Excitation of the Earth's Chandler wobble by the North Atlantic double-gyre

S. Elnaz Naghibi, Sergey A. Karabasov Department of Aeronautics, Imperial College London School of Engineering and Materials Science Queen Mary, University of London

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Imperial College London

Chandler wobble

Chandler Wobble, the main component of polar motion, is a 14-month free motion, the period of which is determined by elliptic geometry and the rigidity of the Earth .

Chandler wobble equations

$$\frac{i}{\sigma_0 + i/2Q} \frac{d\mathbf{m}}{dt} + \mathbf{m} = \mathbf{\Psi} = \left[1 - \frac{i}{\Omega} \frac{d}{dt}\right] \left\{\alpha \mathbf{c} + \beta \mathbf{h}\right\}$$

$$T_0 = \frac{2\pi}{\sigma_0} \approx 433 \text{ days}$$



Background: general circulation ocean models & double-gyre models



Naghibi et al., Geophysical Journal International, 2017 Naghibi et al., Applied Mathematical Modelling, 2019



HYCOM (HYbrid Coordinate Ocean Model):

A general ocean model with realistic continent boundaries, bottom topography and time varying wind forcing

Quasi-geostrophic model (QG):

A mid-latitude approximation of the double-gyre problem with steady wind forcing, flat bottom topography and longitudinal-latitudinal boundary walls

HYbrid Coordinate Ocean Model:

Conservation laws for momentum, energy, salinity, mass and the equation of state

$$\begin{aligned} \frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v}.\nabla)\mathbf{v} + 2\mathbf{\omega} \times \mathbf{v} &= -\frac{\nabla M}{\rho} + \frac{\nabla.\tau}{\rho}, \\ \frac{\partial(\Delta hT)}{\partial t} + \nabla.(\Delta hT\mathbf{v}) &= \nabla.(\nu \Delta h\nabla T) + F^{T}, \\ \frac{\partial(\Delta hS)}{\partial t} + \nabla.(\Delta hS\mathbf{v}) &= \nabla.(\nu \Delta h\nabla S) + F^{S}, \\ \frac{\partial}{\partial t}(\Delta h) + \nabla.(\mathbf{v}\Delta h) &= 0, \\ \rho &= \rho(T, S, P), \end{aligned}$$

The model considers 40 isopycnal layers and the outputs are provided over a uniform 2250×4500 resolution grid.

Quasi-geostrophic double-gyre model:

stratified potential vorticty eqautions with meridional gradient of planetary vorticity, lateral viscosity & bottom friction and wind forcing
$$\begin{split} \partial_t q_i + J(\psi_i, q_i) &= \delta_{1i} F_w - \delta_{i3} \frac{a_v}{H_3^2} \nabla^2 \psi_i + a_h \nabla^4 \psi_i \\ q_i &= \nabla^2 \psi_i + \beta_y - (1 - \delta_{i1}) S_{i1} (\psi_i - \psi_{i-1}) - (1 - \delta_{i3}) S_{i2} (\psi_i - \psi_{i+1}) \end{split}$$

(Karabasov et al. 2009, Shevchenko & Berloff 2015)

boundary conditions

$$\partial_{\mathbf{nn}} \psi_i - \alpha^{-1} \partial_{\mathbf{n}} \psi_i = 0$$

 $i = 1, 2, 3$

numerical solution

CABARET: Compact Accurately Boundary-Adjusting high-REsolution Technique (Karabasov & Goloviznin 2009)

package

PEQUOD: Parallel Quasi-Geostrophic Model, (Maddison, Berloff & Karabasov 2014)



HYCOM outputs





QG outputs





Comparison of Chandler wobble excitation functions using HYCOM & QG velocity fields



Comparison of the meanflow and RMS profiles in HYCOM & QG



Chandler wobble excitation functions: Global oceans vs. the North Atlantic (motion term)



North Atlantic is not the largest contributor in Chandler wobble excitation



(Naghibi et al., 2017)

Future work:

- Calculating Chandler wobble excitation over a longer period of time using HYCOM & QG outputs
- Comparing the predicted excitation with geodetic observation of the Chandler wobble after subtraction of mass terms estimated from geophysical GRACE satellite gravimetry data



Thank you