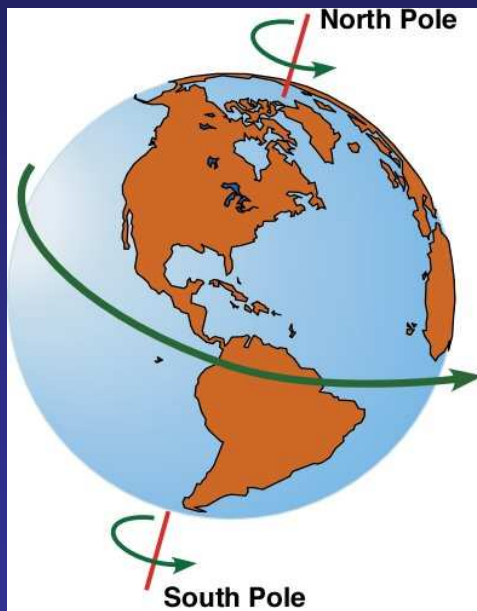


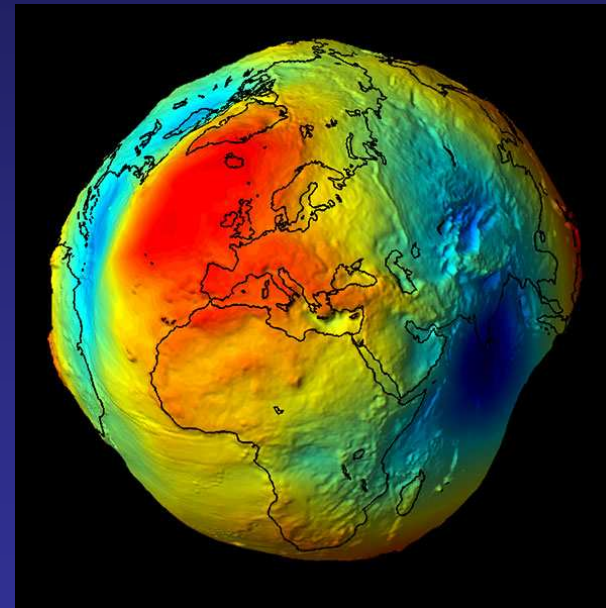
About the MacCullagh relations in Relativity

Michael.Soffel, Sergei.Klioner & Enrico Gerlach
TU Dresden

MacCullagh relations in the Newtonian framework



Earth's rotation



Earth's gravity field

Newtonian MacCullagh relations: spherical representation

$$C_{21} = -\frac{I_{13}}{Ma^2}$$

$$S_{21} = -\frac{I_{23}}{Ma^2}$$

$$C_{22} = \frac{I_{22} - I_{11}}{4Ma^2}$$

$$S_{22} = -\frac{I_{12}}{2Ma^2}$$

$$C_{20} = \frac{1}{Ma^2} \left(\frac{I_{11} + I_{22}}{2} - I_{33} \right)$$

$$U(\mathbf{x}) = \left(\frac{GM}{r} \right) \sum_{l=0}^{\infty} \sum_{m=0}^l \left(\frac{a}{r} \right)^l P_{lm}(\cos \theta) (C_{lm} \cos m\phi + S_{lm} \sin m\phi)$$

$$I_{ij} = \int_B d^3x \rho(\mathbf{x}^2 \delta_{ij} - x^i x^j)$$

Newtonian MacCullagh relations: Cartesian coordinates

$$M_{ij} = -\text{STF}(I_{ij})$$

$$U(\mathbf{x}) = G \sum_{l=0}^{\infty} \frac{(2l-1)!!}{l!} M_L \frac{\hat{n}_L}{r^{l+1}}$$

$$\hat{n}_L \equiv \frac{\hat{x}^L}{r^l}$$

$$\hat{x}^L \equiv \hat{x}^{i_1 \dots i_l} = \text{STF}_{i_1 \dots i_l} (x^{i_1} \dots x^{i_l})$$

$$\hat{x}^{ij} \equiv \text{STF}_{i,j} (x^i x^j) = x^i x^j - \frac{1}{3} \delta_{ij} \mathbf{x}^2$$

Relativity

Post-Newtonian framework (DSX)

Blanchet-Damour quadrupole mass moment

$$M_{ij} = \int_B d^3x \hat{x}_{ij} \sigma + \frac{1}{14c^2} \frac{d^2}{dt^2} \int_B d^3x \hat{x}_{ij} \mathbf{x}^2 \sigma - \frac{20}{21c^2} \frac{d}{dt} \int_B d^3x \hat{x}^{ijk} \sigma^k$$

$$\sigma \equiv \frac{T^{00} + T^{ss}}{c^2}, \quad \sigma^i \equiv \frac{T^{0i}}{c}$$

A post-Newtonian spin vector was defined in DSX III

$$S_i = \epsilon_{ijk} \int_B d^3x x^j \left[\sigma^k \left(1 + \frac{4U}{c^2} \right) - \frac{\sigma}{c^2} \left(4V^k - \frac{1}{2}Q^k \right) \right]$$

$$V^k(\mathbf{x}) \equiv G \int_B d^3x' \frac{\sigma^k(\mathbf{x}')}{|\mathbf{x} - \mathbf{x}'|}$$

$$Q^k(\mathbf{x}) \equiv G \int_B d^3x' \sigma^l(\mathbf{x}') \frac{\delta_{kl} - n_{xx'}^k n_{xx'}^l}{|\mathbf{x} - \mathbf{x}'|}$$

$$n_{xx'}^k \equiv \frac{(x - x')^k}{|\mathbf{x} - \mathbf{x}'|}$$

From this a post-Newtonian moment of inertia tensor can be derived

Question

Is the Newtonian relation

$$M_{ij} = -\text{STF}(I_{ij})$$

valid also to post-Newtonian order ?

To answer this questions we employed a specific Newtonian model for the $1/c^2$ terms:

a homogeneous oblate spheroid with b (in z -direction) \ll

the internal graviational potential U is given by

$$U(\mathbf{x}) = \pi G\rho [\mathcal{I} - \mathcal{A}(x^2 + y^2) - \mathcal{B}z^2]$$

where

$$\begin{aligned}\mathcal{A} &= \frac{(1 - e^2)^{1/2}}{e^3} \arcsin(e) - \frac{1 - e^2}{e^2} = \frac{2}{3} - \frac{2}{15}e^2 - \frac{8}{105}e^4 + \dots \\ \mathcal{B} &= \frac{2}{e^2} - \frac{2(1 - e^2)^{1/2}}{e^3} \arcsin(e) = \frac{2}{3} + \frac{4}{15}e^2 + \frac{16}{105}e^4 + \dots\end{aligned}$$

$(2\mathcal{A} + \mathcal{B} = 2)$ and

$$\mathcal{I} = 2\mathcal{A}a^2 + \mathcal{B}b^2.$$

The geometrical eccentricity e of the spheroid is given by

$$e^2 = 1 - \frac{b^2}{a^2}.$$

RESULT

$$M_{ij} \neq -\text{STF}(I_{ij})$$

For our model:

$$M_{ij} = -\text{STF}(I_{ij}) - \left(\frac{148}{875}e^2\right) \left(\frac{GM}{c^2a}\right) (Ma^2) \hat{e}_{ij} + \mathcal{O}(e^4)$$

$$\hat{e}_{ij} \equiv \text{diag}(1, 1, -2)$$

PN MacCullagh relations, spherical representation

For our specific model only one MC relation needs to be modified

$$C_{20}^* = \frac{1}{Ma^2} \left(\frac{I_{11} + I_{22}}{2} - I_{33} \right)$$

$$C_{20}^* = C_{20} - \frac{444}{875} \left(\frac{GM_{\odot}}{c^2 a} \right) e^2$$



δ - term

Numbers for the Earth:

$$e^2 = 0.0066944$$

Taking the EGM96 value for GM_{\oplus} ($3.986004418 \cdot 10^{14} \text{ m}^3\text{s}^{-2}$) the relativistic scale factor reads

$$\left(\frac{GM_{\oplus}}{c^2 a}\right) = 6.9534851 \cdot 10^{-10}$$

so that

$$\delta \equiv \frac{444}{875} \left(\frac{GM_{\oplus}}{c^2 a}\right) e^2 = 2.362 \times 10^{-12}$$

Implications for a post-Newtonian SMART-97 like nutation model

In SMART97 GEM-T3 values for the potential coefficients are used:

$$\begin{aligned} C_{20} &= -1082.626\,074\,59 \times 10^{-6} \\ C_{22} &= 1.574\,410\,20 \times 10^{-6} \\ S_{22} &= -0.903\,757\,18 \times 10^{-6}. \end{aligned}$$

The principle moments of inertia, A , B and C are obtained in the following way: one starts with a certain initial value for C :

$$C = 0.180\,548\,385\,370 \times 10^{-14} M_{\odot} \text{AU}^2,$$

where M_{\odot} is the mass of the Sun,

$$\frac{M_{\odot}}{M_{\oplus}} = 332\,946.045,$$

and the astronomical unit, AU, was taken as

$$\text{AU} = 149\,597\,870\,610 \text{ m}.$$

$B - A$ is then obtained from

$$\frac{B - A}{M_{\oplus} a^2} = 4 \sqrt{C_{22}^2 + S_{22}^2} = 7.261\,454 \times 10^{-6}$$

that would still be valid relativistically (note, that the potential coefficients refer to axes that differ in the equatorial plane by an angle $\alpha = 14^{\circ}.95$ (West) with respect to the principle axes of inertia). Then the Newtonian MacCullagh relation for the dynamical ellipticity H_d is used for SMART97:

$$H_d \equiv \frac{2C - A - B}{2C} = -M_{\oplus} a^2 \frac{C_{2,0}}{C}$$

that leads to

$$\begin{aligned} A &= 0.179955329763 \times 10^{-14} M_{\odot} \text{AU}^2 \\ B &= 0.179959294245 \times 10^{-14} M_{\odot} \text{AU}^2. \end{aligned}$$

If the post-Newtonian MacCullagh relation (30) is used one would get

$$\begin{aligned} A &= 0.179955329761 \times 10^{-14} M_{\odot} \text{AU}^2 \\ B &= 0.179959294243 \times 10^{-14} M_{\odot} \text{AU}^2. \end{aligned}$$

Implications for a post-Newtonian SMART-97 like nutation model

In SMART97 GEM-T3 values for the potential coefficients are used:

$$\begin{aligned} C_{20} &= -1082.626\,074\,59 \times 10^{-6} \\ C_{22} &= 1.574\,410\,20 \times 10^{-6} \\ S_{22} &= -0.903\,757\,18 \times 10^{-6}. \end{aligned}$$

The principle moments of inertia, A , B and C are obtained in the following way: one starts with a certain initial value for C :

$$C = 0.180\,548\,385\,370 \times 10^{-14} M_{\odot} \text{AU}^2,$$

where M_{\odot} is the mass of the Sun,

$$\frac{M_{\odot}}{M_{\oplus}} = 332\,946.045,$$

and the astronomical unit, AU, was taken as

$$\text{AU} = 149\,597\,870\,610 \text{ m}.$$

$B - A$ is then obtained from

$$\frac{B - A}{M_{\oplus} a^2} = 4 \sqrt{C_{22}^2 + S_{22}^2} = 7.261\,454 \times 10^{-6}$$

that would still be valid relativistically (note, that that the potential coefficients refer to axes that differ in the equatorial plane by an angle $\alpha = 14^{\circ}.95$ (West) with respect to the principle axes of inertial). Then the Newtonian MacCullagh relation for the dynamical ellipticity H_d is used for SMART97:

$$H_d \equiv \frac{2C - A - B}{2C} = -M_{\oplus} a^2 \frac{C_{2,0}}{C}$$

that leads to

$$\begin{aligned} A &= 0.179955329763 \times 10^{-14} M_{\odot} \text{AU}^2 \\ B &= 0.179959294245 \times 10^{-14} M_{\odot} \text{AU}^2. \end{aligned}$$

If the post-Newtonian MacCullagh relation (30) is used one would get

$$\begin{aligned} A &= 0.179955329761 \times 10^{-14} M_{\odot} \text{AU}^2 \\ B &= 0.179959294243 \times 10^{-14} M_{\odot} \text{AU}^2. \end{aligned}$$

12th digit
after the comma

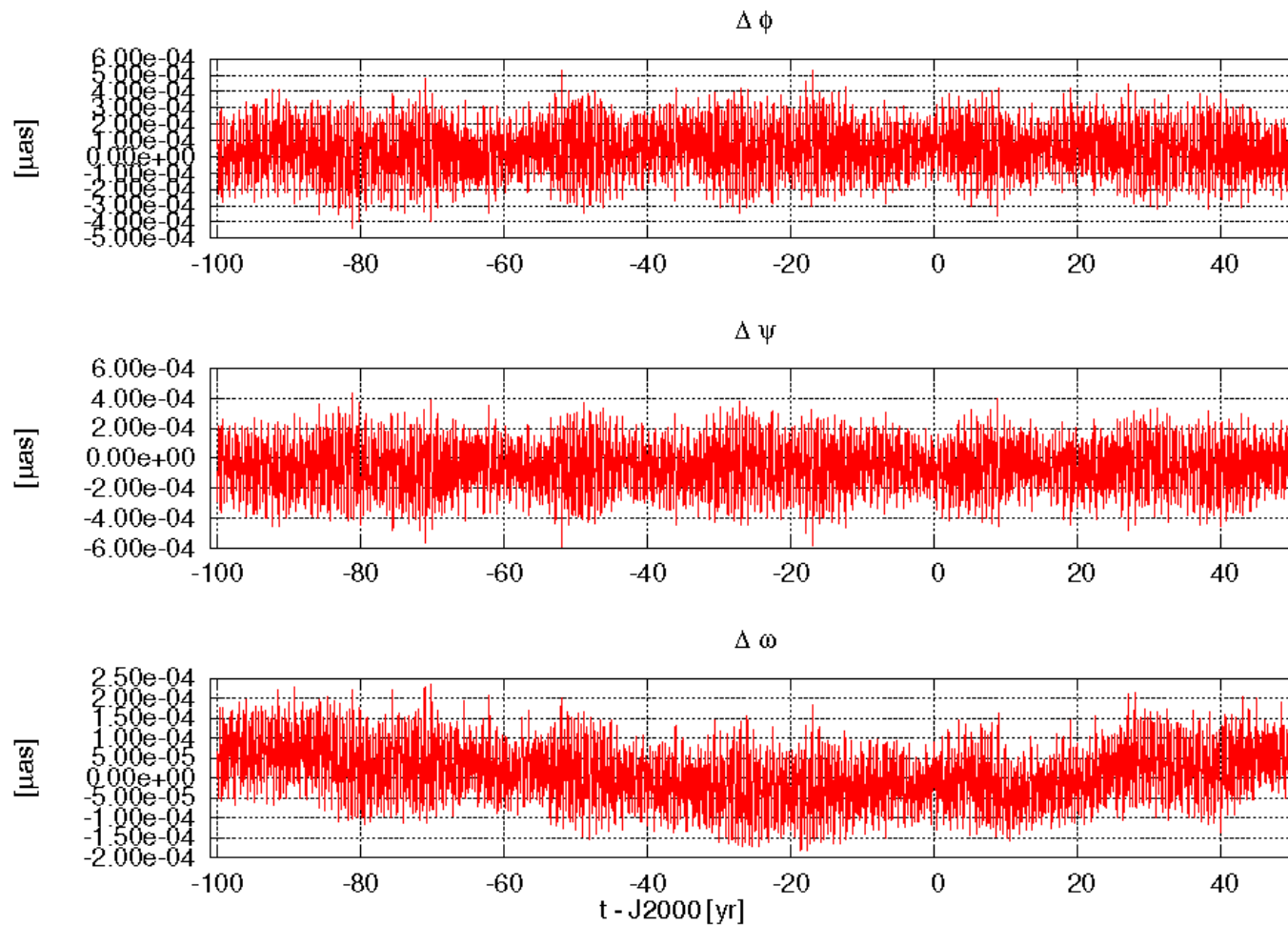


Figure 1: Differences (without $\delta/$ with δ) for the three Euler angles $\Delta\phi$, $\Delta\psi$ and $\Delta\omega$ are shown as functions of time. The δ was magnified by a factor of 1000, so that effects for the real Earth will be about one thousand times smaller.

CONCLUSION

A violation of the classical Mac Cullagh relations in Relativity can safely be neglected for a post-Newtonian Prec/Nut theory at $0.1 \mu\text{s}$ accuracies