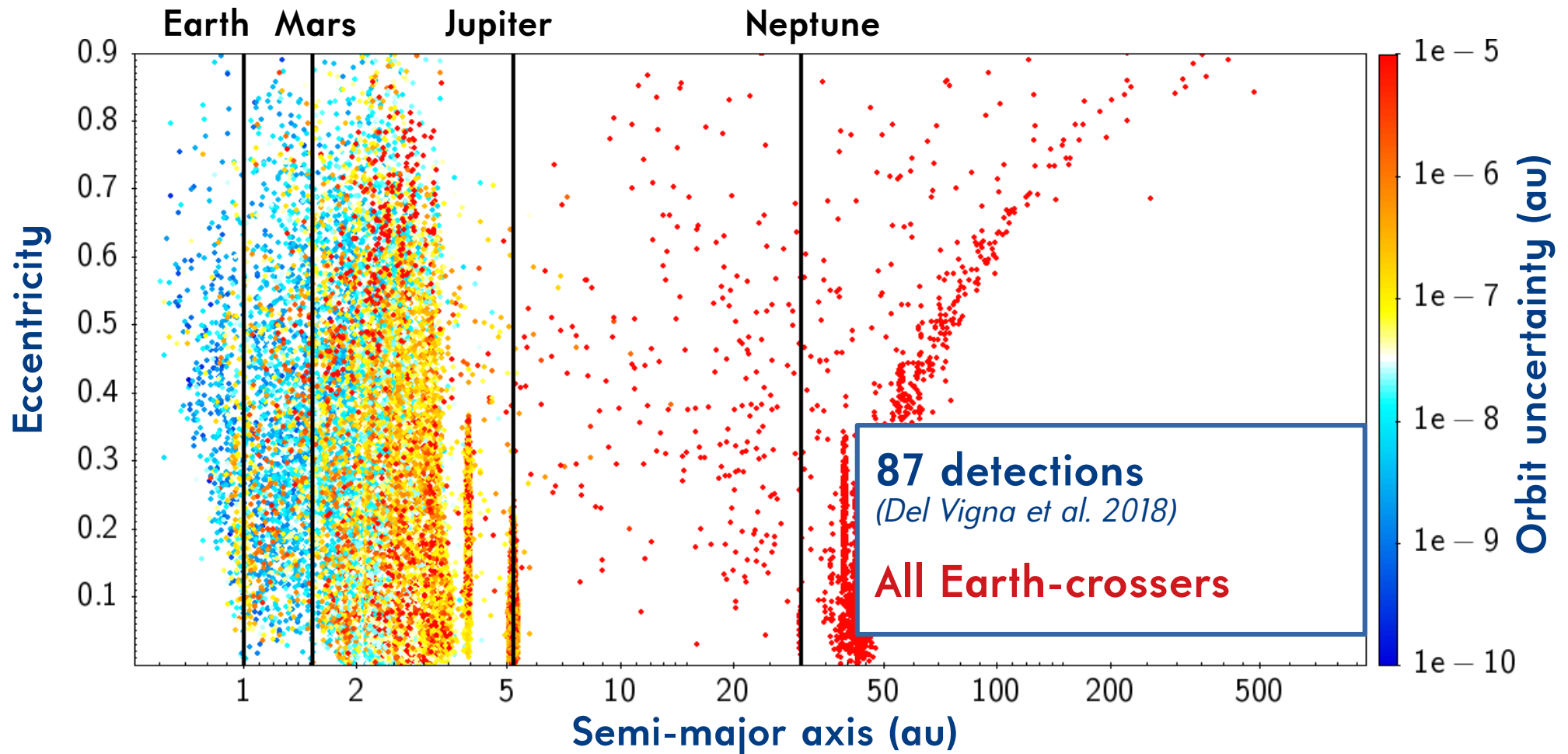
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### Detections

- **Detected from the astrometry:** least-square orbit determination (6 orbital elements + Yarkovsky parameter)
- **Very accurate orbits & long time span**
- **Small objects** (proportional to  $1/D$ )

# Yarkovsky detections (before Gaia)

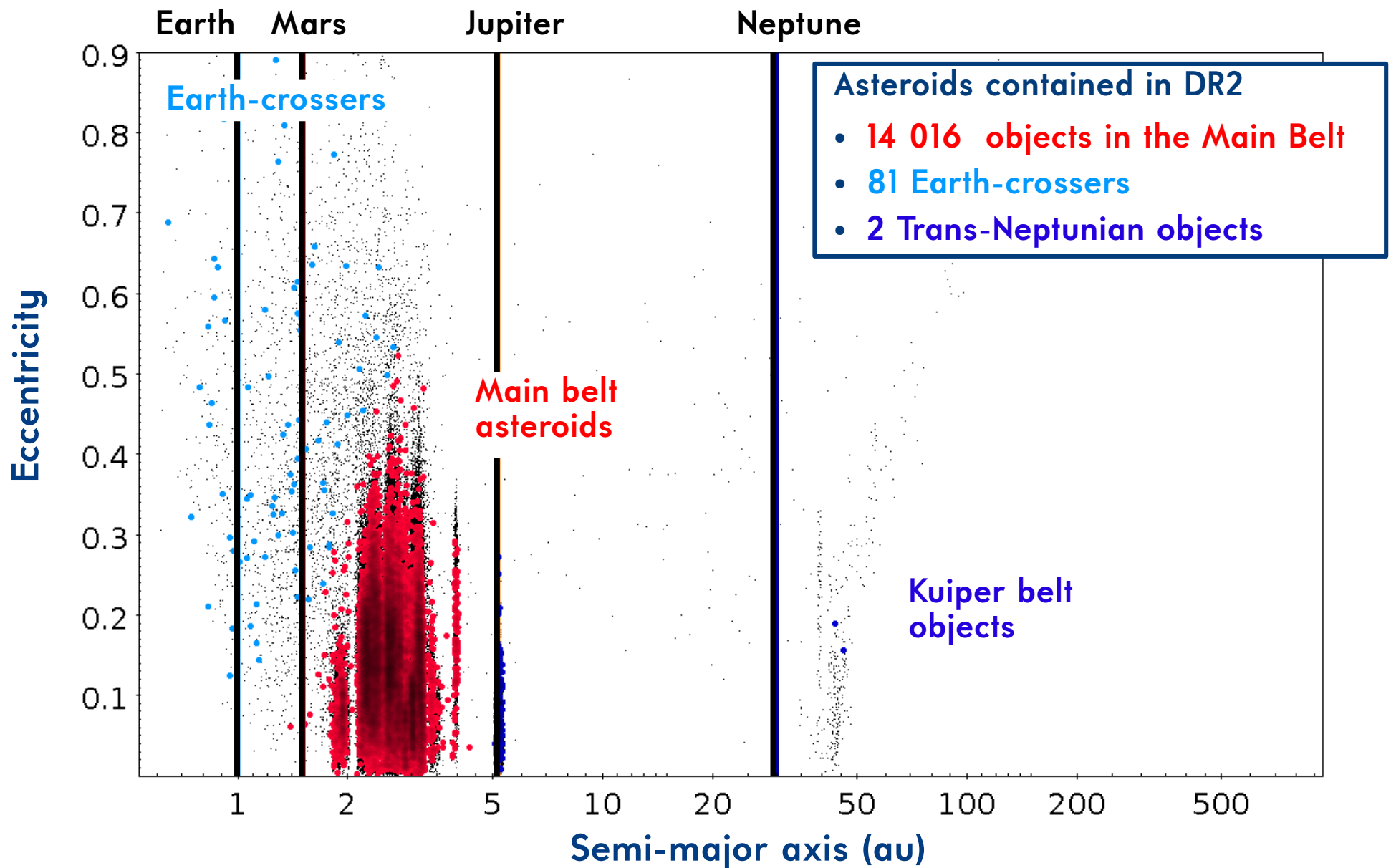


800 000 asteroids

900 small with « good » orbits  
uncertainty  $< 3e-9$  au ( $\sim 450$  m)

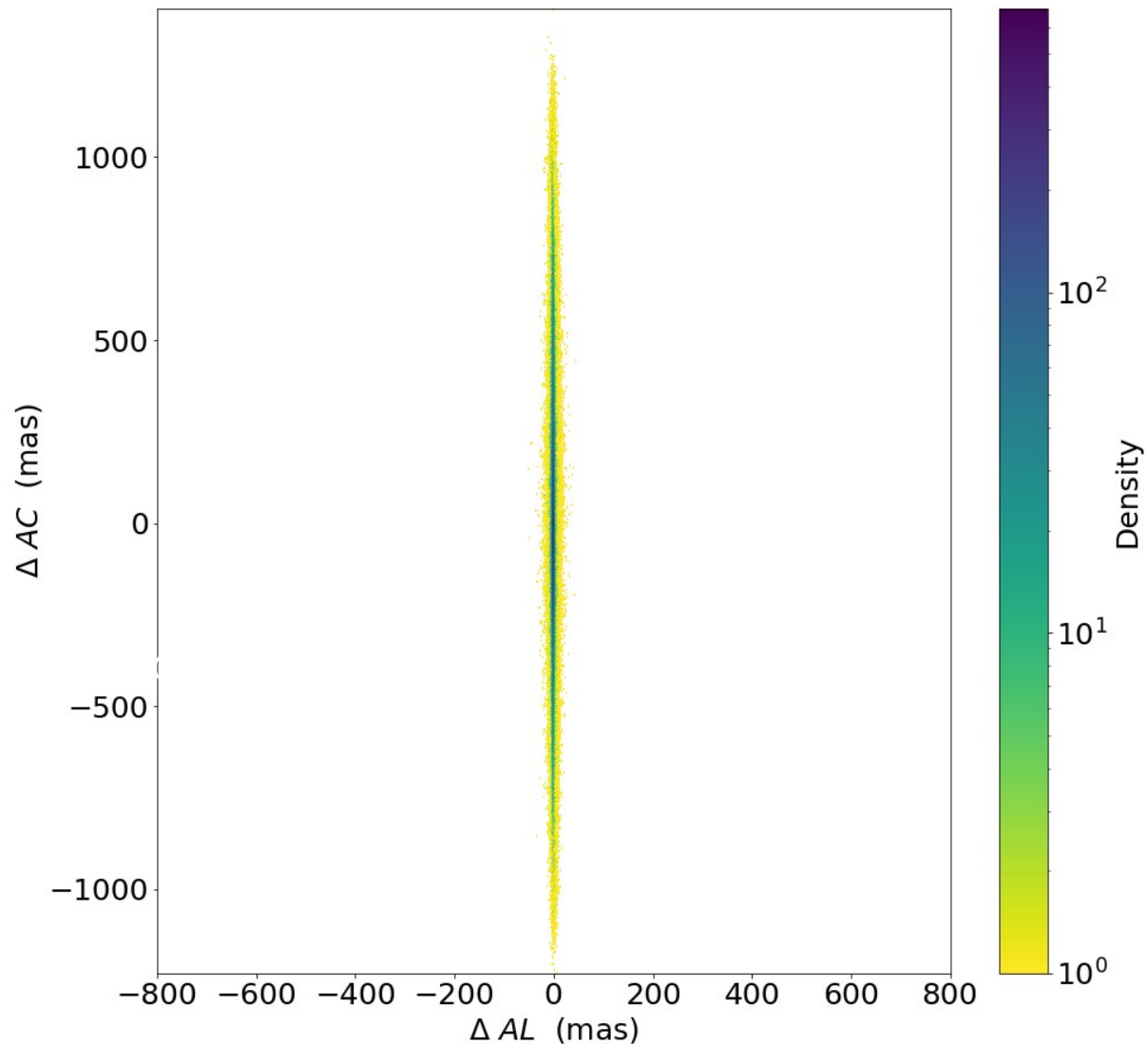
No Main Belt asteroids







## Post-fit residuals in the **ALong** scan – **ACross** scan plane

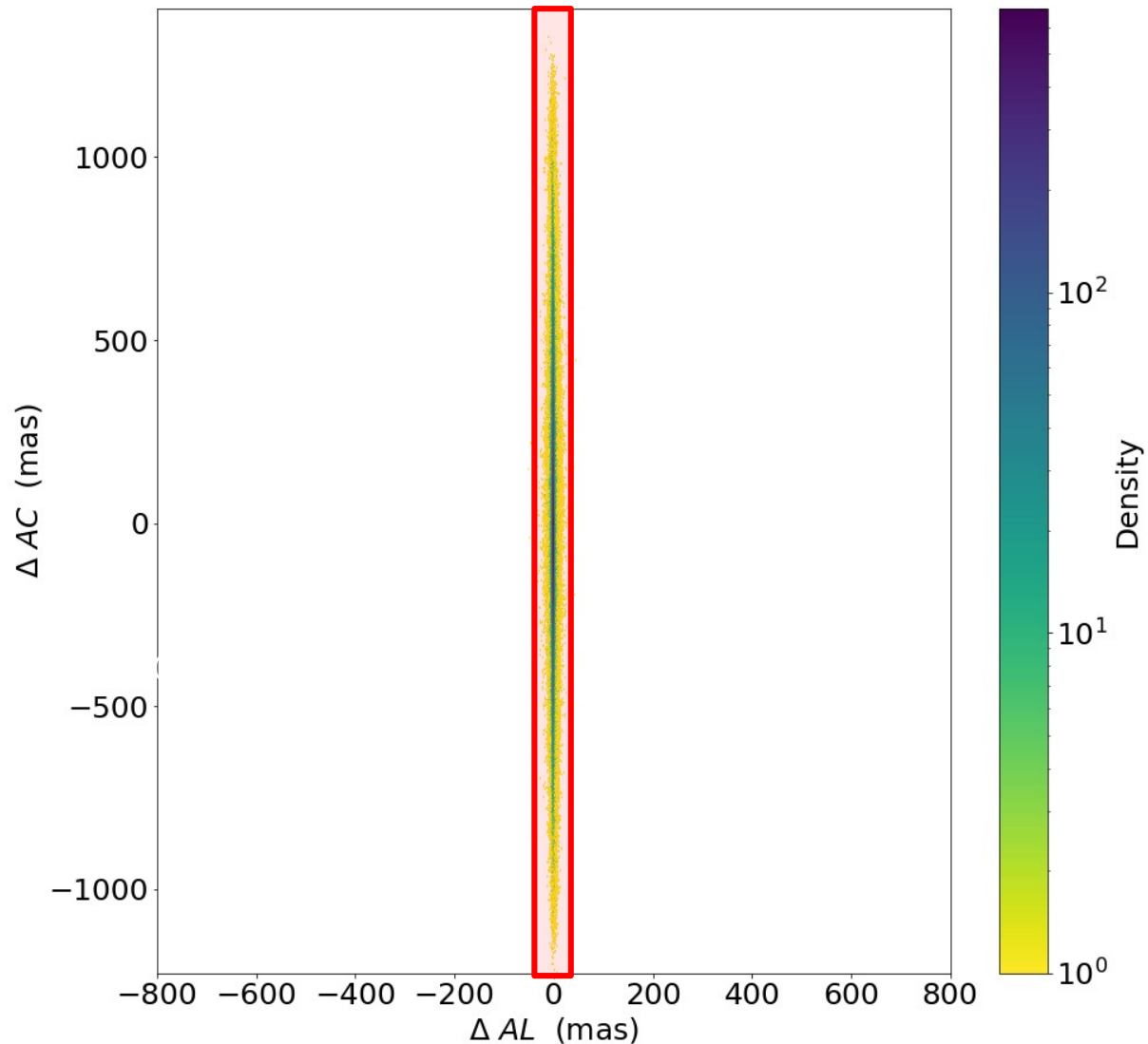


### Gaia asteroid astrometry

- ~ **2 millions** of observations
- 22 months
- Accuracy is in the **ALong** scan direction



Post-fit residuals in the **ALong** scan – **ACross** scan plane

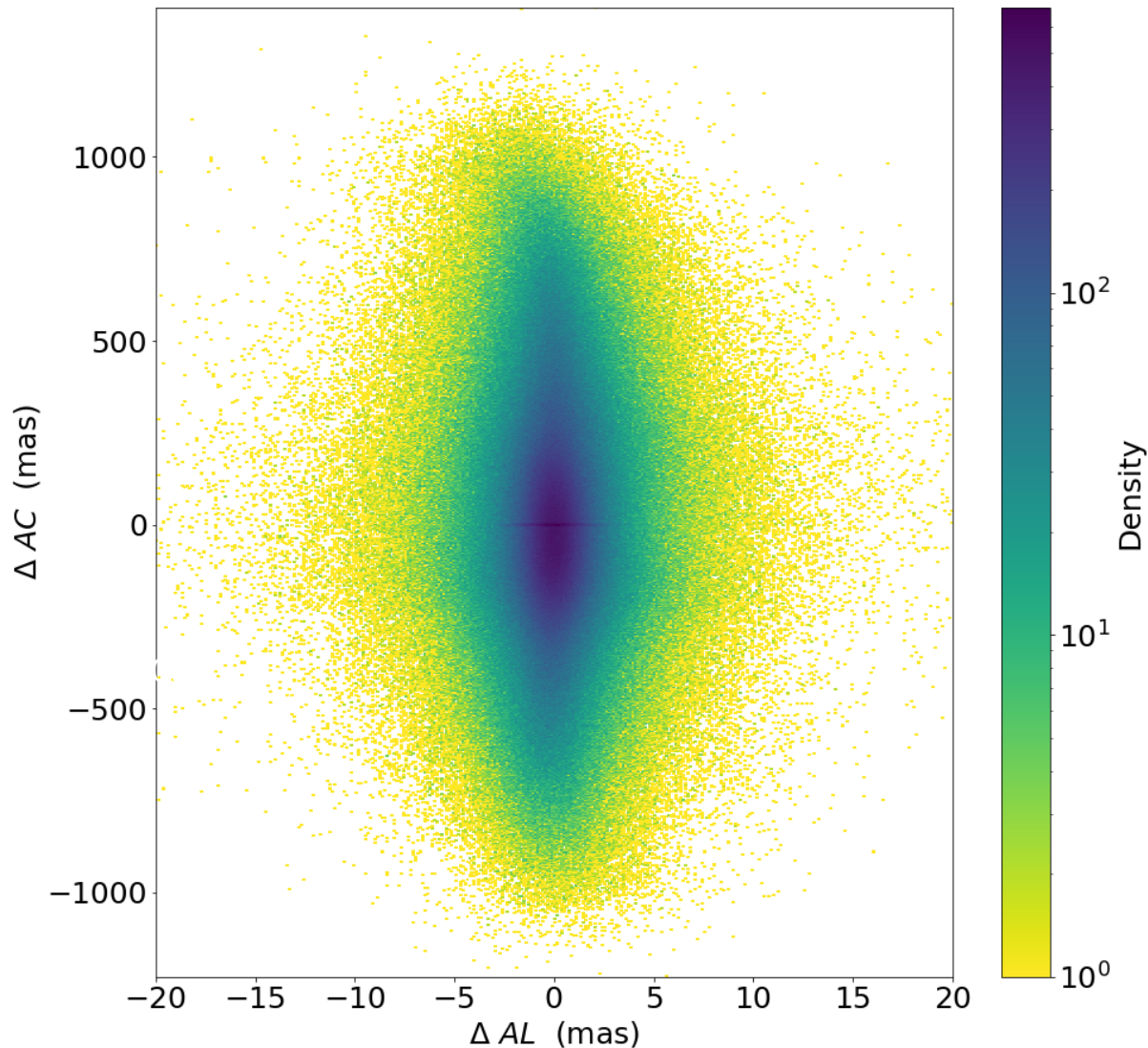


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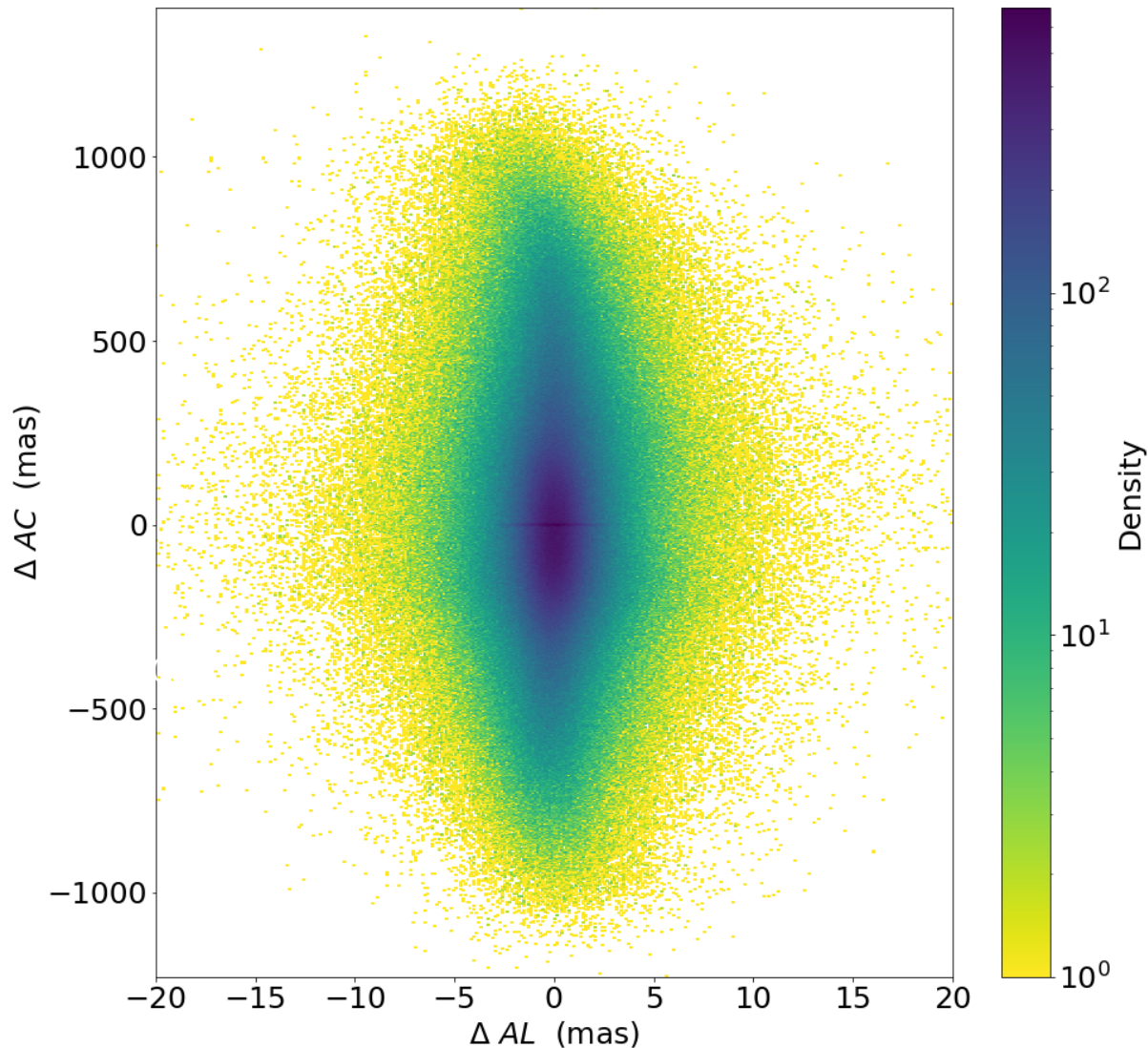
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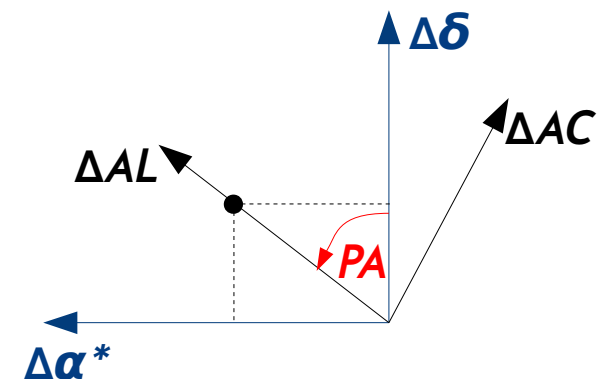
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### Gaia asteroid astrometry

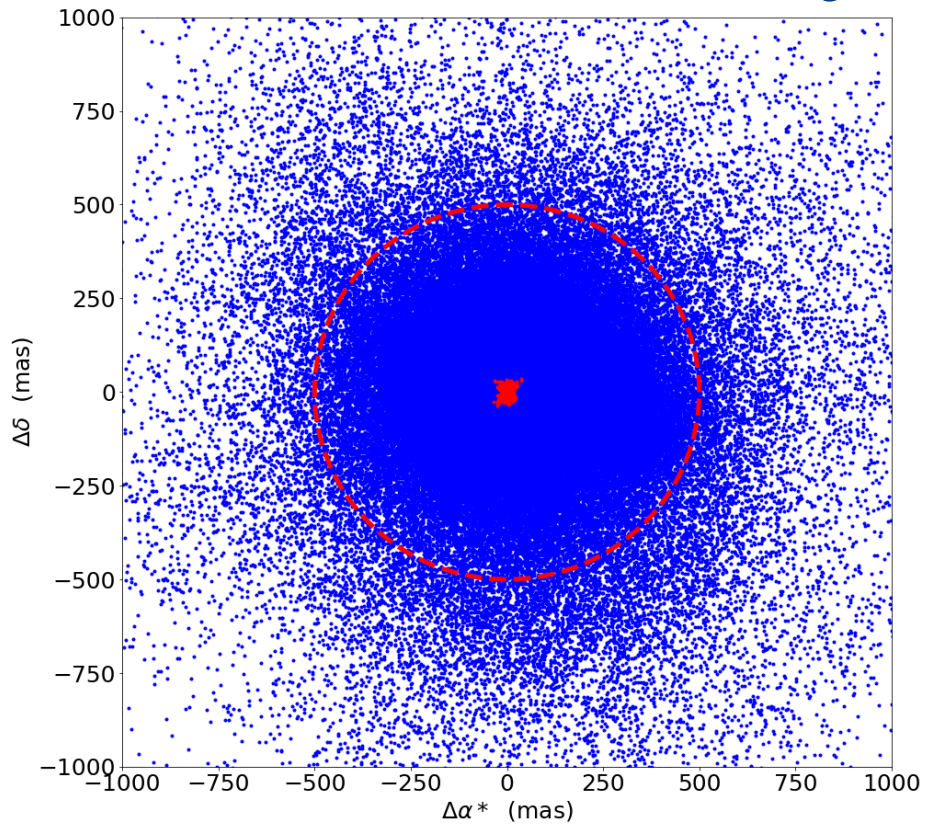
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*How can we make a comparison between Gaia and ground-based astrometry ?*





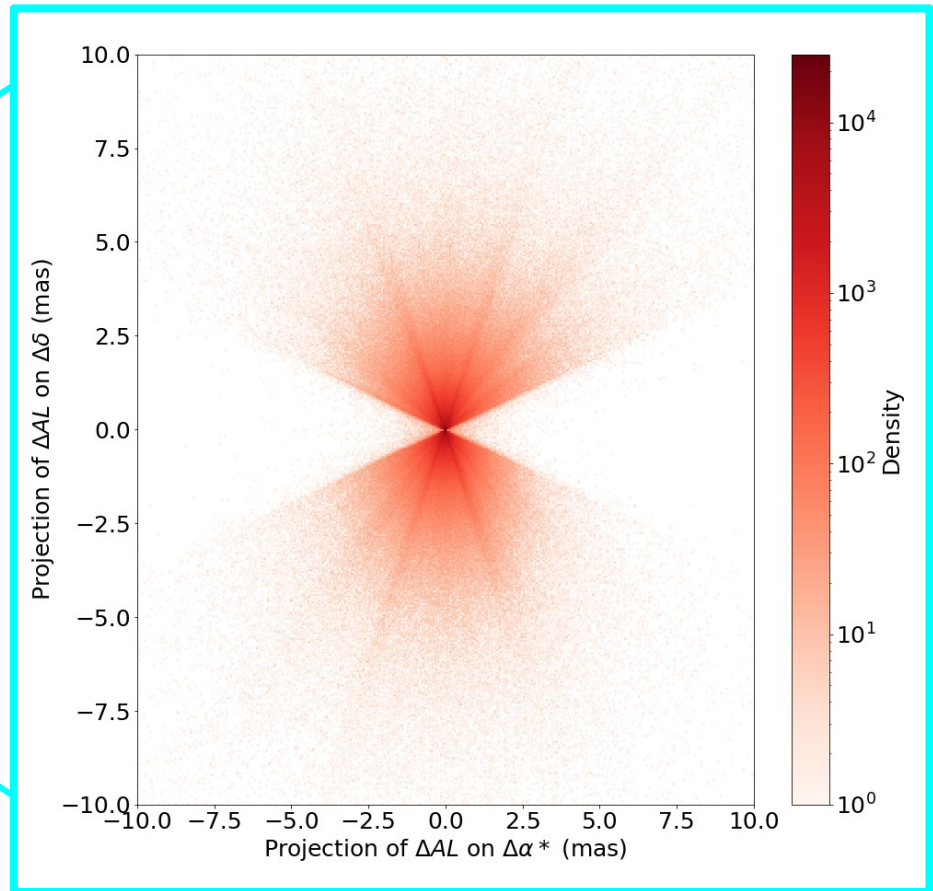
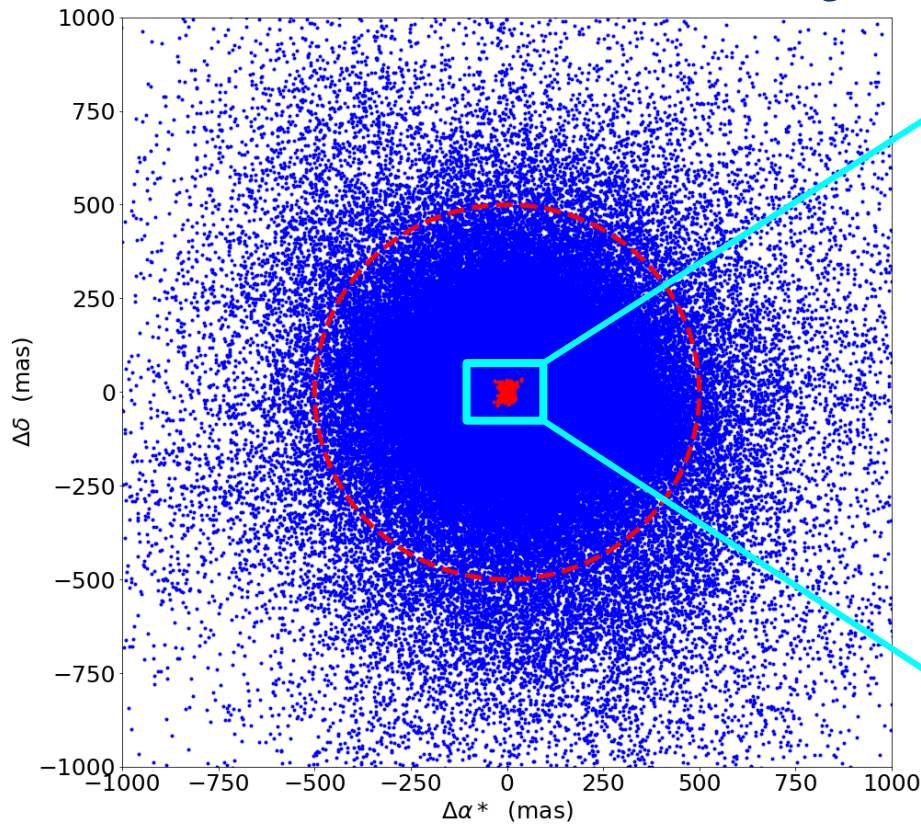
## Post - fit residuals on the sky



# Gaia DR2 vs ground-based asteroid astrometry



## Post - fit residuals on the sky



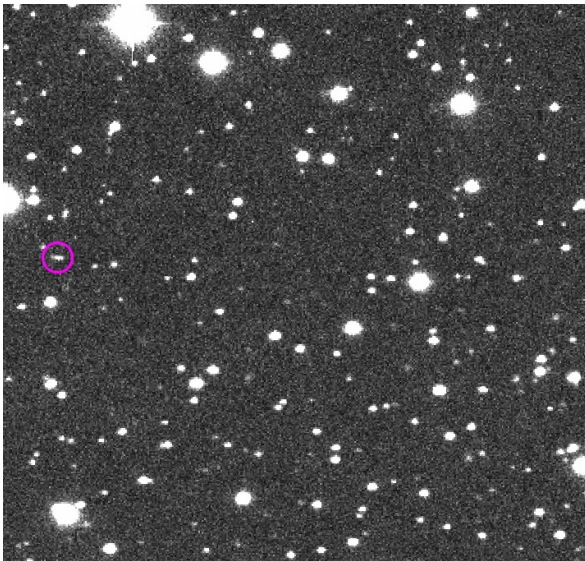
### Main « issues » :

- 2 millions of very accurate observations
- New vision of asteroid astrometry
- Short observational arc

# How to combine Gaia and ground-based observations?



## 1. Debiasing of old stellar catalogs



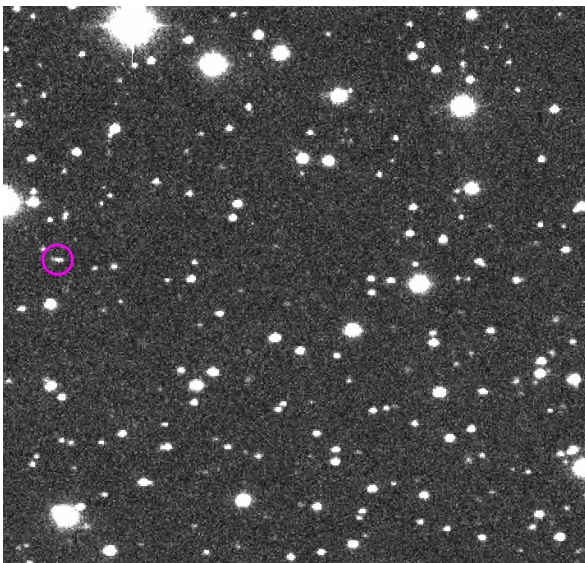
Discovery  
observations of the  
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Discovery  
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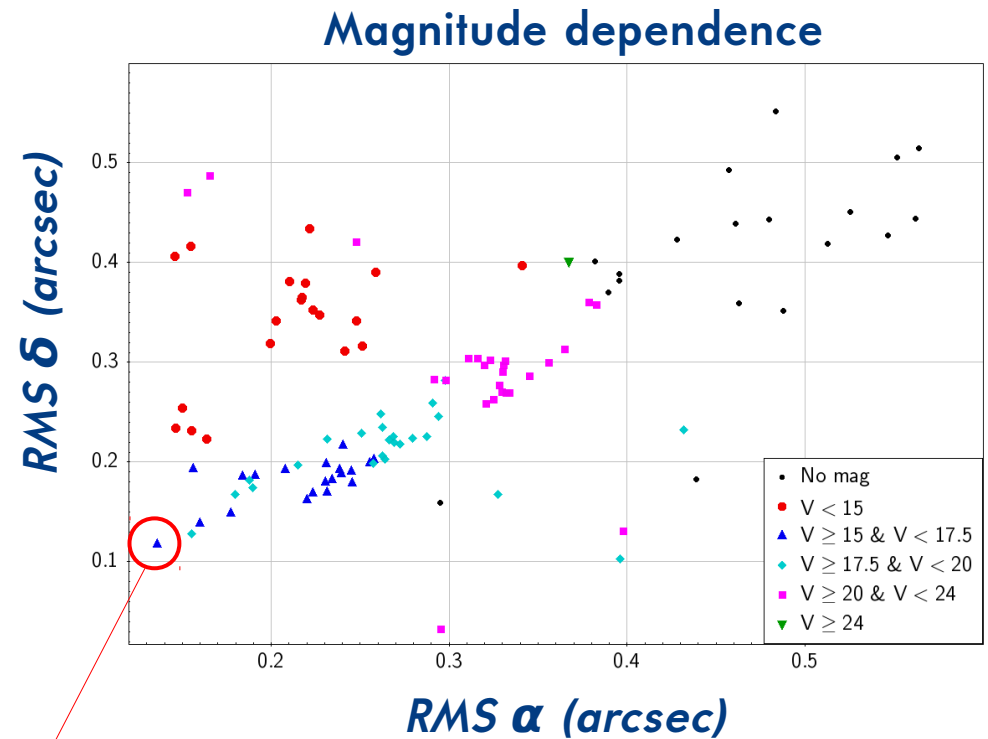
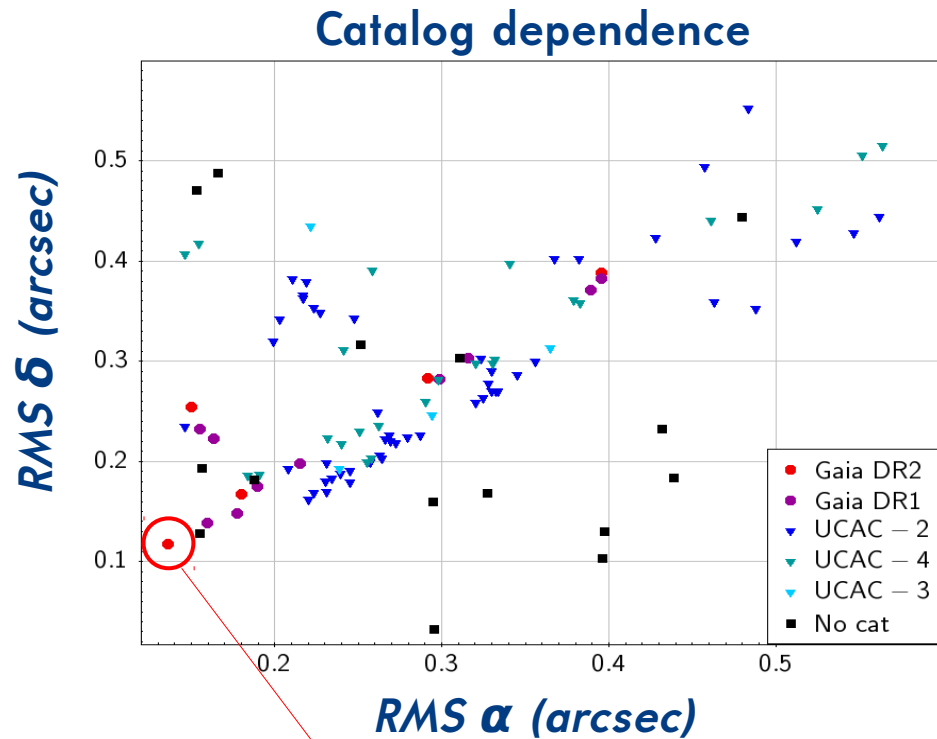


Next talk : Paolo Tanga

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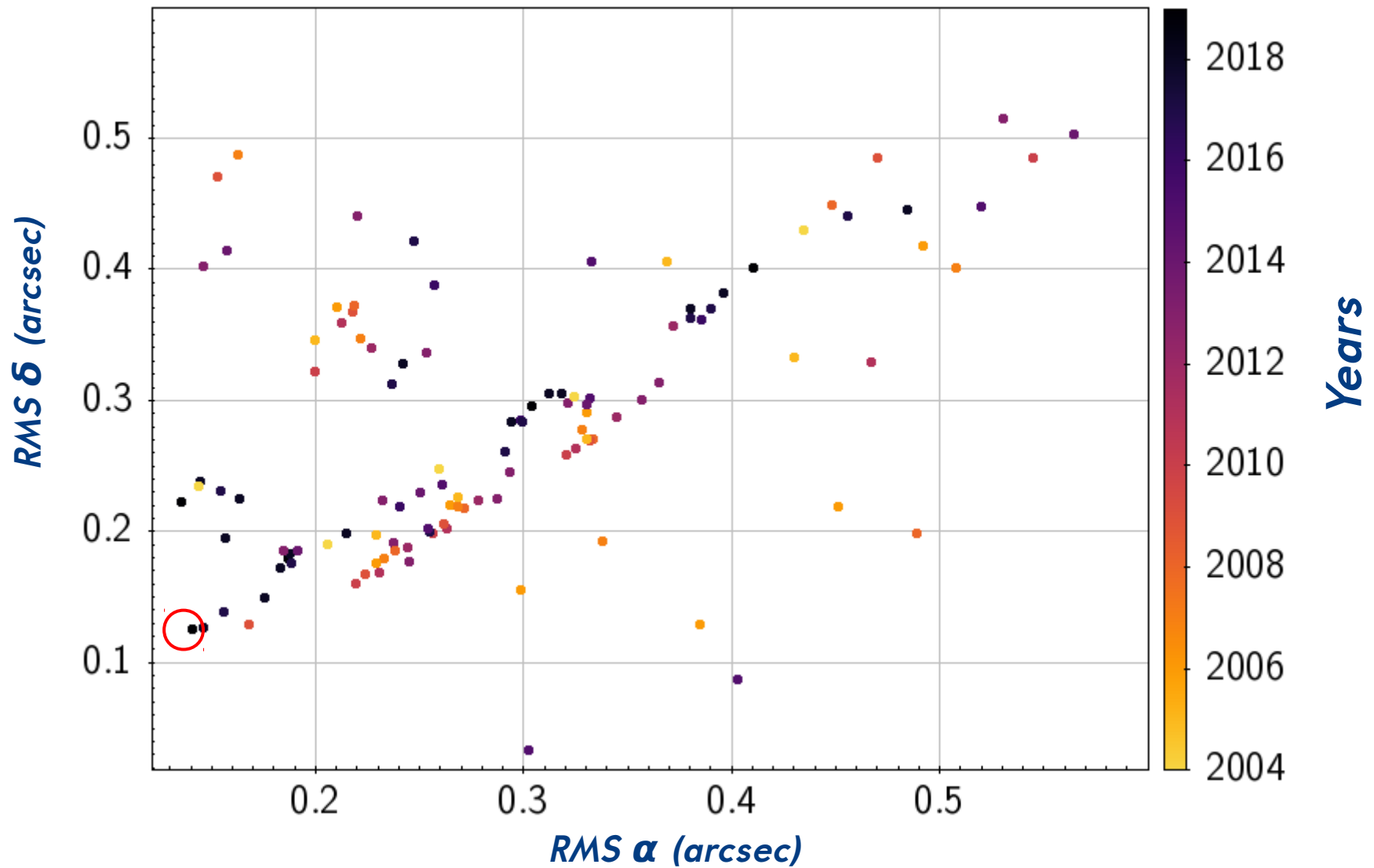
## 2. Error model : weights to give to each observatory



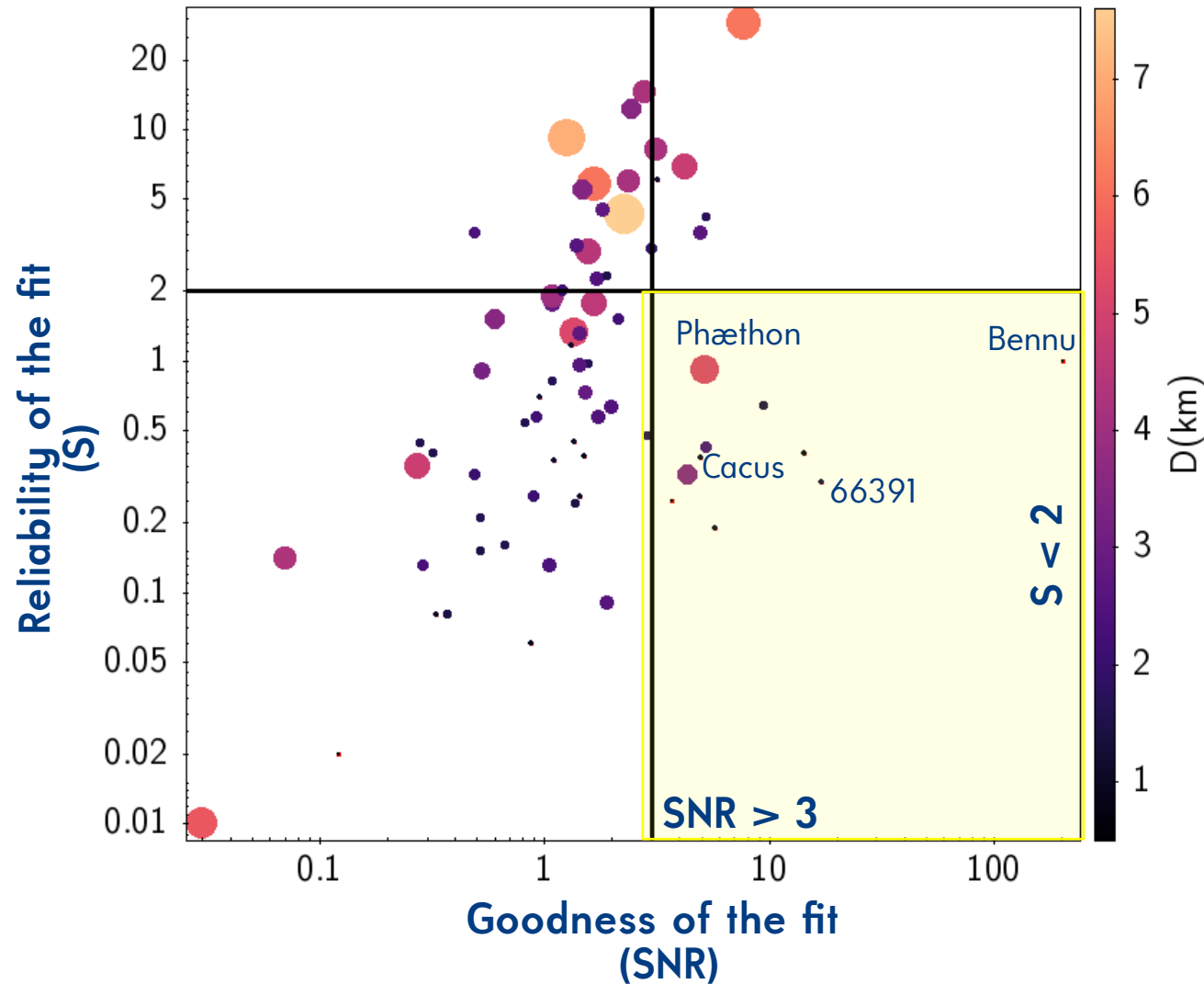
(Observatory, year, catalog, magnitude class)  
(G96, 2019, V, 2)



## 2. Error model : weights to give to each observatory



# Results : new Yarkovsky detections



## Initial sample from Gaia DR2

- ~ 60 objects
- Diameter < 10 km
- Orbit uncertainty < 1 km

## Detections

- 10 new + densities
- Cases not possible before Gaia



# Conclusions

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- **Gaia** has already changed our view of the asteroid astrometry
- Our knowledge of the Main Belt is still very limited:  
**we are missing quantity and quality**
- We are on the verge of a revolution : Gaia is producing ultra-accurate astrometry for millions of observations
- We need to combined Gaia and tens of ground-based observations to detect subtle non-gravitational perturbations like the Yarkovsky effect
- We have analyzed and corrected all the available astrometry
- The combination has already produced amazing results, but moreover it shows that now we are ready for the next Gaia releases
- **To the Main Belt and beyond.**