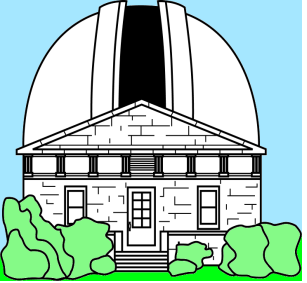


Deep South Telescope DST

Norbert Zacharias
Charlie Finch
(for the team)

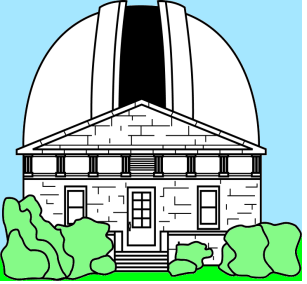
U.S. Naval Observatory
Astrometry Department
Optical Reference Frame Div.

norbert.zacharias@navy.mil
nzIAUc8@gmail.com



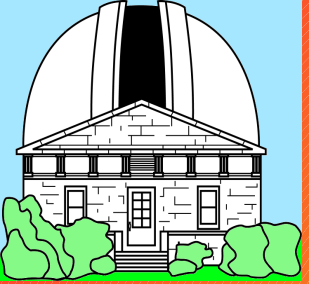
the team ...

- USNO instrument shop:
 - Gary Wieder, Chris Kilian, Phil Eakens
- FRAMEx collaboration:
 - Valeri Makarov, Nathan Secrest, Bryan Dorland, Jean Souchay
- Contractors:
 - Greg Bredthauer, Marion Zacharias
- CTIO: Esteban Parkes ...
- PlaneWave:
 - Ralph Emerson, Kevin Ivarsen, Matt Dietrich ...
- Princeton Instruments: Mike Melle ...

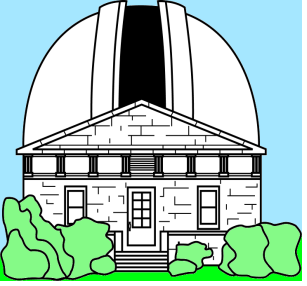


Layout of talk

- History of DST project
- Deployment of instrument
- Science program
- Target selection
- Status of operations
- Conclusion

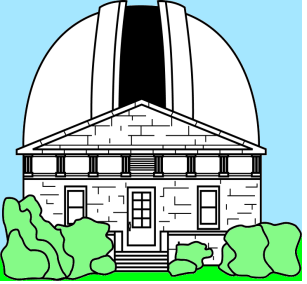


History of DST project



timeline

- 2015 / 2016 begin planning
 - era of astrometric all-sky surveys from ground comes to an end due to Gaia success
 - need large aperture (“going deep”) telescope in the south
- 2017 market research, budget approved, get PW1000
- 2018 :
 - order 4k CCD camera from Princeton Instruments
 - construction of telescope completed
 - December: test observing at PW factory in California near LA



PlaneWave CDK1000

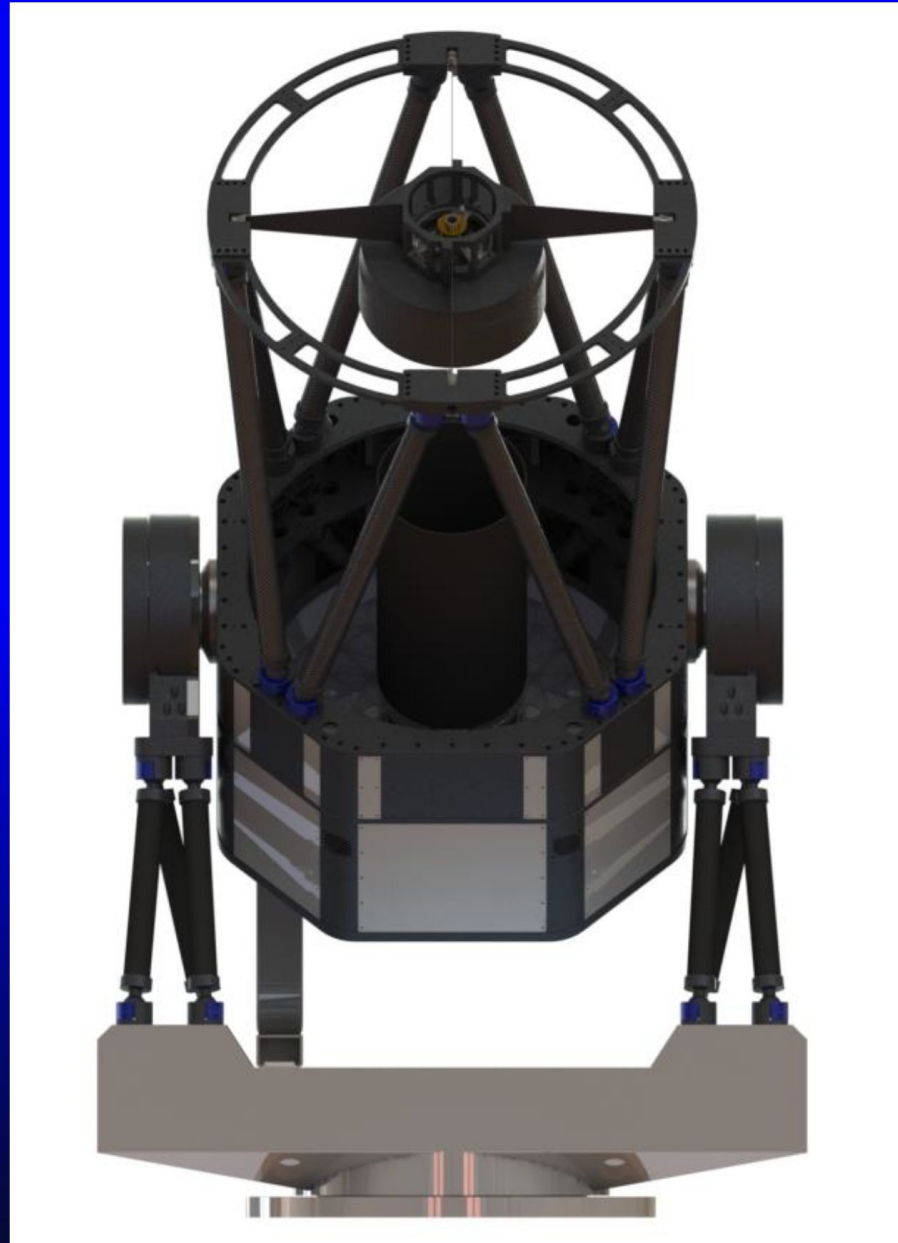
Fits in 16' dome

Focal plane is at eye height

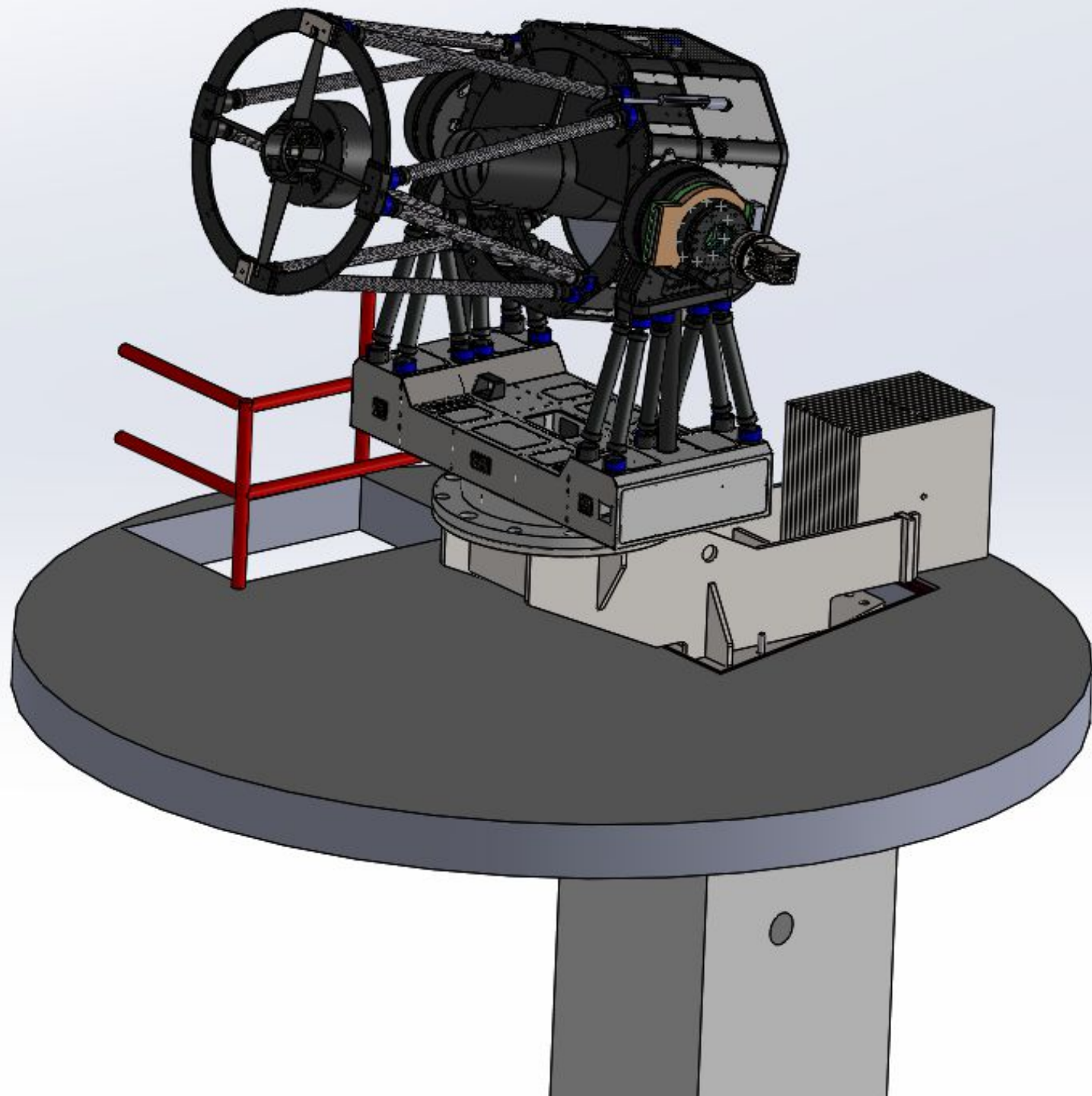
Optical benches on the sides are de-rotated

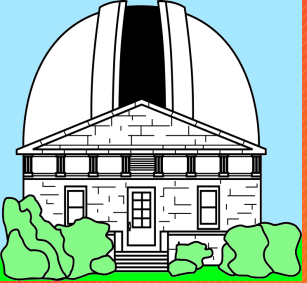
All cables and electronics are inside truss poles, forks, base and cable tray

Easy access to cables and electronics

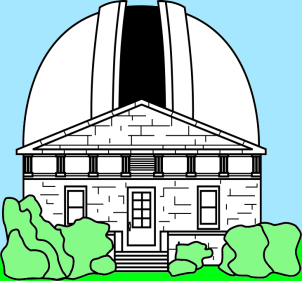


- 1.0 m apert.
- $f = 6.0$ m
- 2 foci
 - optical
 - near-IR
- robotic





Deployment of Instrument at CTIO



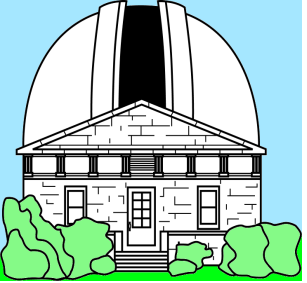
2019

- 4k camera delayed: accident during shipment
- March: deploy telescope at CTIO w. small test camera
- June, July: 2nd trip to CTIO:
 - finish system integration (except 4k camera), all devices
 - implement and test operations software
- August: 3rd trip to CTIO:
 - install 4k camera, finish up deployment
 - Aug 13: first light
 - Aug 24: operational



**March
2019
CTIO**





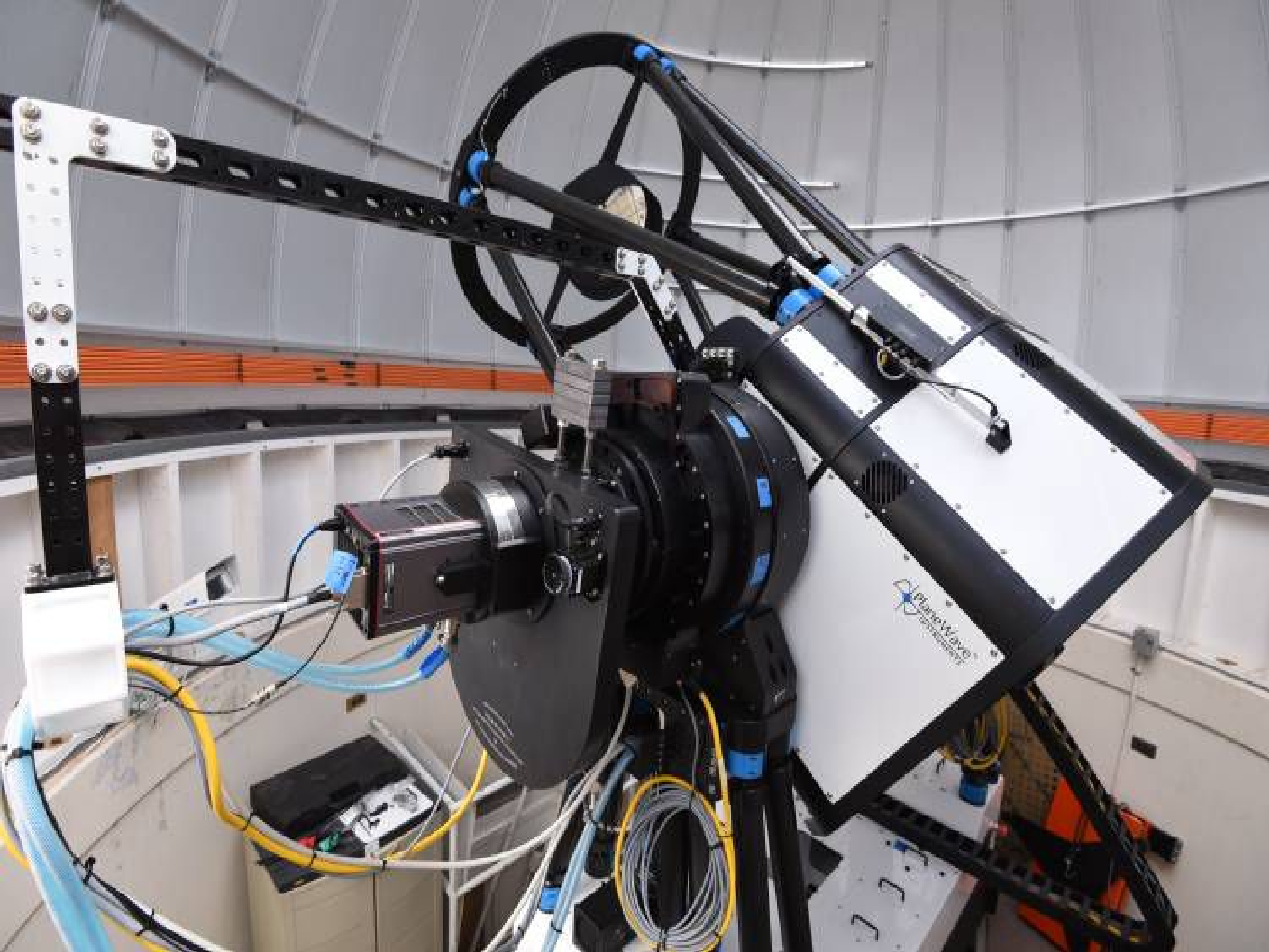
Princeton Instruments camera

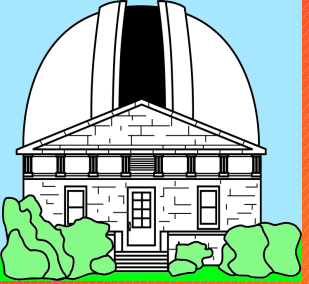
Products: SOPHIA Ultra-Low Noise CCD Cameras



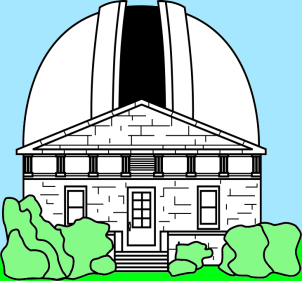
- 4k by 4k E2V CCD
- 4 port low-noise read
- lifetime vacuum
- thermoelectric cooling to -90 C
- no more liquid N2





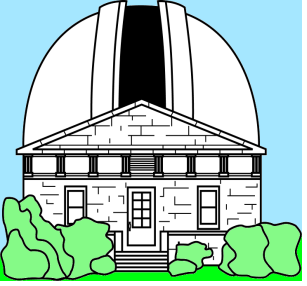


Science Program



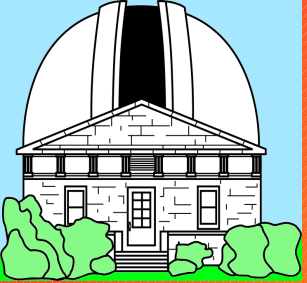
DST is used for FRAME_x

- follow-up on Gaia – VLBI “odd sources”
 - target extragalactic sources with **radio-optical position offsets**
 - **high cadence** observing with various filters (B,V,R,I)
 - is there a connection between astrometric offset and other parameters (brightness, color, structure ...)?
 - do optical observations from the ground (arcsec resolution) “see” a different center of light than Gaia does?

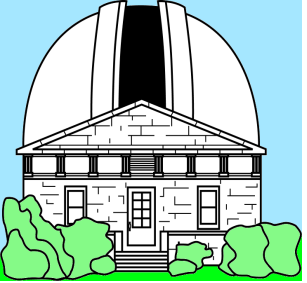


other DST projects

- 10% of time to **Chile** (likely begin next semester)
- 2nd port (focus) of DST for **near-IR camera**
- mount can move non-sidereal, arbitrary tracks:
 - capable of observing **asteroids** incl. **NEOs, PHAs**
- other programs TBD in future

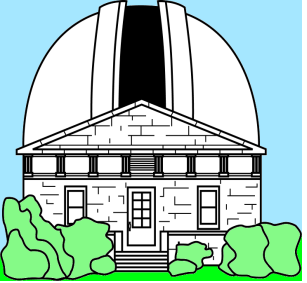


Target selection for FRAMEx program



DST extragalactic targets

- most sources are in [Gaia DR2](#) and have [VLBI](#) observ.
- 4 groups:
 - [nearby AGN](#) targets (Nathan Secrest), some need VLBI obs.
 - “odd” sources from [ICRF3](#) (Valeri Makarov)
 - [other](#) sources with significant radio-optical pos.difference (RFC)
 - “[control sample](#)” = no significant radio-optical pos.difference
- selected sources as bright as possible ($R \leq 18.5$)
- about even distribution along RA, $\text{Dec} \leq +25^\circ$

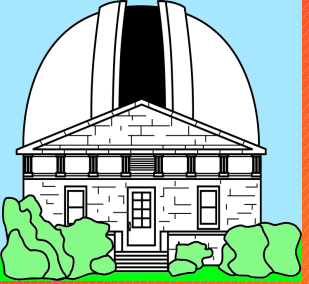


numb. of sources by
radio-optical pos.dif

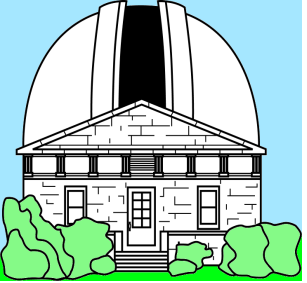
no info 23
up to 2.9 mas 50
2.9+ to 5.0 mas 14
5.0+ to 10.0 mas 32
10.0+ to 30.0 mas 38
30.0+ to 100.0 mas 15
100.0+ to 300.0 mas 8
over 12

total 192
= 4 per 0.5 h RA

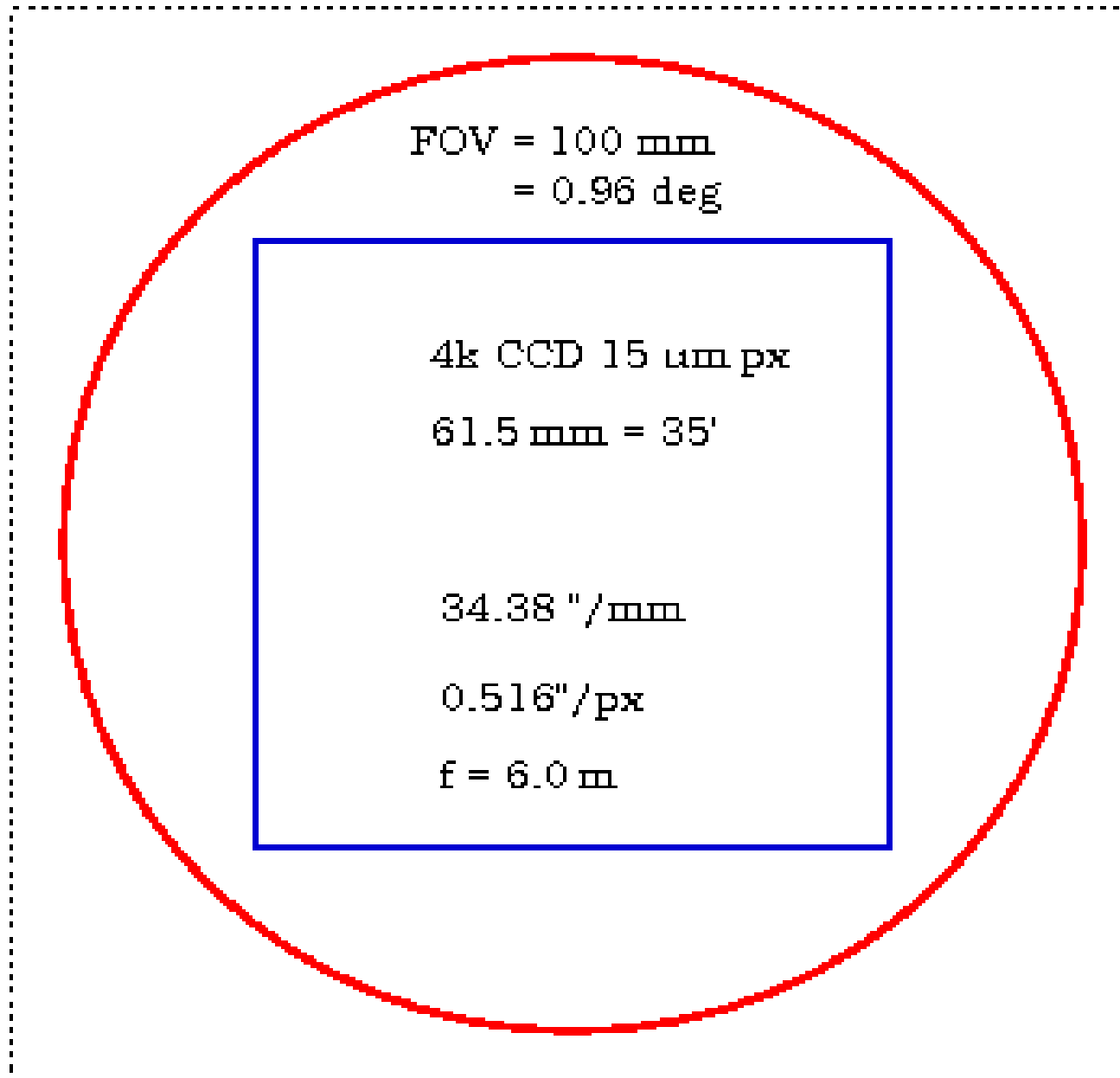


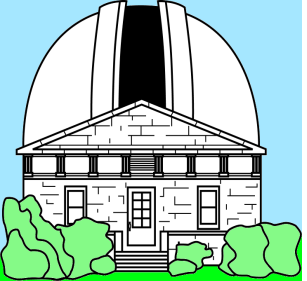


Status of operations



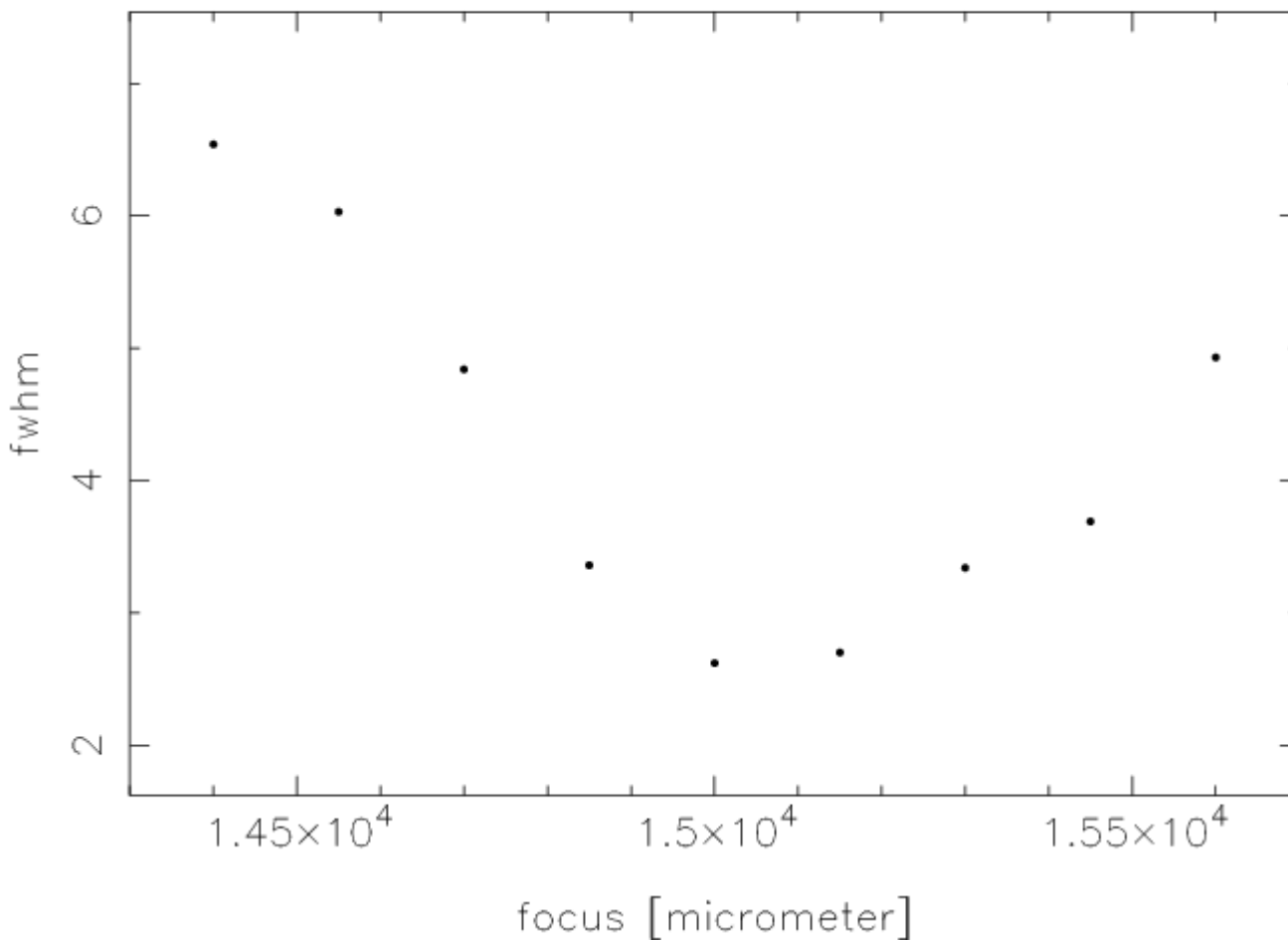
focal plane with single CCD

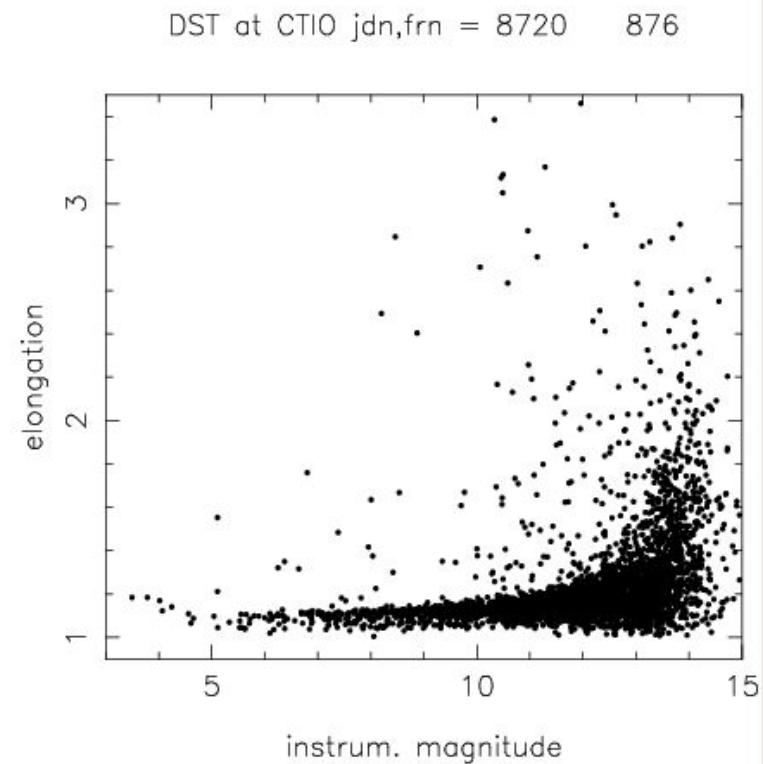
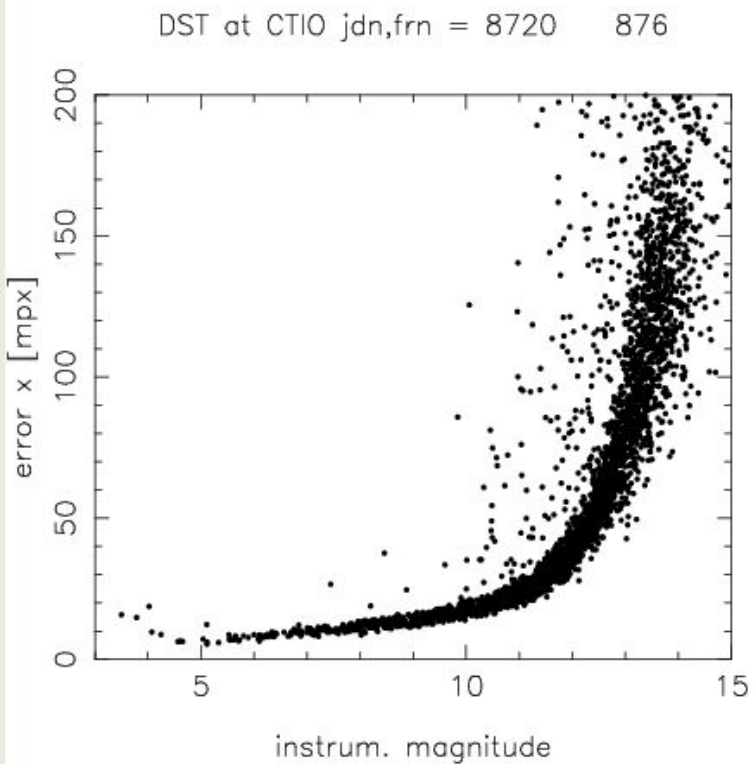
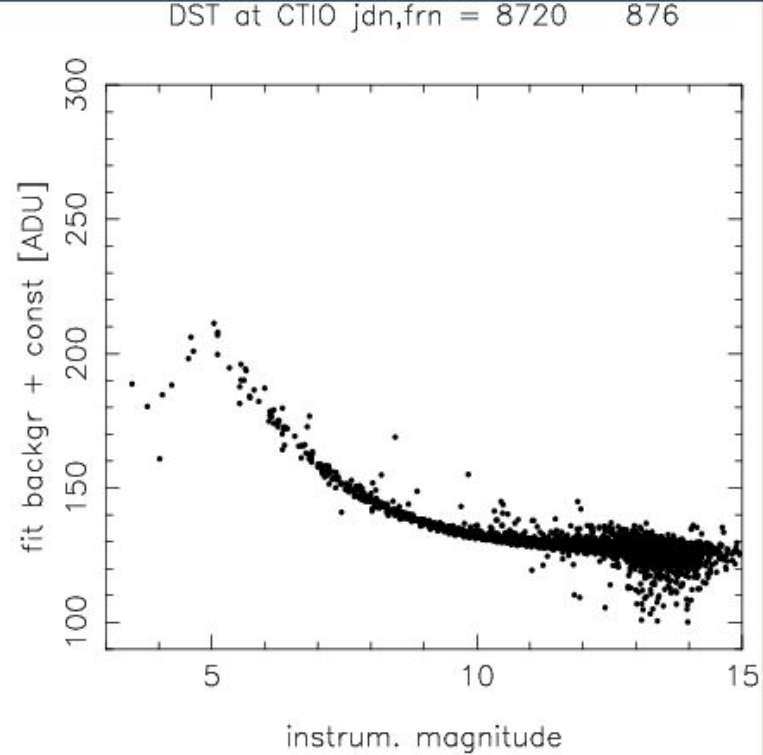
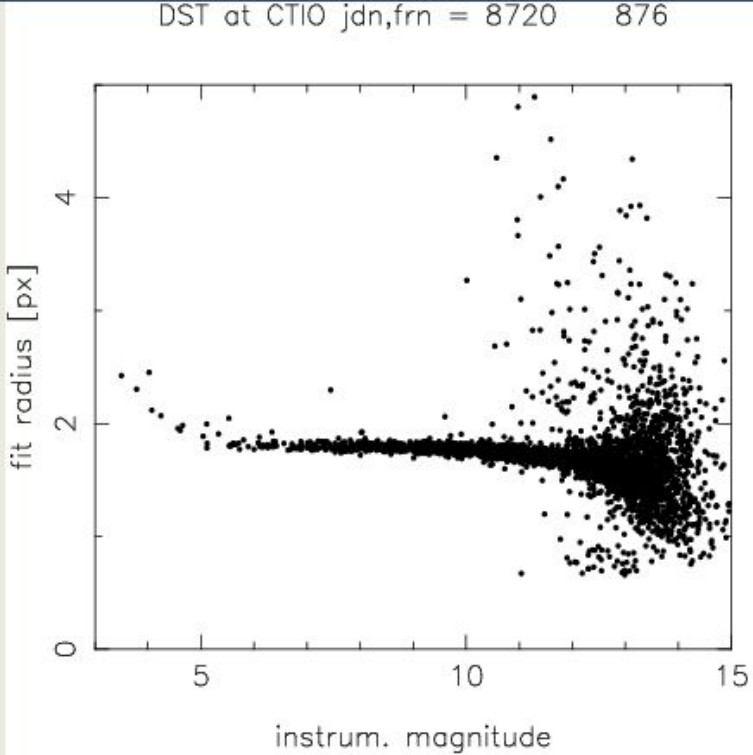




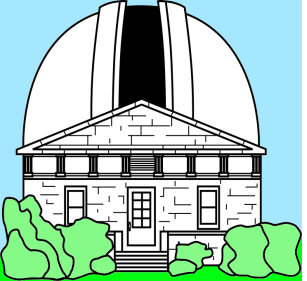
focus curve

DST focus plot:/home/urat/DST2/focp/foc10.dat

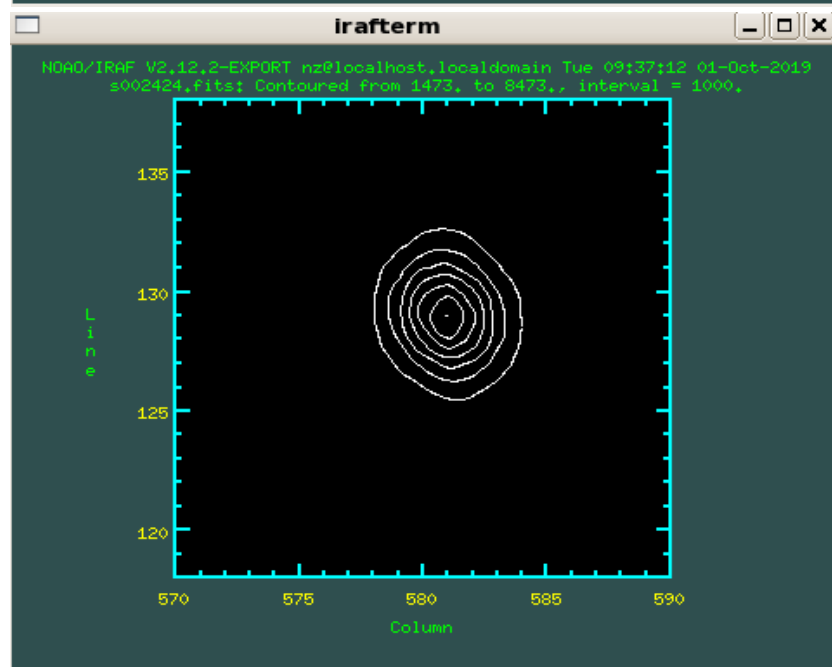
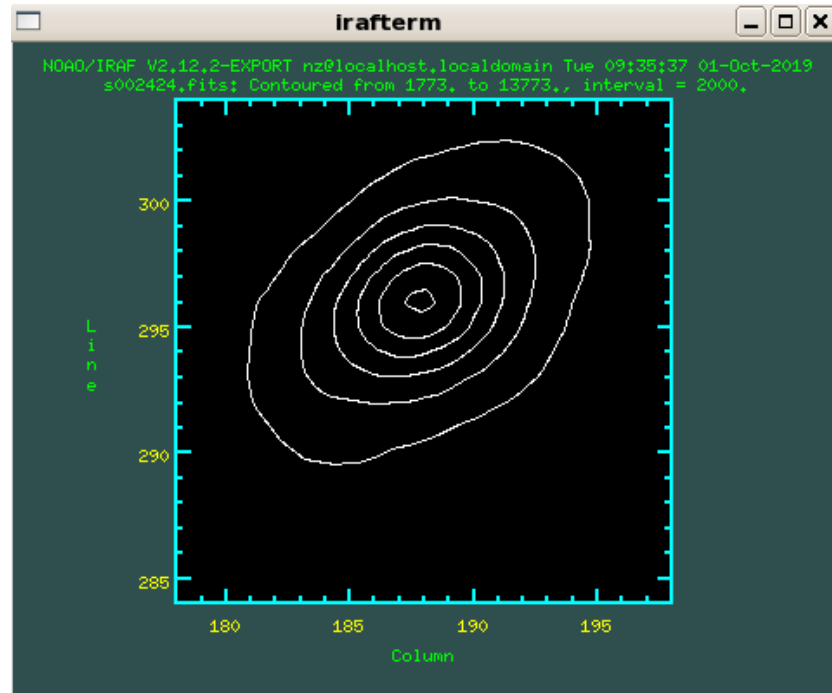
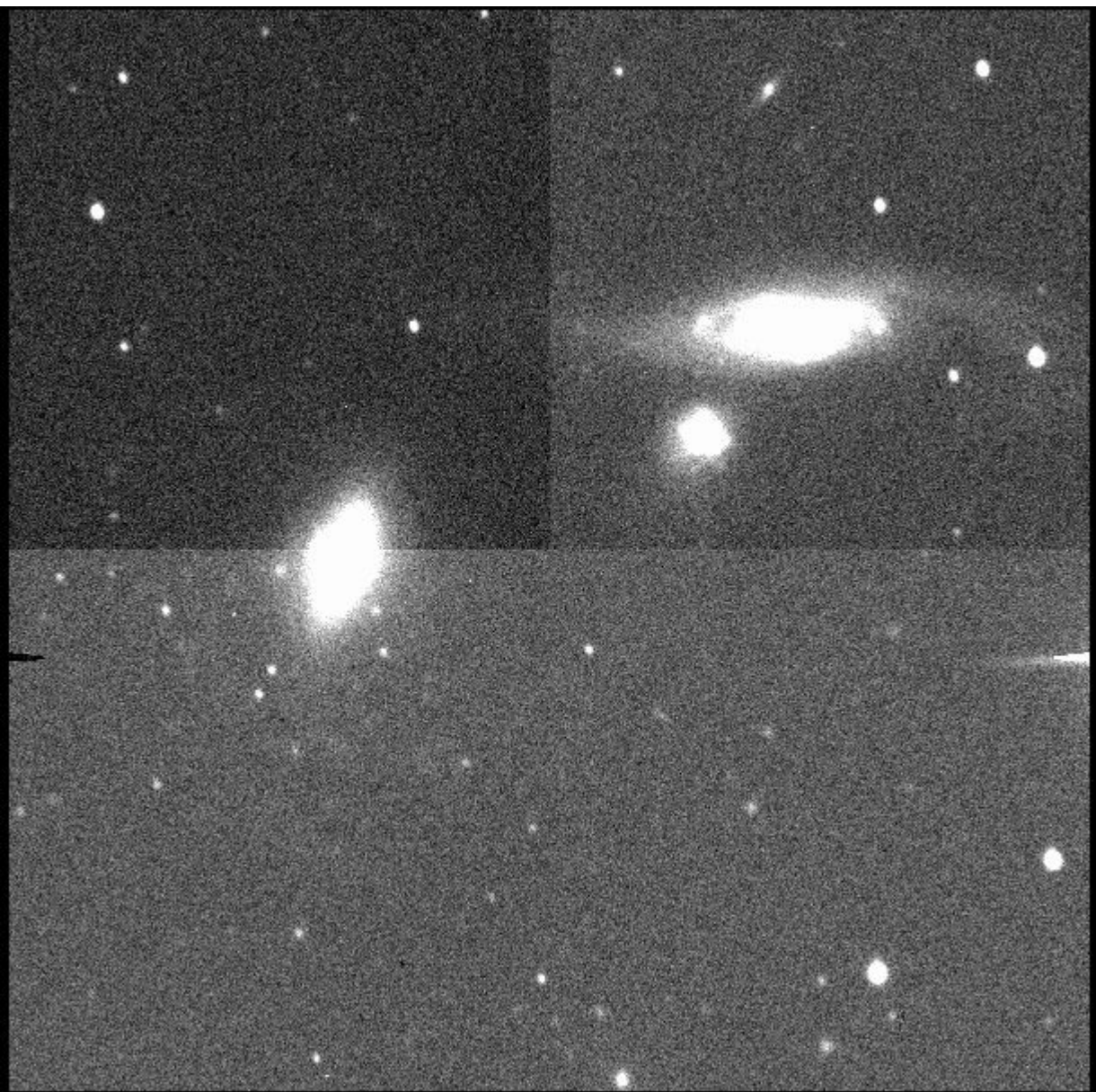


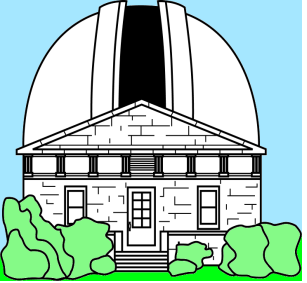


**stellar
image
profile fit
results**



NGC 7465





observing statistics

observing from Aug 24 - Sept 30

total number of nights 38

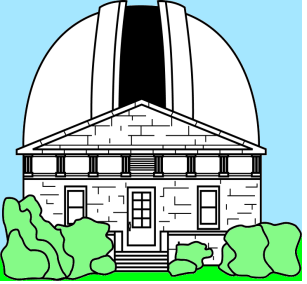
lost due to weather 4

lost due to instrument 4

nights with observations 30

total number of exposures 2797

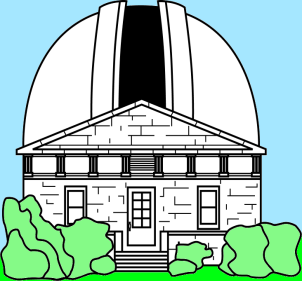
total number of sources ~ 100



issues

- reading of mount (mirror) temperatures
 - just becomes available now
 - significant improvement in **focus prediction** expected
- **image elongation**
 - about 30% of exposures show unacceptable image quality
 - more frequent for longer than shorter integrations, but do see perfect images for some exposures up to 10 min
- “hang-ups”
 - occasional loss of communication to some device
 - typical process for a new telescope to improve robustness
- reduction pipeline: in progress





Conclusions

- **DST** is operational at Cerro Tololo
 - current program: 199 FRAMEx targets
 - **high cadence** (every other night) observing (astrometry & photom.)
 - about 100+ exposures per night
 - reduction pipeline in progress
 - still some **challenges** (not always round images, extended sources)
 - expected single measure precision: **4 - 10 mas** (dep.on brightness)
- if your **favorite target** is not on the list:
 - **now is the time to act !**